



# Do I care about this thing?

## A PS1 star/galaxy catalog for ZTF



AAM+17a

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#### Seriously Adam - Do We Need This?!



PTF "holes" in coverage

single filter

large depth variations

credit: PTF

PS1 - full 3pi coverage 5 filter coverage deep, "good" seeing



Chambers+16

#### "Classic" s/g Method



credit: PS1 website

## (Brief) Intro to Machine Learning



#### HST/PS1

~75k sources

faint sources

small area (1 sq deg)

"unbiased"

morphological classifications

#### SDSS spec/PS1

~3.5M sources

spec limit ~21 mag

large footprint

target selection bias

spectroscopic classifications







#### "Simple" Model



Tachibana & Miller 18

#### "Simple" Model



Tachibana & Miller 18

#### **Creating Features**

Shape parameters from PS1 stack images

white feat =  $\frac{\sum_{f}^{g,r,i,z,y} \text{feat}_{f} \times \text{SNR}_{f}^{2} \times \delta_{f}}{\sum_{f}^{g,r,i,z,y} \text{SNR}_{f}^{2}}$ 

### **Creating Features**



#### 11 "white" features

#### reduce color dependency

#### no mag measurements

#### **Figure of Merit** Maximize TPR @ FPR = 0.005



Kasliwal+16

#### Results



#### Results



Blah Blah Frickin Blah

Tell me what I need to know



#### ~1 >> likely star

#### = 0.5 >> absolutely no idea

## ~0 >> likely galaxy

Note - all bets are off near Galactic plane



| FPR   | TPR  | s/g score |
|-------|------|-----------|
| 0.005 | 0.7  | 0.76      |
| 0.01  | 0.74 | 0.65      |
| 0.02  | 0.79 | 0.53      |
| 0.05  | 0.85 | 0.36      |
| 0.10  | 0.90 | 0.24      |



(imperfect) ML models are superior to low-dimensional cuts Never, ever, ever, ever, ever use hard cuts

PS1 provides ideal dataset for ZTF s/g separation  $3\pi$ , 5 filters (red coverage very important), deep, high quality

HST training set > SDSS training set Less biased

Model optimized for LIGO follow up FoM = TPR @ FPR = 0.005 (we can provide you with alternate thresholds)

70% of stars removed while rejecting only ~0.5% of galaxies