

Observations of AT2019wey

The Mysterious Galactic Low-mass X-ray Binary

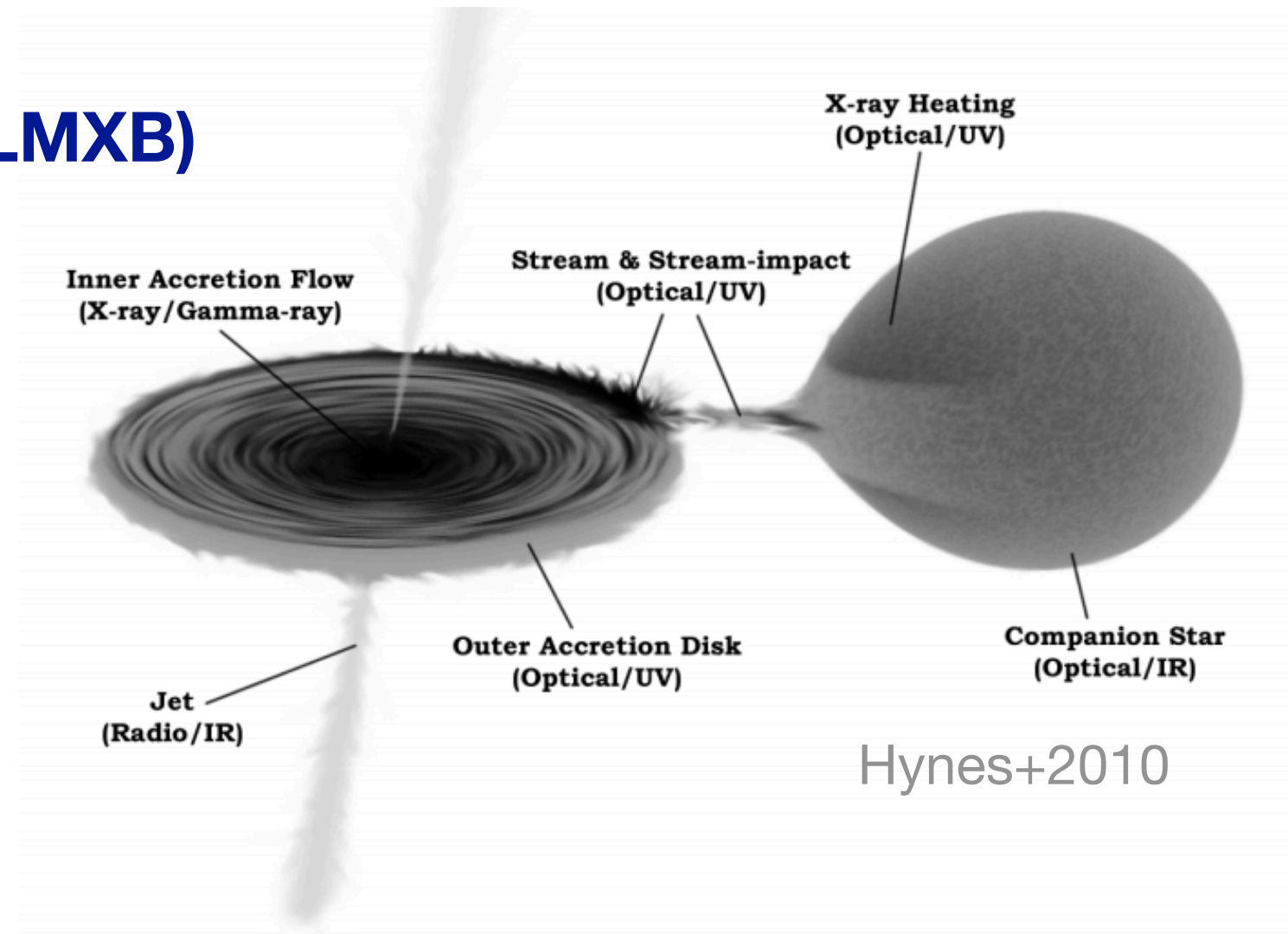
Yuhan Yao

Caltech

Oct 20, 2020

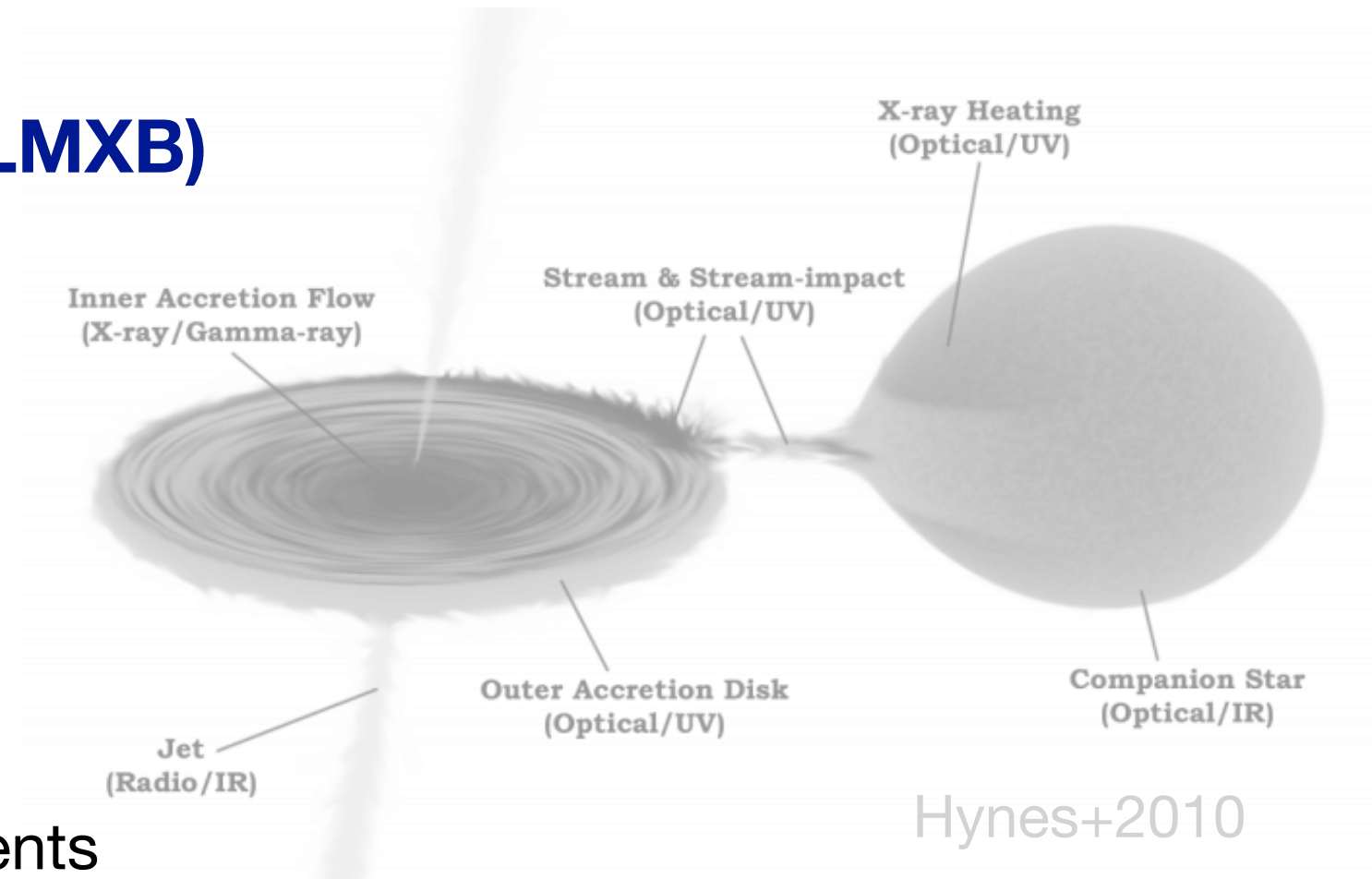
Low-mass X-ray Binaries (LMXB)

- BH / NS accretor
- Low mass companion
- Accretion disk

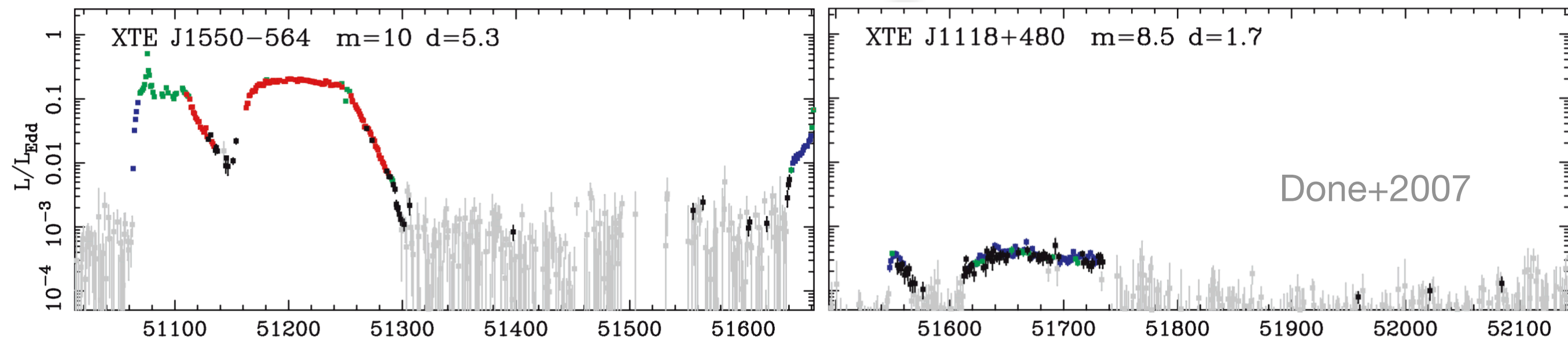


Low-mass X-ray Binaries (LMXB)

- BH / NS accretor
- Low mass companion
- Accretion disk

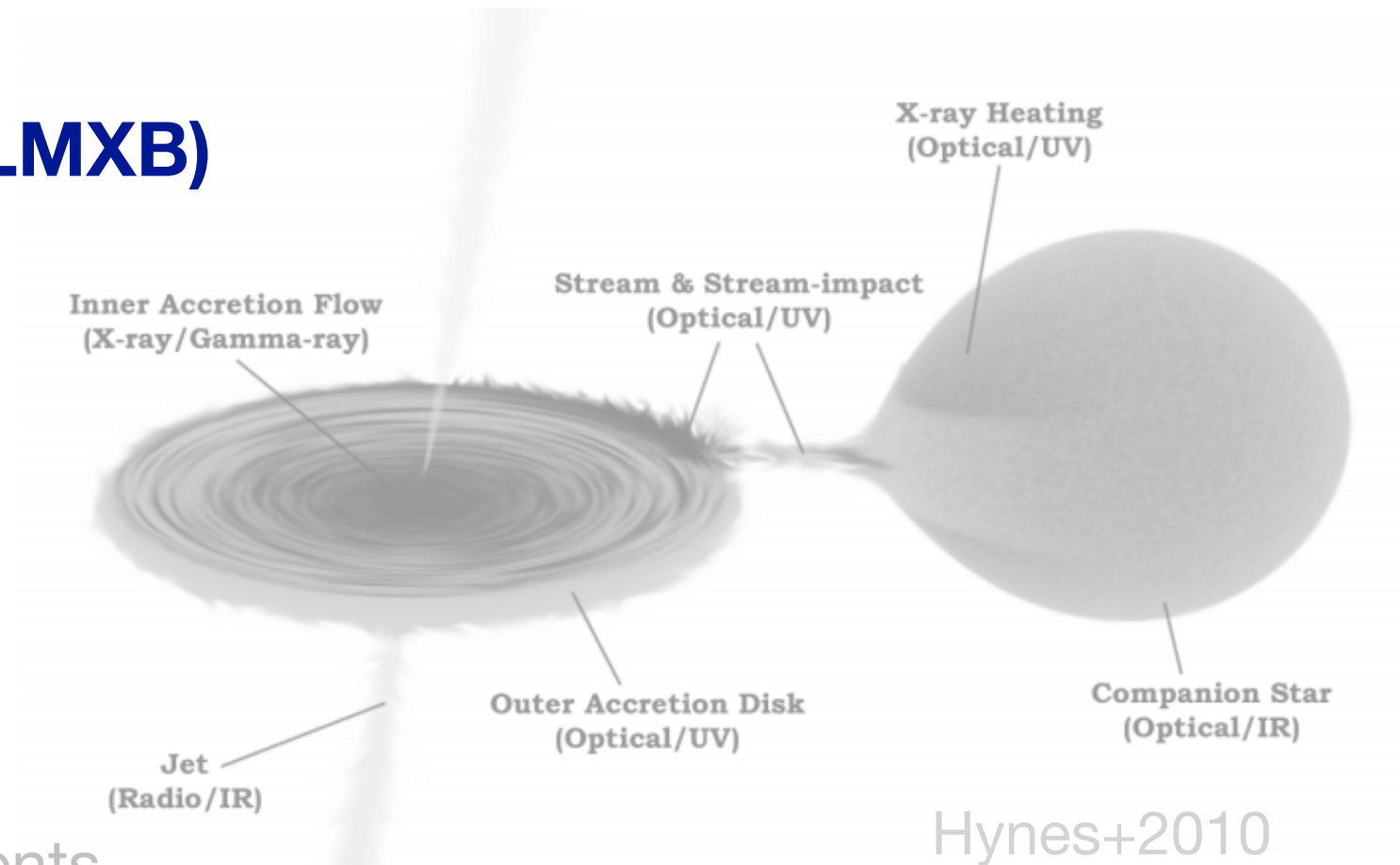


- All known LMXB BHB are transients

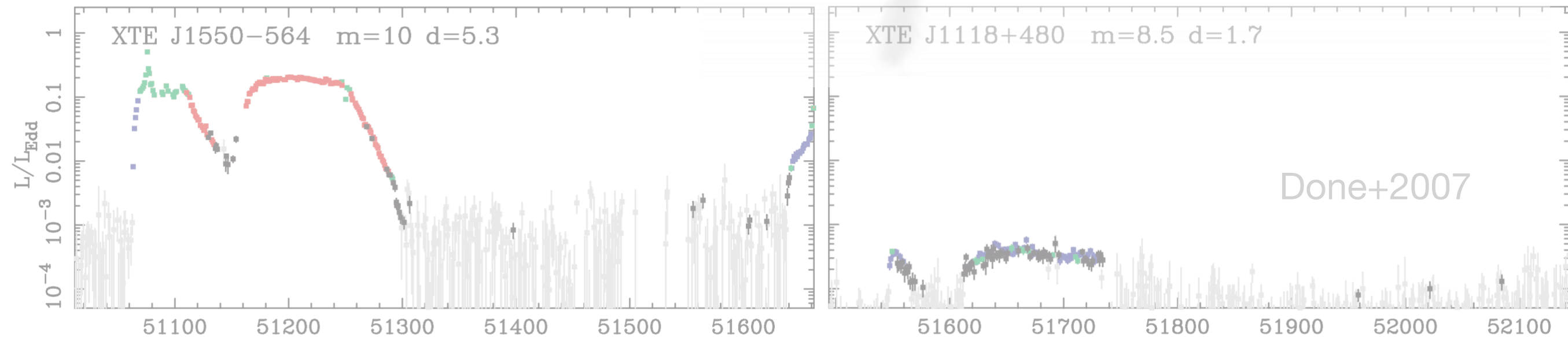


Low-mass X-ray Binaries (LMXB)

- BH / NS accretor
- Low mass companion
- Accretion disk



- All known LMXB BHB are transients

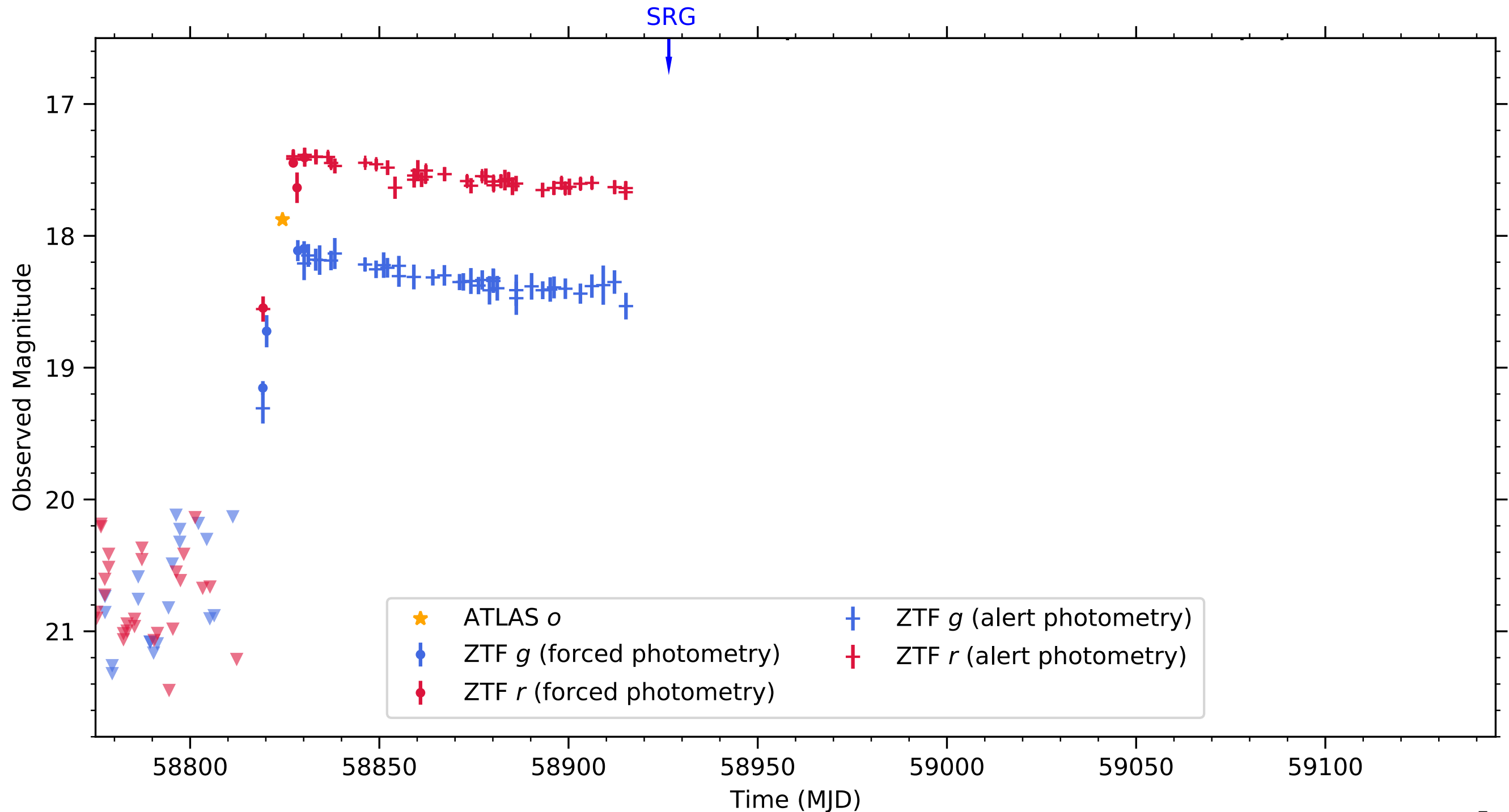


Find faint LMXB outbursts { Using more sensitive X-ray surveys → **SRG** satellite
Searching for outbursts in optical/IR → **ZTF, ATLAS, Gattini ...**

Discovery of an Optical and X-ray Transient

- ZTF 1st detection: Dec 2, 2019
- SRG discovery: Mar 18, 2020. ~1 mCrab.

AT2019wey / ZTF19acwrvzk

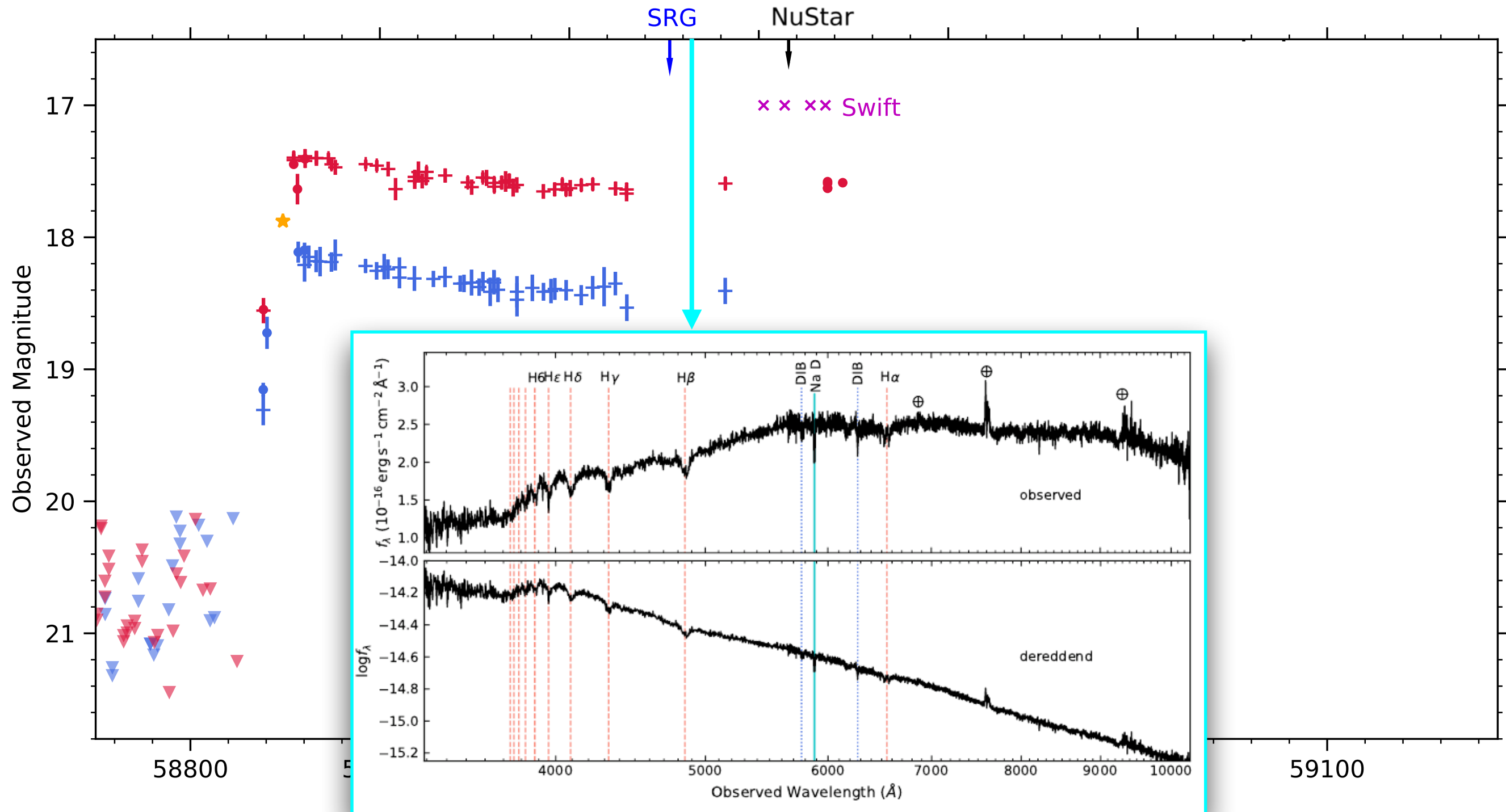


Discovery of an Optical and X-ray Transient

Galactic Origin $E(B-V) \sim 0.8-1.3$ mag

AT2019wey / ZTF19acwrvzk

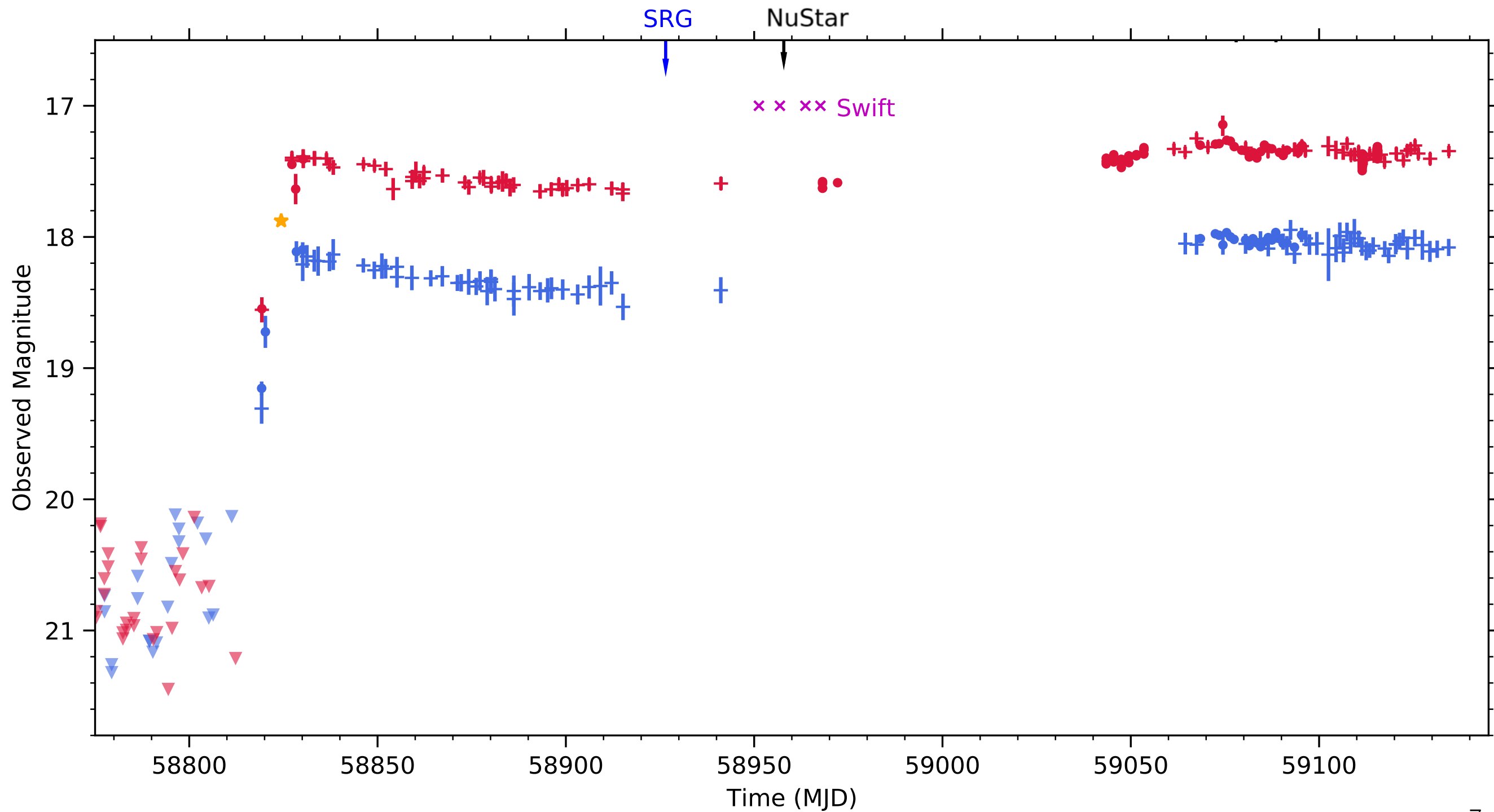
- Blue continuum + broad Balmer lines
- *Swift*/*NuSTAR* spectrum in Apr: power-law $\Gamma \sim 1.8$



Discovery of an Optical
and X-ray Transient

Galactic Origin
 $E(B-V) \sim 0.8\text{--}1.3$ mag

AT2019wey / ZTF19acwrvzk

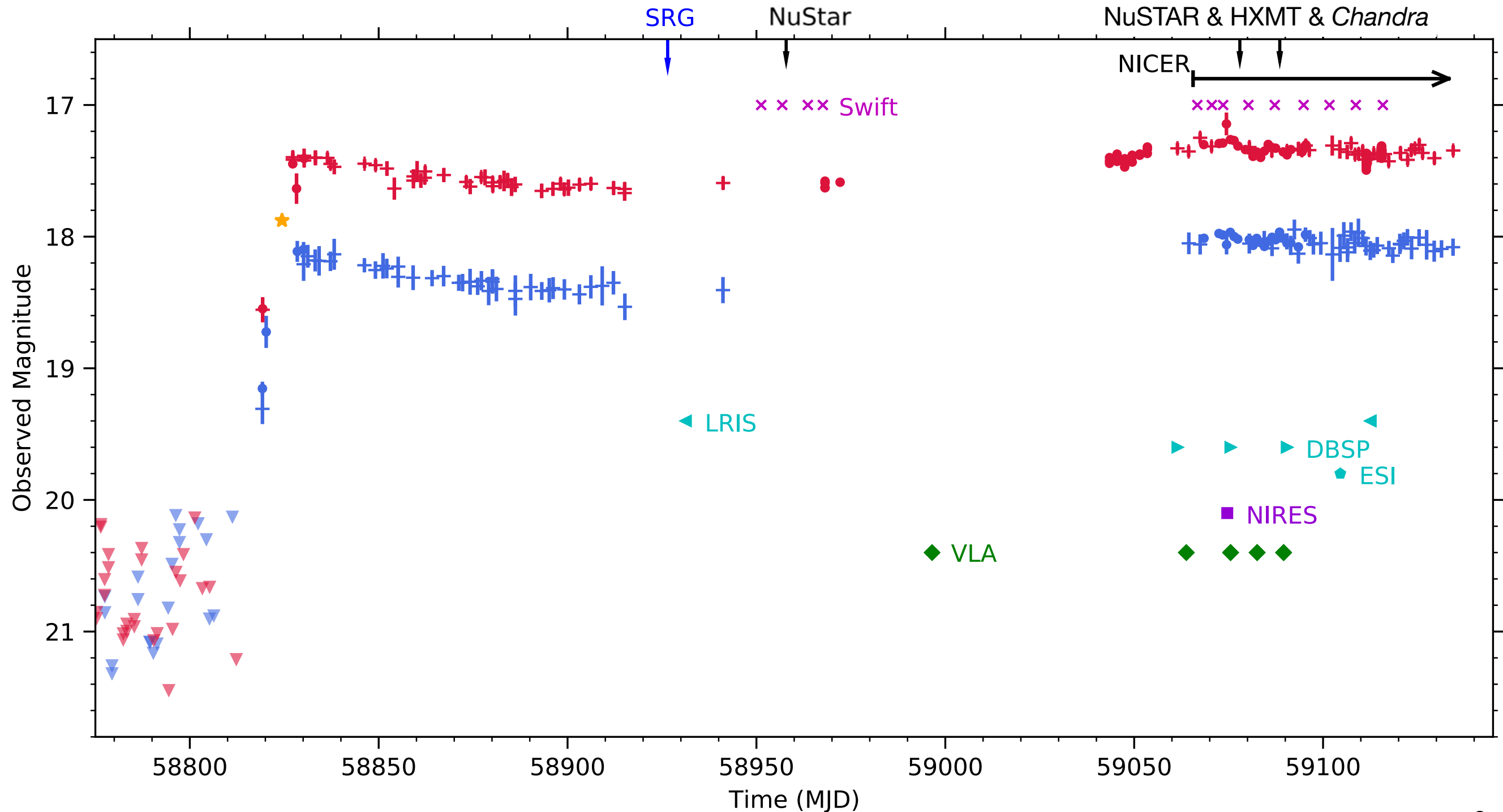


Discovery of an Optical
and X-ray Transient

Galactic Origin
 $E(B-V) \sim 0.8\text{--}1.3$ mag

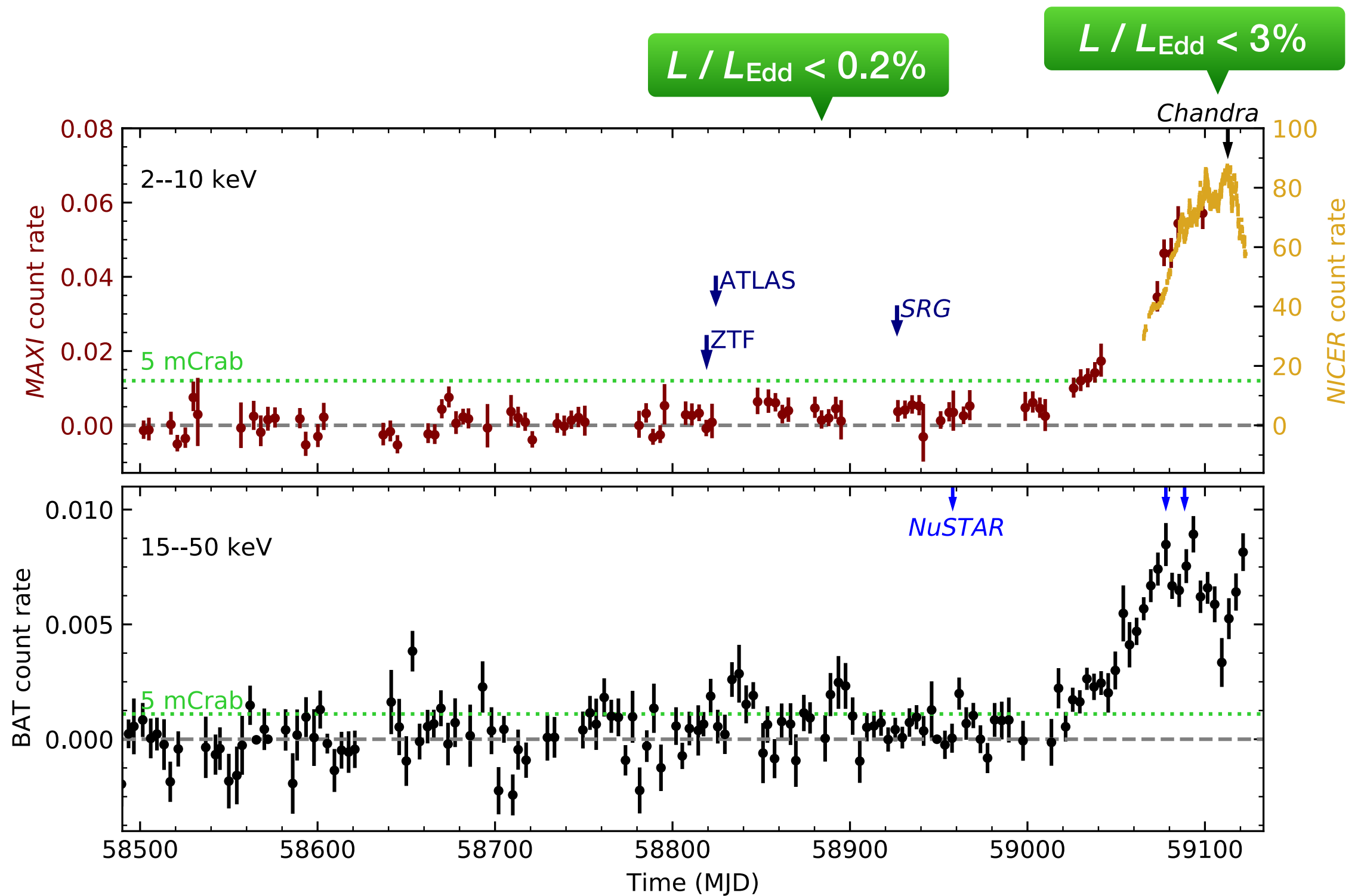
**Comprehensive Multi-
wavelength Follow-up**

AT2019wey / ZTF19acwrvzk



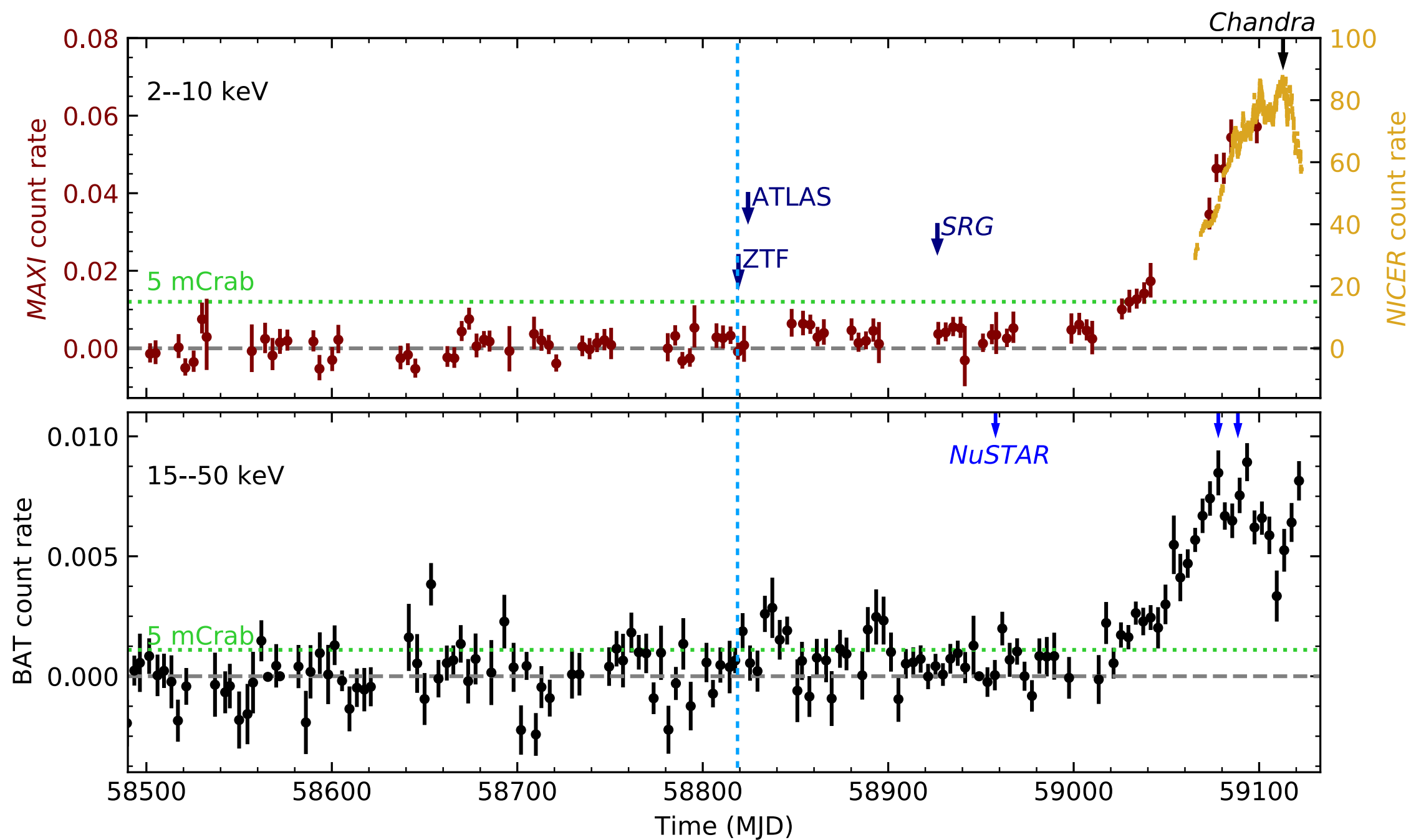
X-ray Properties

- Brightened from 1mCrab to ~25mCrab



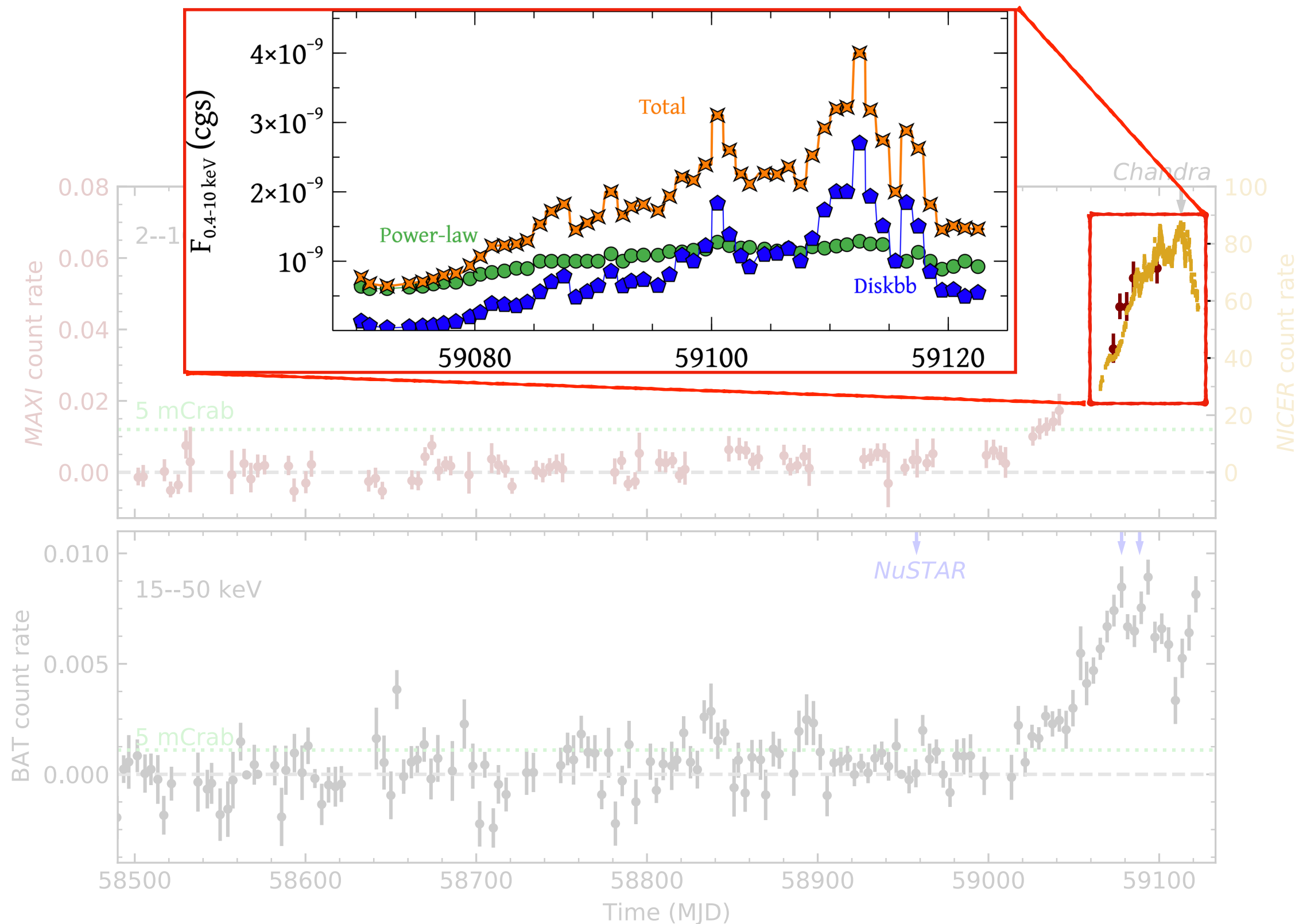
X-ray Properties

- Brightened from 1mCrab to ~25mCrab
- No X-ray nova outburst



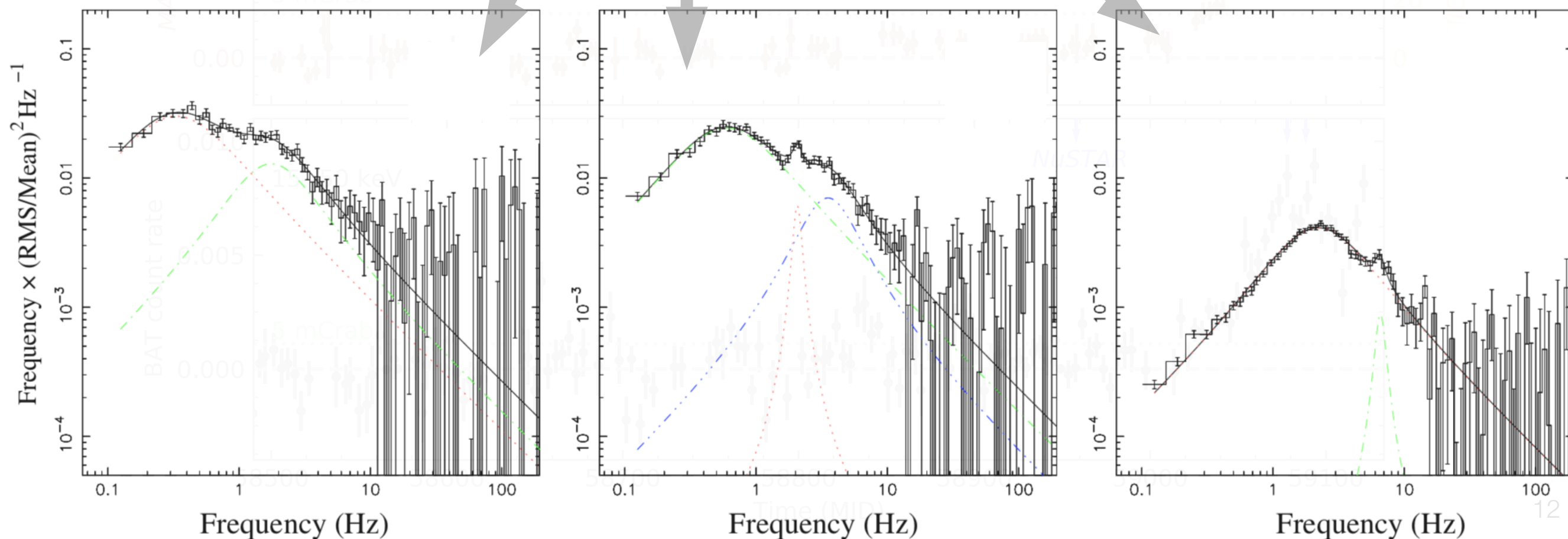
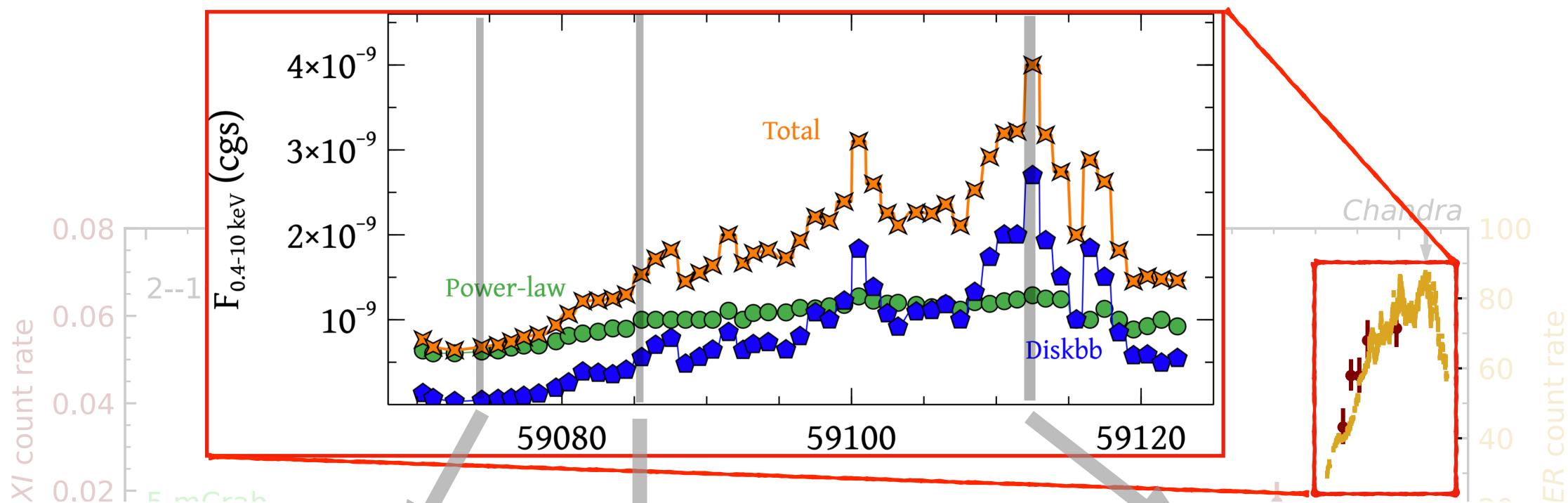
X-ray Properties

- Brightened from 1mCrab to ~25mCrab
- No X-ray nova outburst
- Soft X-ray enhancements from thermal emission



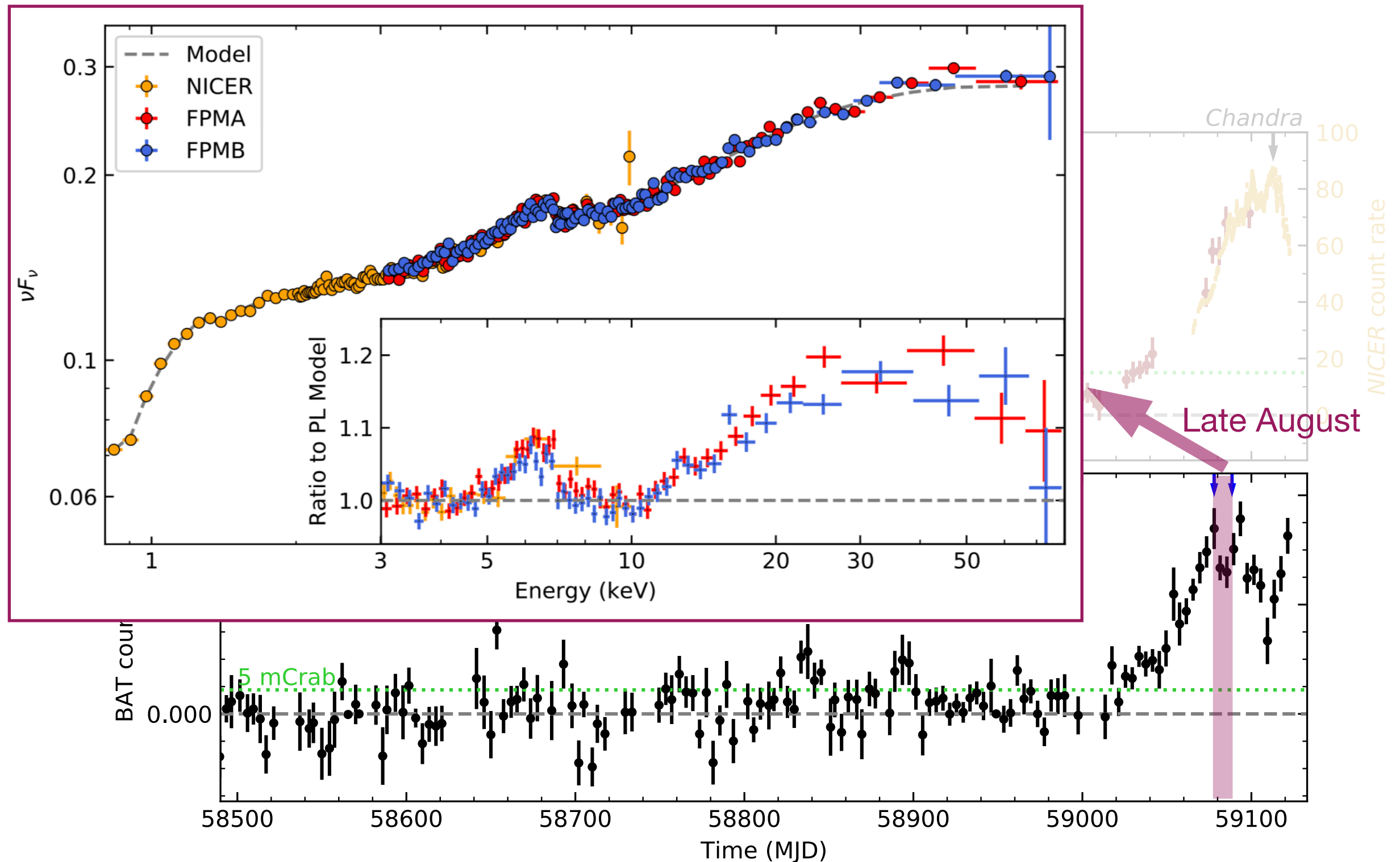
X-ray Properties

- Brightened from 1mCrab to ~25mCrab
- No X-ray nova outburst
- Soft X-ray enhancements from thermal emission
- Low-frequency QPO (higher ν as the source gets softer)



X-ray Properties

- Brightened from 1mCrab to ~25mCrab
- No X-ray nova outburst
- Soft X-ray enhancements from thermal emission
- Low-frequency QPO (higher ν as the source gets softer)
- Relativistic reflection spectrum ($i < 30^\circ$) \rightarrow NS or BH binary

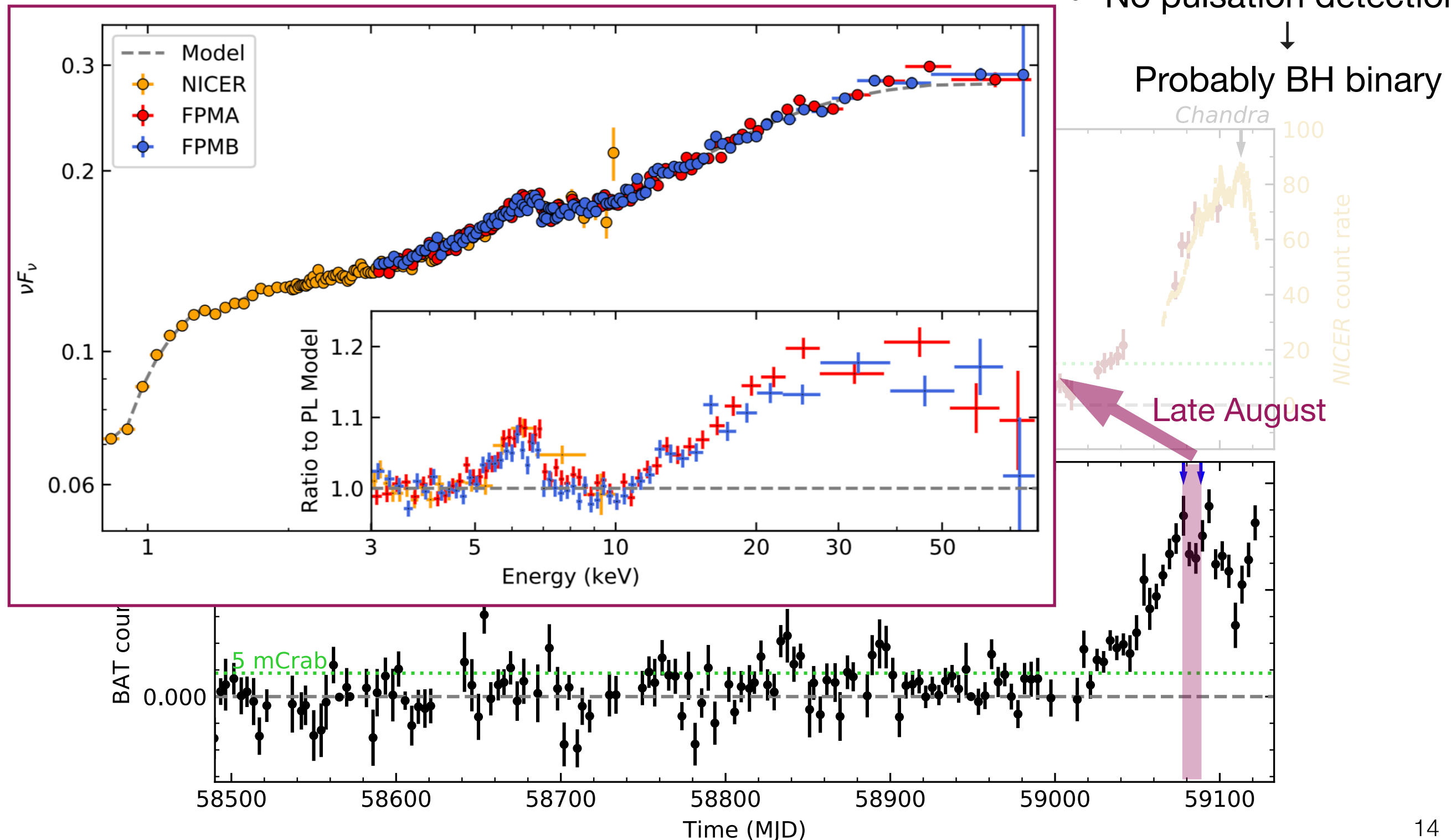


X-ray Properties

- Brightened from 1mCrab to ~25mCrab
- No X-ray nova outburst
- Soft X-ray enhancements from thermal emission
- Low-frequency QPO (higher ν as the source gets softer)

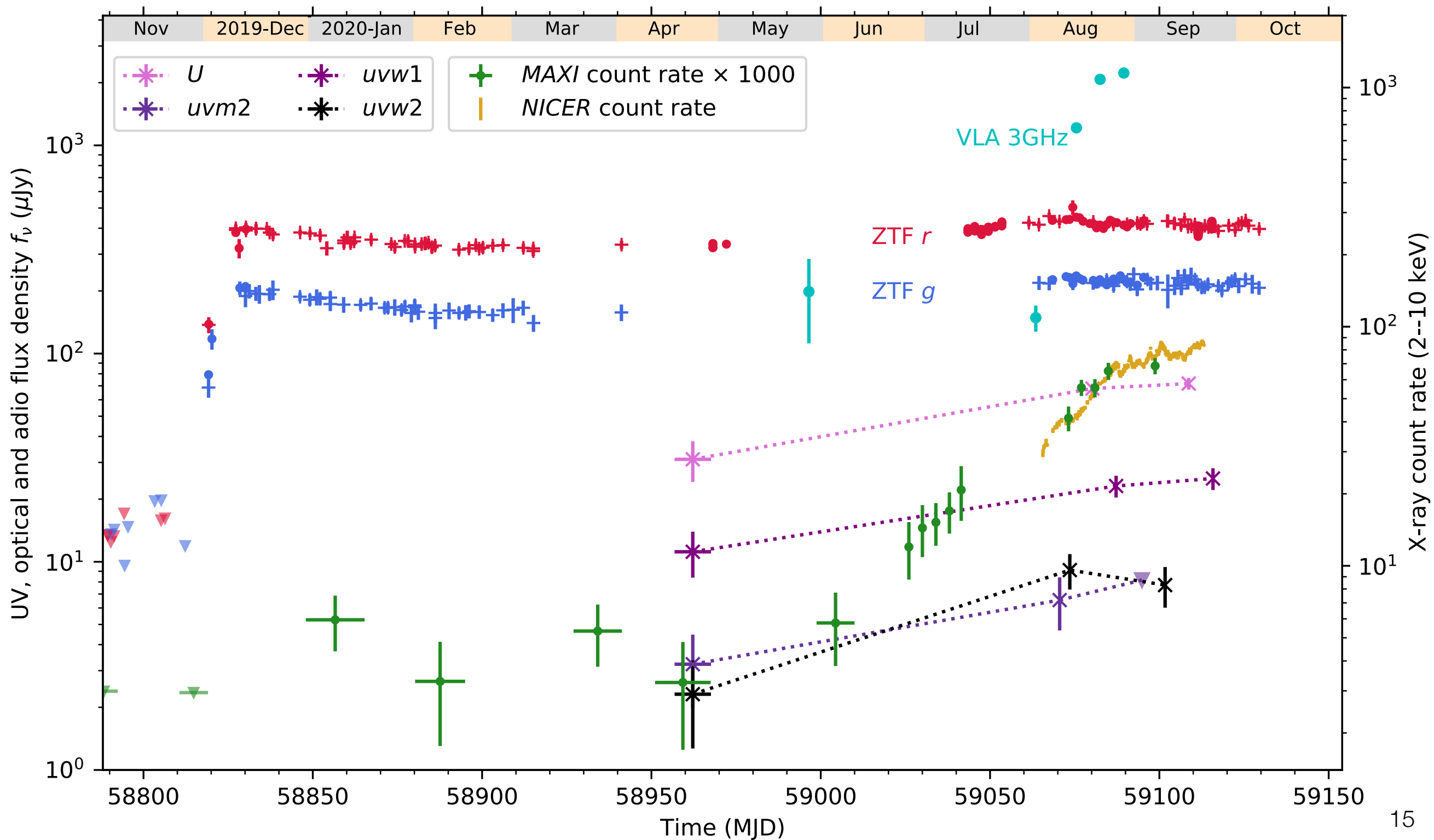
- Relativistic reflection spectrum ($i < 30^\circ$) \rightarrow NS or BH binary

- No cutoff < 300 keV
- No pulsation detection



The ***Mysterious Part:***
Nearly Constant in Optical Despite Significant Brightening in Radio/ X-ray!!!

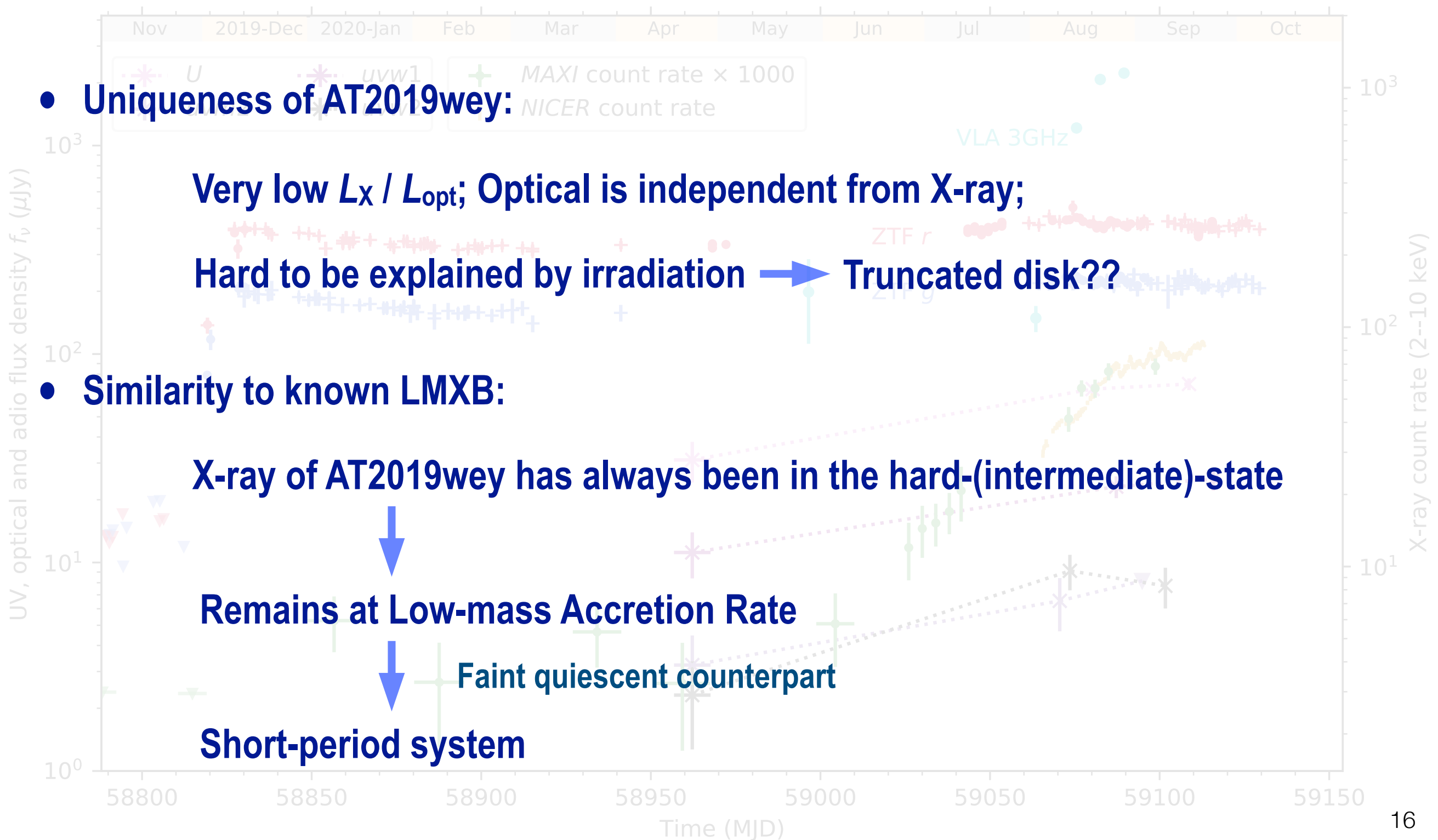
Brightening: X-ray $\uparrow > 20$, radio $\uparrow > 10$, UV $\uparrow \sim 3$, Opt $\uparrow < 1.3$



The *Mysterious* Part:

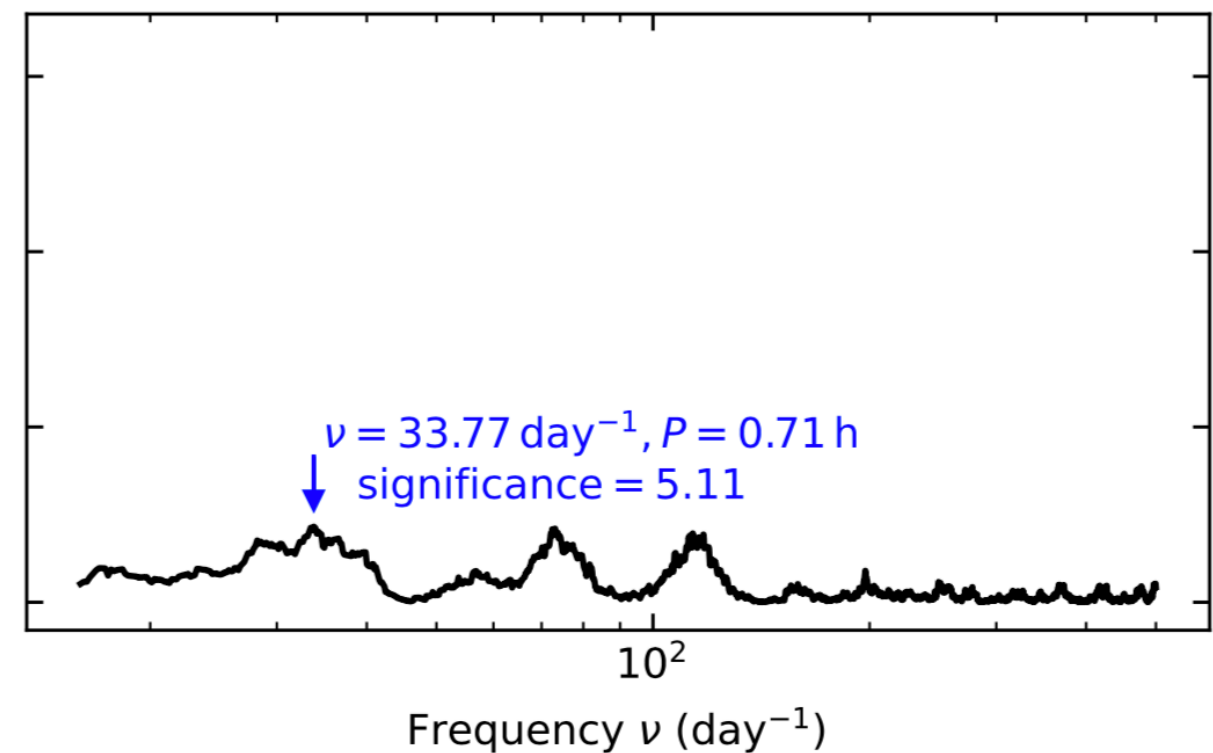
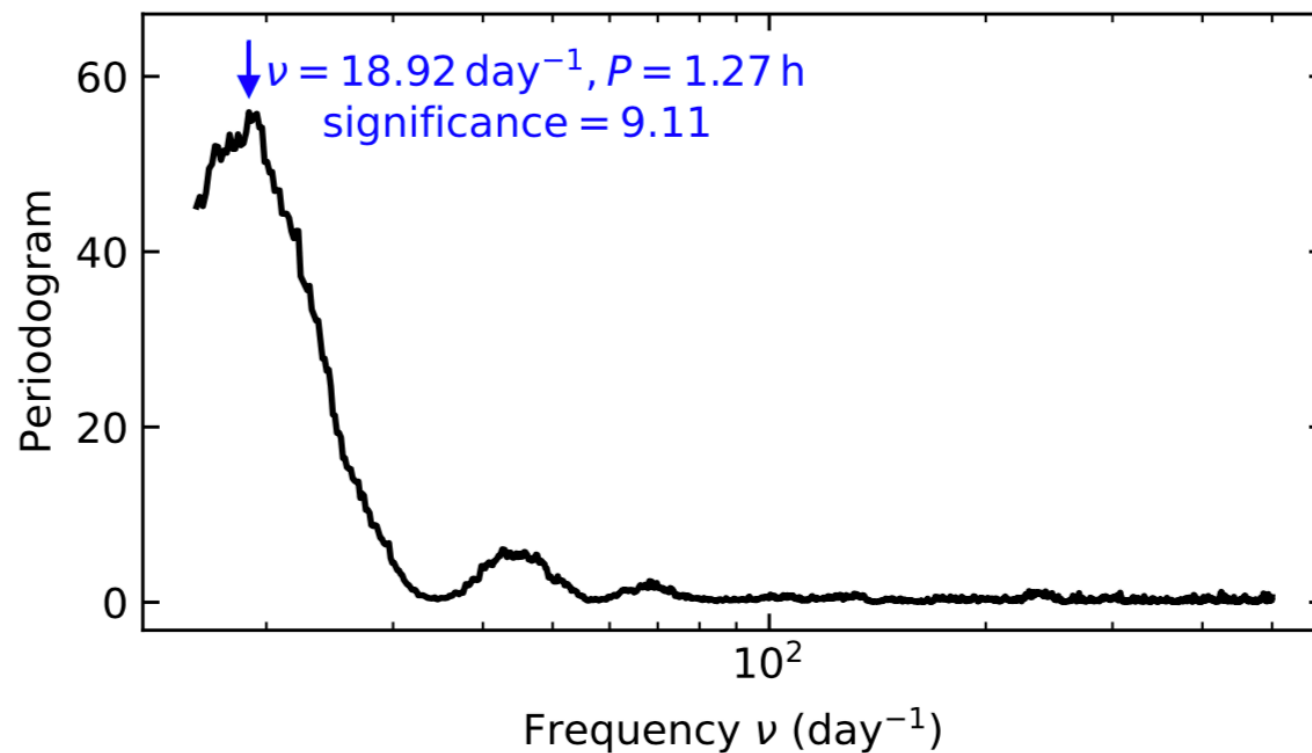
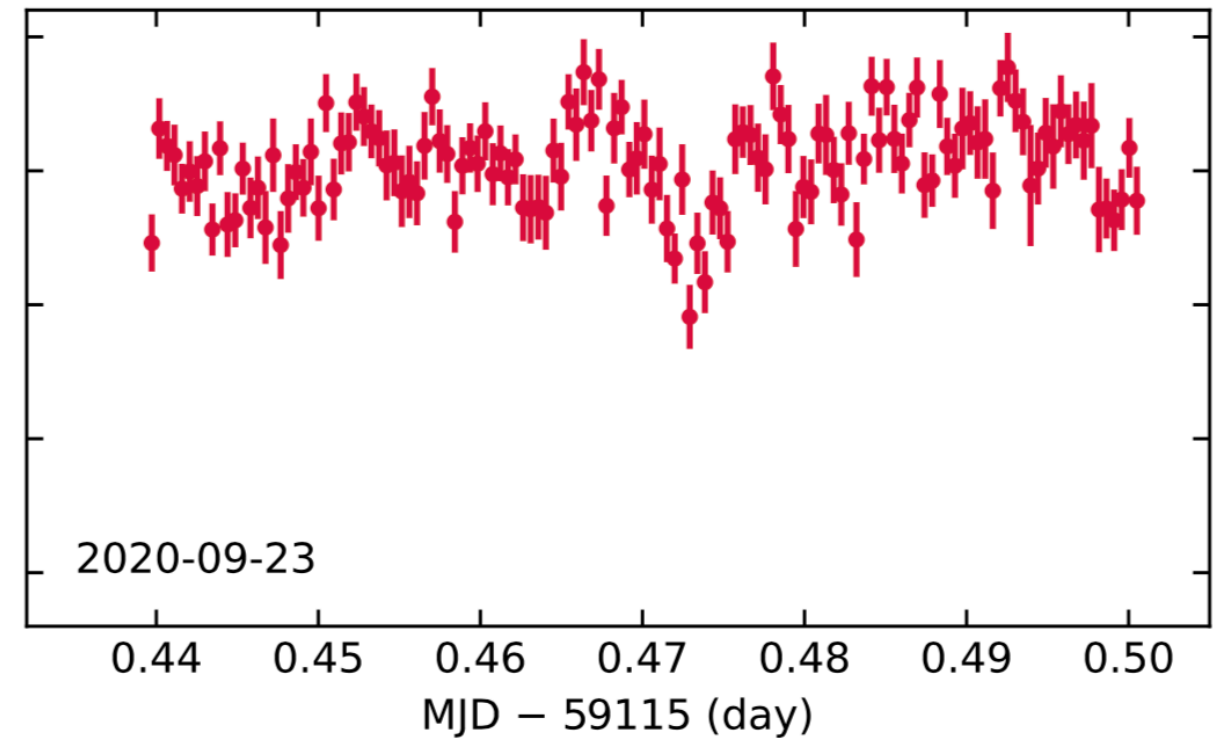
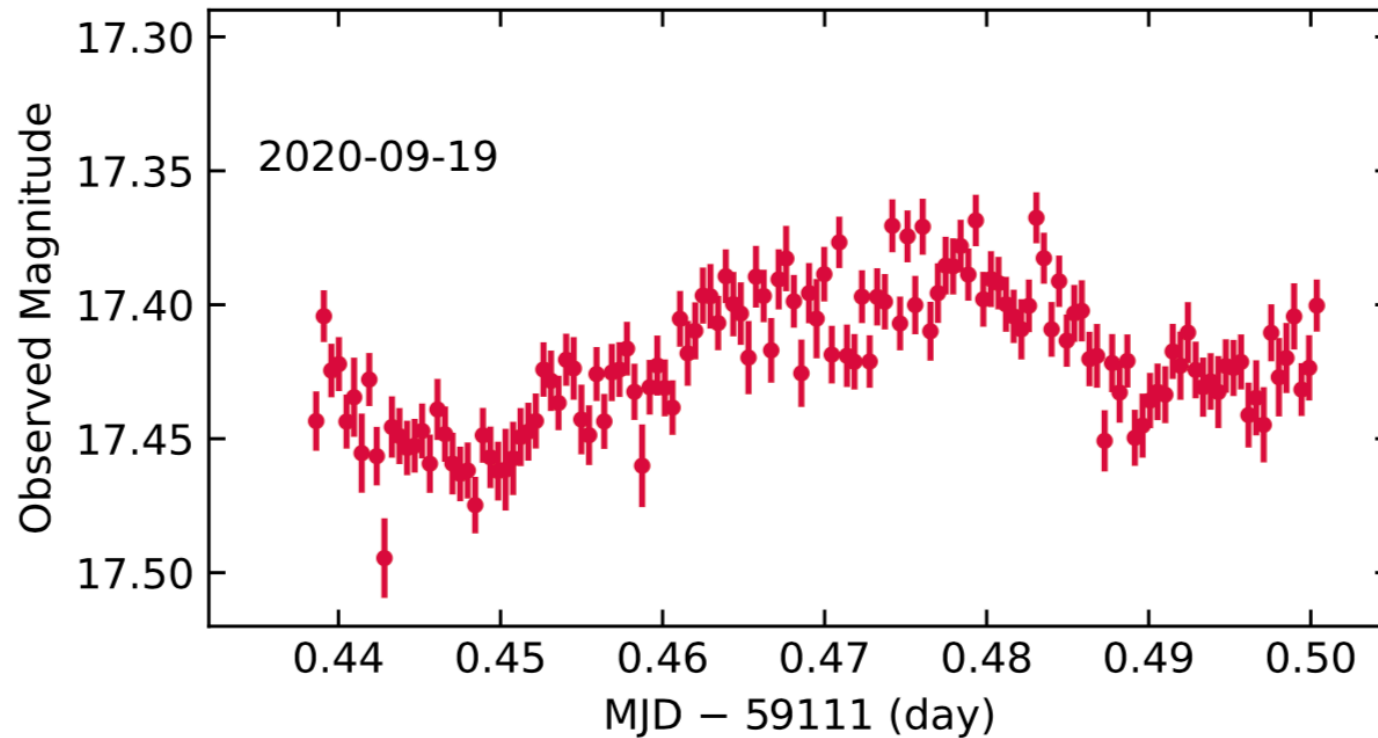
Nearly *Constant in Optical* Despite *Significant Brightening in Radio/ X-ray!!!*

Brightening: X-ray $\uparrow > 20$, radio $\uparrow > 10$, UV $\uparrow \sim 3$, Opt $\uparrow < 1.3$

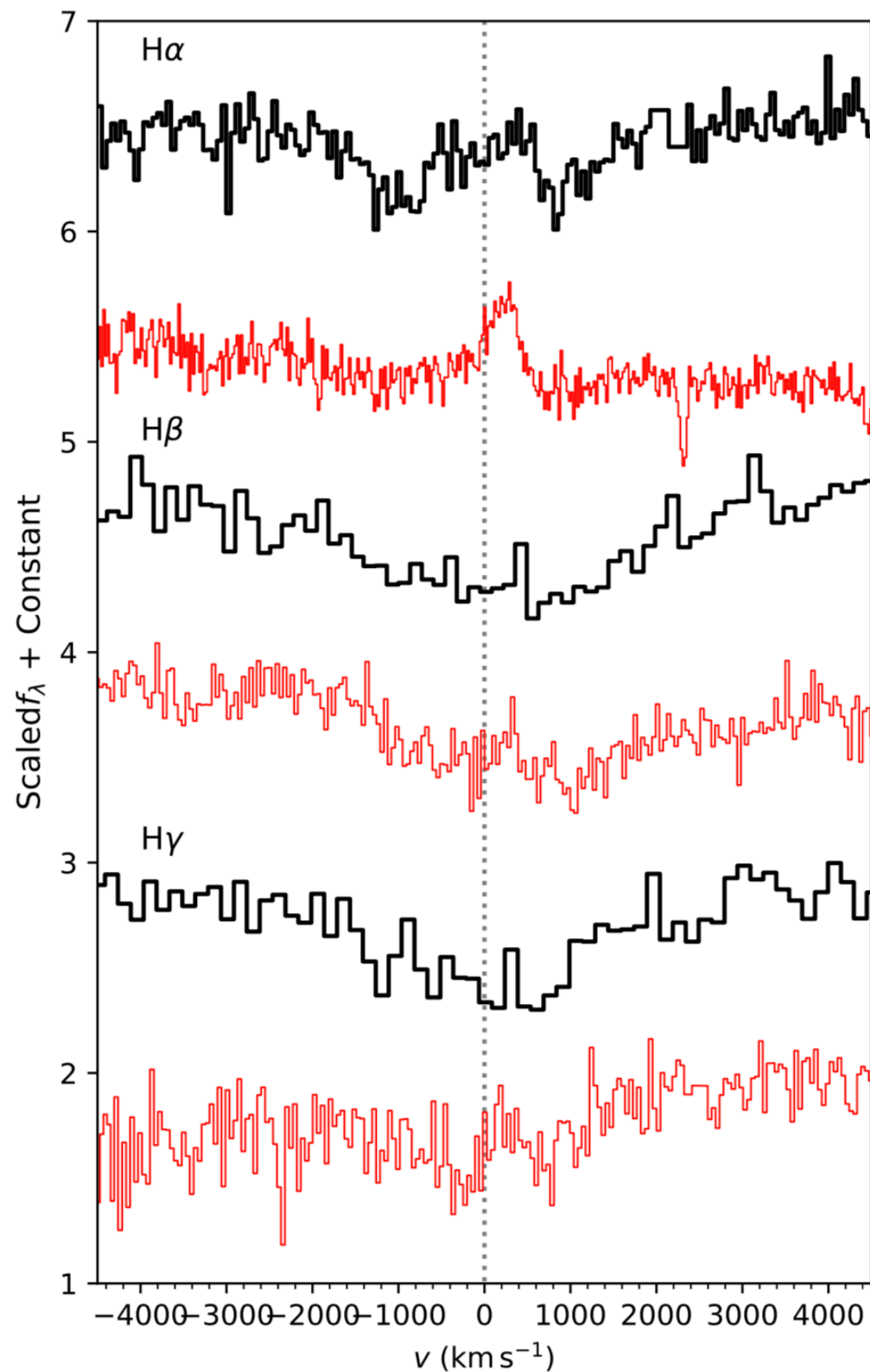


Orbital Modulation at 1.3 hr? Not conclusive...

ZTF deep drill observation



Optical: Variable Balmer

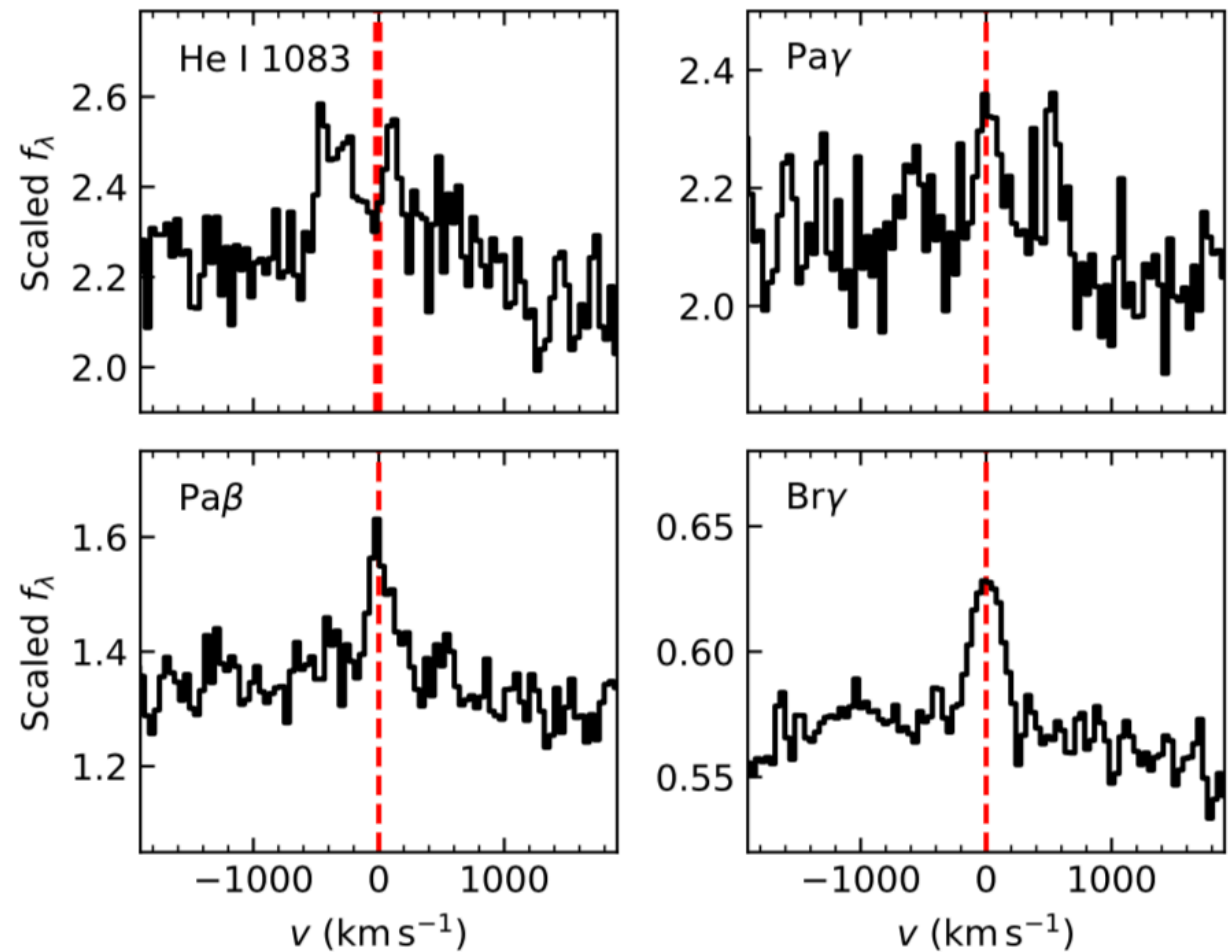


2020-03-23
KeckI-LRIS

2020-09-12
KeckII-ESI

- Broad emission and absorption lines
- Observed in LMXB BHB & DN

NIR: Helium and Hydrogen



Summary

- AT2019wey is a LMXB discovered by ATLAS in optical & *SRG* in X-ray
- Its X-ray spectral and timing properties are consistent with LMXBs
- It has been in the X-ray hard state since discovery
- Its optical emission is independent of X-ray/radio, remained ~constant, which has never been seen in known LMXBs
- Its uniqueness broadens our knowledge for the Galactic LMXB population

Thank you!