# **SCHEDLLING ZTF**

Eric Bellm March 19, 2018

## What is the scheduling problem we're trying to solve?

Decide what fields to observe in what order. while maintaining the desired cadences while maximizing image quality while avoiding the moon while minimizing slew time while minimizing filter changes while interleaving multiple observing programs while equalizing observing time between programs while responding to TOOs while recovering from weather losses

45 seconds per exposure over a 10-hour night: 800 exposures/night  $\Rightarrow$  800!  $\approx$  10<sup>1977</sup> possible orderings

## NP hard!

# ZTF is conducting an unusually large number of surveys simultaneously.

# MSIP

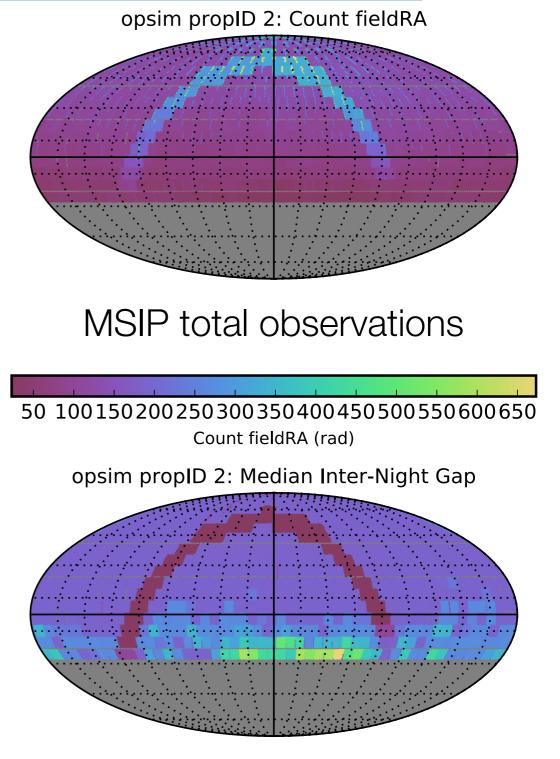
Northern Sky Survey Galactic Plane Survey

#### Partnership

High Cadence i-band HiCAPS Twilight Survey SFR survey TOOs

# Caltech

multiple rotating programs...



MSIP cadence

3

### The simplest scheduler is a static list.

[(time, field, filter, exposure time), (time, field, filter, exposure time)]

### Pros

conceptually simple technically robust

# Cons

labor intensive hard to reproduce

#### Greedy schedulers always take the next best option.

Recompute score before each new observation and observe the field with the highest value.

[(field, score), (field, score)]

### Greedy schedulers always take the next best option.

Recompute score before each new observation and observe the field with the highest value.

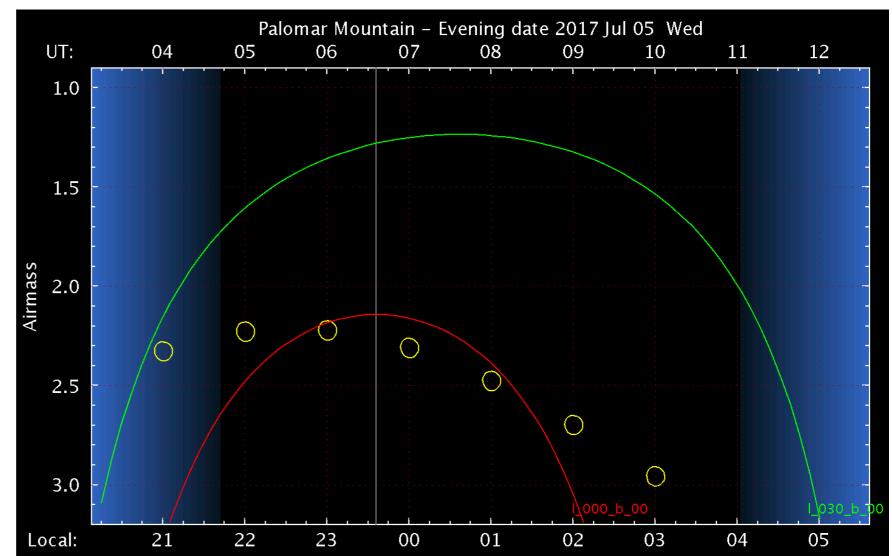
[(field, score), (field, score)]

#### Pros

simple to implement provides some optimization

# Cons

no lookahead; cannot optimize globally



## We can optimize ZTF for discovery.

Maximize volume surveyed per image:

 $V = \frac{4\pi}{3} d^3$   $\propto 10^{0.6m_{\text{lim}}} \qquad \text{(to maximize SNR, use 10^{0.8m_{\text{lim}}})}$ 

Limiting magnitude depends on: filter, sky brightness, airmass, seeing

So: maximize the *volume-weighted* number of images observed in acceptable cadence windows.

## We can optimize ZTF for discovery.

Maximize volume surveyed per image:

 $V = \frac{4\pi}{3}d^3$  $\propto 10^{0.6m_{
m lim}}$ (to maximize SNR, use  $10^{0.8m_{\text{lim}}}$ ) Limiting magnitude depends on: filter, sky brightness, airmass, seeing objective function optimization algorithm So: maximize the volume-weighted number of images observed in acceptable cadence windows.

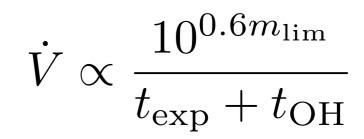
observing strategy

## We can optimize ZTF for discovery.

Maximize volume surveyed per image:

 $V = \frac{4\pi}{3}d^3$ (to maximize SNR, use  $10^{0.8m_{\text{lim}}}$ )  $\propto 10^{0.6m_{\text{lim}}}$ Limiting magnitude depends on: filter, sky brightness, airmass, seeing objective function optimization algorithm So: maximize the volume-weighted number of images observed in acceptable cadence windows. observing strategy

For greedy schedulers, choose highest



## We compute volume-weighting with a predictive model.

compute sky brightness (-> limiting mag) xgboost model trained on PTF using the following features: filter (g, r, or i) sun altitude altitude, azimuth moon illumination fraction moon altitude moon distance

# Acceptable cadence windows are defined per observing program.

Each observing (sub-)program specifies

- a survey footprint (either as discrete field ids or a selection function)
- a minimum number of days between revisits
- a "request set" per field: the number of observations per filter to be observed tonight
- the fraction of total time to be devoted to this subprogram

Intra-night separations are set to be simply larger than a fixed interval (30 minutes) to enable asteroid discrimination (previous efforts to define on-off cadence windows based on time since last observations proved too brittle.)

Each night, each program proposes request sets according to their observability and inter-night cadence, and the scheduler constructs an optimal observing plan from among them.

#### The MSIP observing programs provide a simple example.

```
{"program_name": "MSIP",
"subprogram_name": "all_sky",
"program_observing_fraction": 0.4,
"subprogram_fraction": 0.85,
"field_selections":{"dec_range":[-31,80],
             "abs_b_range":[7,90],
             "grid_id":0},
"filter_ids": [1, 2],
"internight_gap_days": 3,
"n_visits_per_night": 2},
{"program_name": "MSIP",
"subprogram_name": "nightly_plane",
"program_observing_fraction": 0.4,
"subprogram_fraction": 0.15,
"field_selections":{"dec_range":[-31,80],
             "abs_b_range":[0,7],
             "grid_id":0},
"filter_ids": [1, 2],
 "internight_gap_days": 1,
 "n visits per night": 2}
```

Request Sets (Fields)

#### **Time Blocks**

13

	to	t <sub>1</sub>	t <sub>2</sub>	t <sub>3</sub>	t4	<b>t</b> 5	t <sub>6</sub>	t7
r <sub>o</sub> (gggg)								
r <sub>1</sub> (gggg)								
r <sub>2</sub> (©r)								
r <sub>3</sub> (©r)								
r4 (i)								
r <sub>5</sub> (oro)								
r <sub>6</sub> ( <b>rgr</b> )								

# We use Integer Programming techniques to perform nightly optimization.

 $V_{rtf}$  Volume factor for request field r at time t in filter f  $Y_{rtf}$  ("yes") =1 if we observe r at t in f, 0 otherwise



$$\sum_{r \in R} \sum_{t \in T} \sum_{f \in F} V_{rtf} Y_{rtf}$$



subject to

$$\sum_{t \in T} Y_{rtf} \le n_{rf} \ \forall \ r \in R, f \in F$$

number of requests in this set

$$\sum_{r \in R} Y_{rtf} \le n_{\max} \ \forall \ t \in T$$

number of observations in this slot

And enforce one filter per slot + program balance

#### **Time Blocks**

	to	t <sub>1</sub>	t <sub>2</sub>	t <sub>3</sub>	t4	<b>t</b> 5	t <sub>6</sub>	t7
r <sub>0</sub> (gggg)								
r <sub>1</sub> (gggg)								
r <sub>2</sub> (gr)								
r <sub>3</sub> (gr)								
r4 (i)								
r <sub>5</sub> (org)								
r <sub>6</sub> ( <b>r</b> gr)								

#### **Time Blocks**

	to	t <sub>1</sub>	t <sub>2</sub>	t <sub>3</sub>	t4	t <sub>5</sub>	t <sub>6</sub>	t7
r <sub>0</sub> (gggg)								
<b>r</b> 1 (gggg)								
r <sub>2</sub> (gr)								
r <sub>3</sub> (gr)								
r <sub>4</sub> (i)								
r <sub>5</sub> (grg)								
r <sub>6</sub> (rgr)								



		to	t <sub>1</sub>	t <sub>2</sub>	t <sub>3</sub>	t4	t <sub>5</sub>	t <sub>6</sub>	t7
r <sub>o</sub> (gggg)	Y <sub>0</sub> =1								
r <sub>1</sub> (gggg)	Y <sub>1</sub> =1								
r <sub>2</sub> ( <b>gr</b> )	Y <sub>2</sub> =1								
r <sub>3</sub> ( <b>gr</b> )	<b>Y</b> 3 <b>=0</b>								
r4 (i)	Y <sub>4</sub> =1								
r <sub>5</sub> (grg)	Y <sub>5</sub> =1								
r <sub>6</sub> (ror)	Y <sub>6</sub> =0								

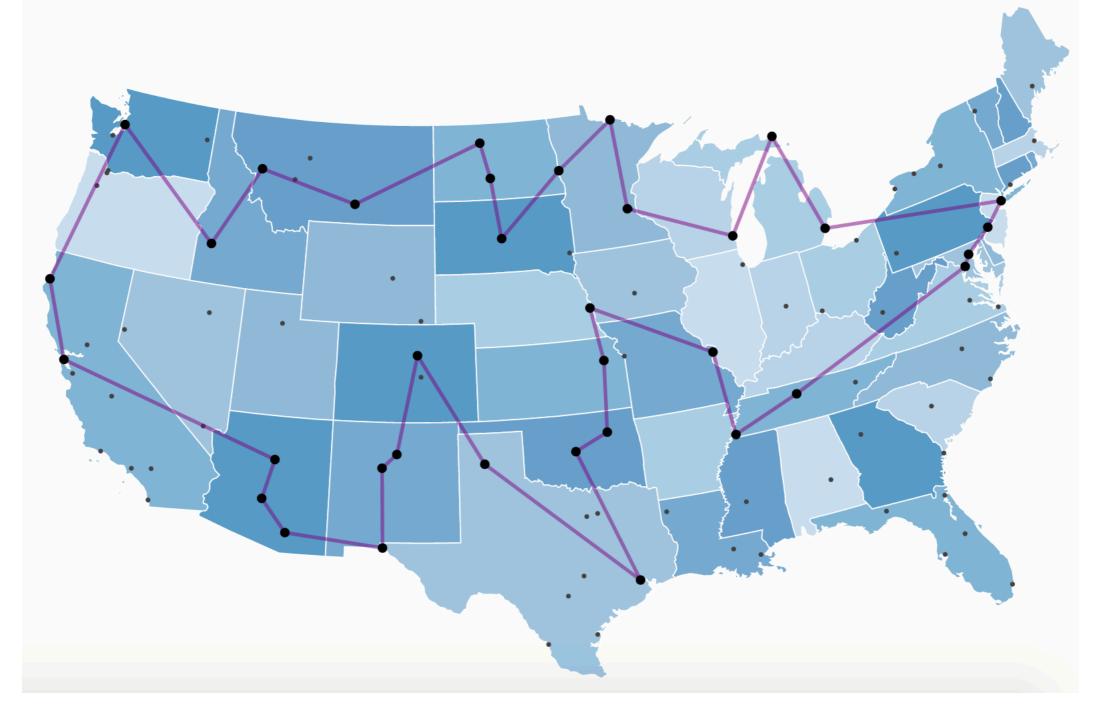


		to	t <sub>1</sub>	t <sub>2</sub>	t <sub>3</sub>	t4	t <sub>5</sub>	t <sub>6</sub>	t7
		Ŋ	r	g	g	i	Ŋ	g	r
r <sub>o</sub> (gggg)	Y <sub>0</sub> =1			Х	Х		Х	Х	
r <sub>1</sub> (gggg)	<b>Y</b> 1=1	Х		Х	Х		Х		
r <sub>2</sub> (gr)	Y <sub>2</sub> =1							Х	
r <sub>3</sub> (gr)	<b>Y</b> 3 <b>=0</b>								
r <sub>4</sub> (i)	Y <sub>4</sub> =1					Х			
r <sub>5</sub> (grg)	Y <sub>5</sub> =1						Х	Х	Х
r <sub>6</sub> (ror)	<b>Y</b> <sub>6</sub> <b>=0</b>								

Request Sets (Fields)

## Then we sequence each block by solving the TSP.

Distances defined by slew time between requests in this block.



HA and Declination slews don't change with slot, but dome slews do.

## **Optimization requires tradeoffs!**

This scheduler is very good at: scheduling observations near zenith maintaining cadence between nights obtaining the requested number of observations *(in fact, a field is only scheduled if its request set can be fully satisfied!)* maintaining balance between programs minimizing slew time\*

#### This scheduler cannot:

guarantee filter order within the night enforce exact times between observations schedule continuous blocks of observations\* guarantee that there are observations at all times

## I've implemented this scheme in the production system.

#### https://github.com/ZwickyTransientFacility/ztf\_sim

C This repository Se	arch	Pull requests Issues Gist		♠ +• 👰•		
ZwickyTransient	acility / ztf_sim		O Unwatch → 2	Star 1 VFork 2		
<> Code (!) Issues	0 1 Pull requests 0	Projects 0 🗉 Wiki 🔸 Pu	llse 🔟 Graphs 🔅 Set	tings		
🔭 Schedule simulate	or for the Zwicky Transient Fac	ility — Edit				
🗇 97 commits	P 2 branches	So releases	2 contributors	മ് BSD-3-Clause		
Branch: master - New	v pull request	Create	e new file Upload files Find	file Clone or download -		
ebellm bugfixes: mai	ke hard copies of queue dataframes to	o avoid SettingWithCo	Latest co	mmit 6c60688 20 hours ago		
🖿 bin	improve metrics			12 days ago		
🖬 data	add R20 files			2 months ago		
notebooks	script for quick basic analysis of simulations a month ago					
🖬 sims	add .keep file to populate sims di	irectory		2 months ago		
<pre>sims stf_sim</pre>		irectory Jeue dataframes to avoid SettingV	VithCo	2 months ago 20 hours ago		
		ueue dataframes to avoid SettingV	VithCo			
ztf_sim	bugfixes: make hard copies of qu	ueue dataframes to avoid SettingV	VithCo	20 hours ago		

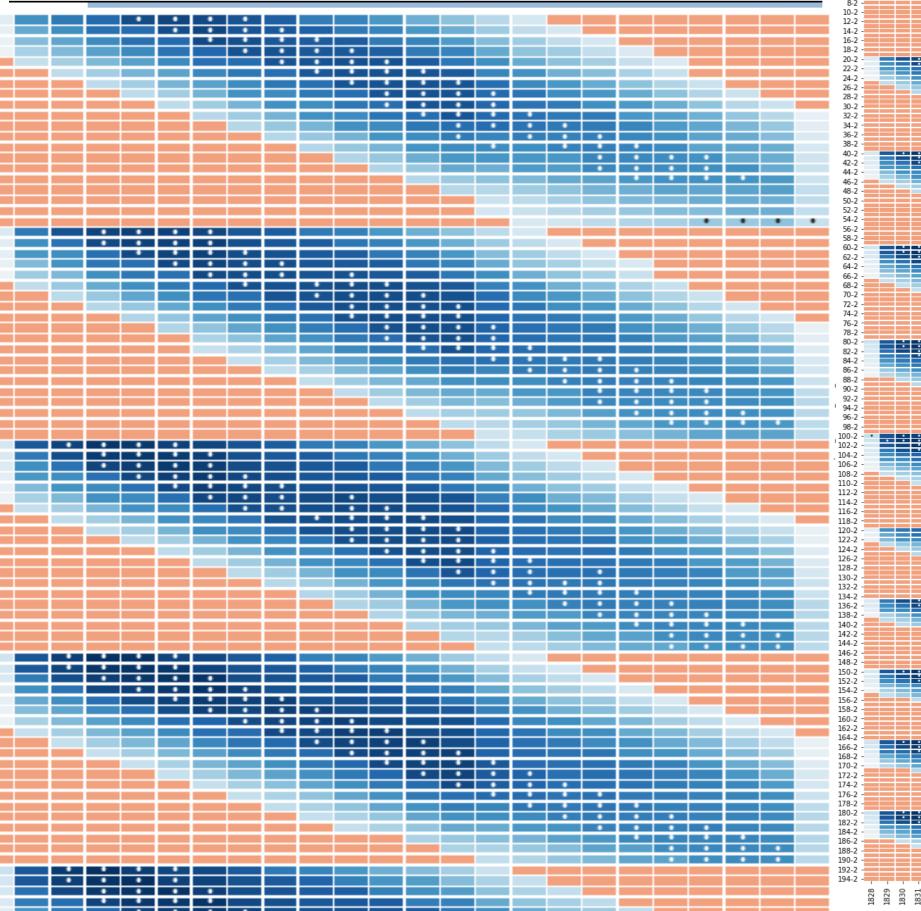
## The core functionality is complete.

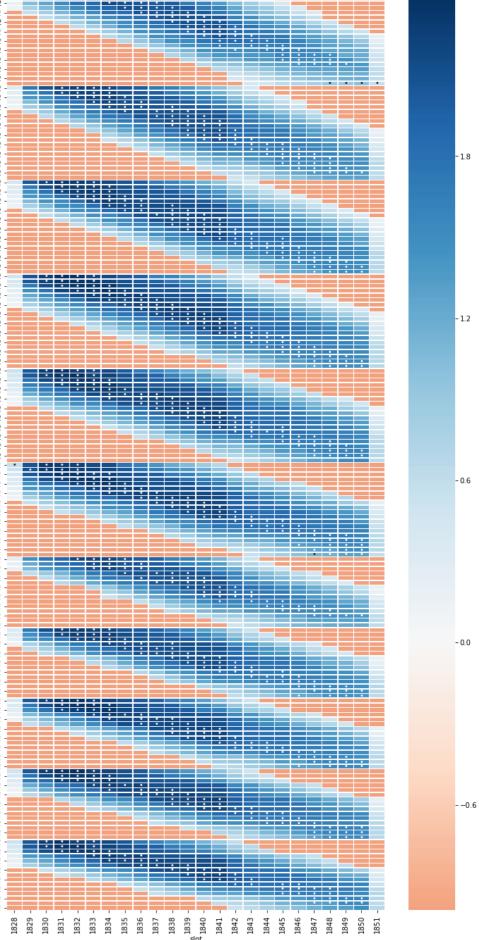
## swappable list, greedy, and slot-based schedulers

runs in simulation mode weather losses (from PTF & iPTF) realistic telescope behavior exports pointing history to LSST MAF

integrated with the ZTF robotic system webserver client that uses the simulator as a library

# We've been running the scheduler throughout commissioning.

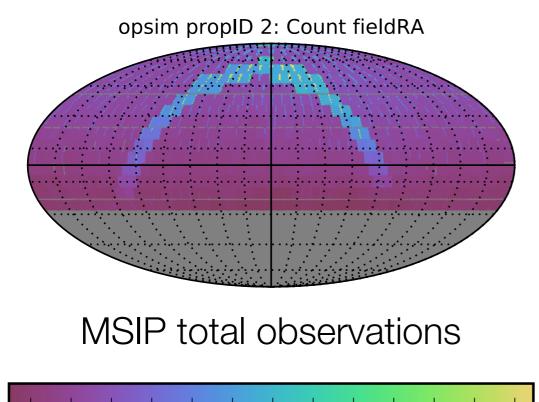




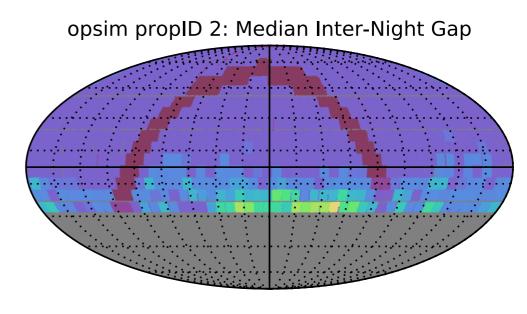
#### A nine-month simulation shows good results.

	Fields	Intra-night cadence	Nightly cadence
MSIP Northern Sky Survey	All fields > Dec -31 &  b  > 7	3 days	2 visits ( <b>gr</b> )
MSIP Galactic Plane	All fields > Dec -31 &  b  < 7	1 day	2 visits ( <b>gr</b> )
Collaboration High-Cadence	72 fields	1 day	6 visits ( <b>gggrrr</b> )
Collaboration i-band	249 fields	4 days	1 visit (i)
Caltech TAC	All fields > Dec -31	1 day	2 visits ( <b>rr</b> )

#### Sky coverage is reasonable.

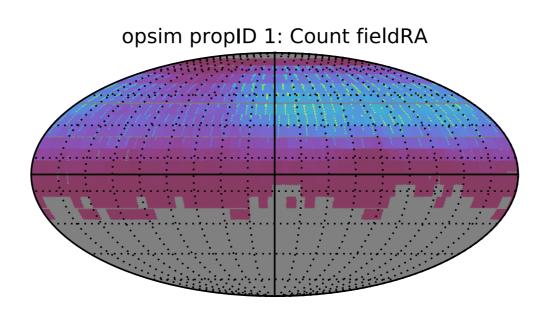


50 100150200250300350400450500550600650 Count fieldRA (rad)



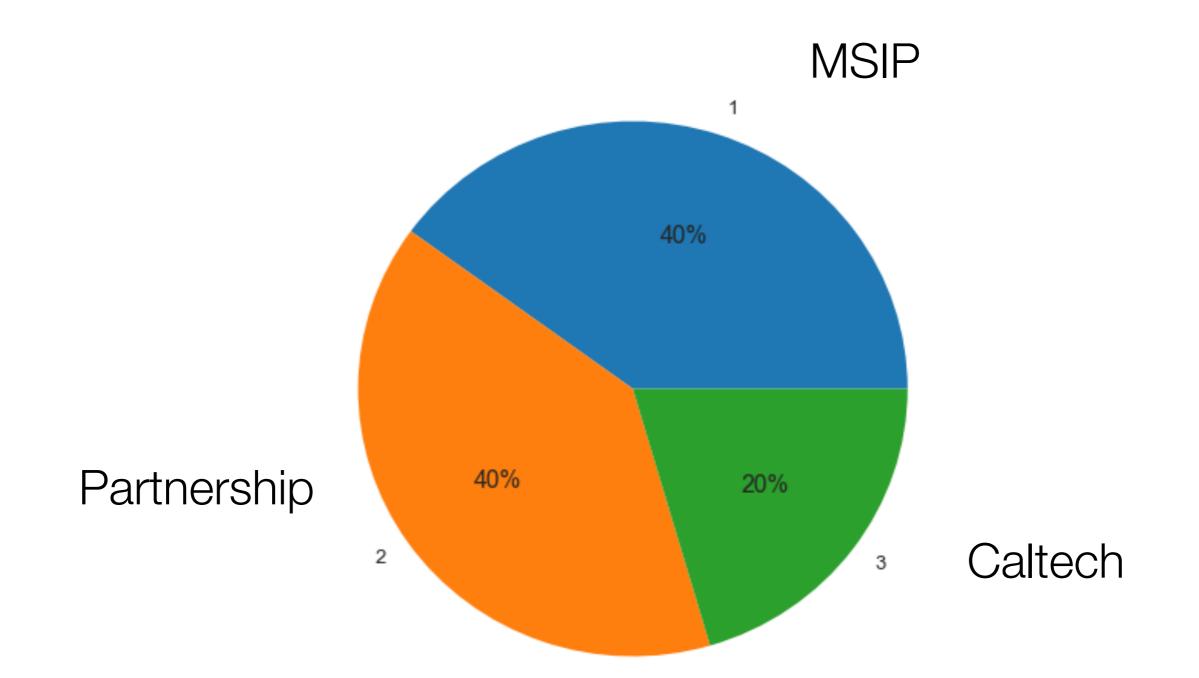
**MSIP** cadence

1.2 1.8 2.4 3.0 3.6 4.2 4.8 5.4 6.0 6.6 7.2 7.8 8.4 Median Inter-Night Gap (days)

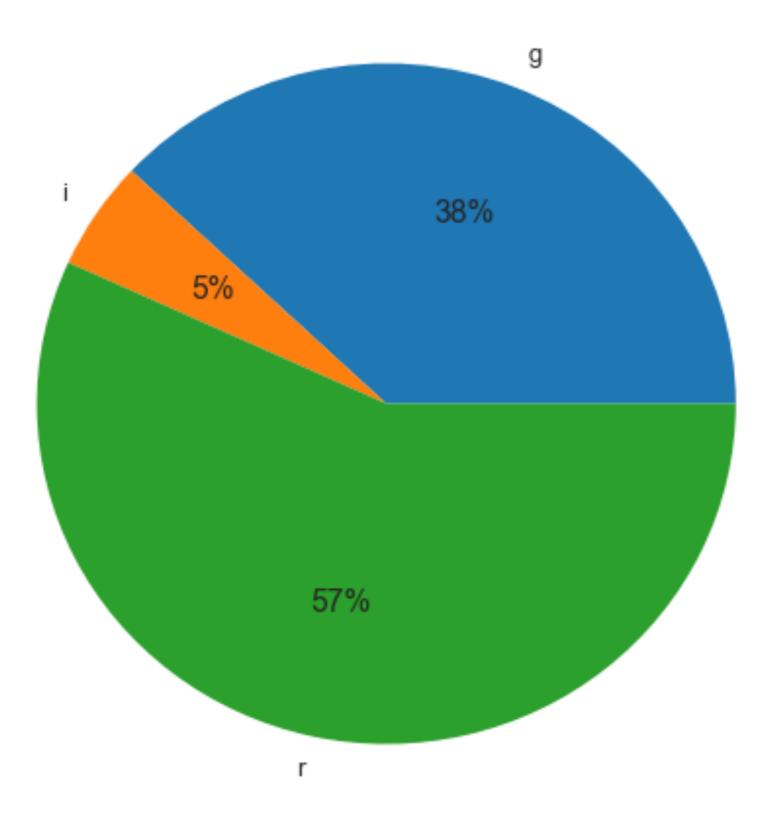


Collaboration total observations

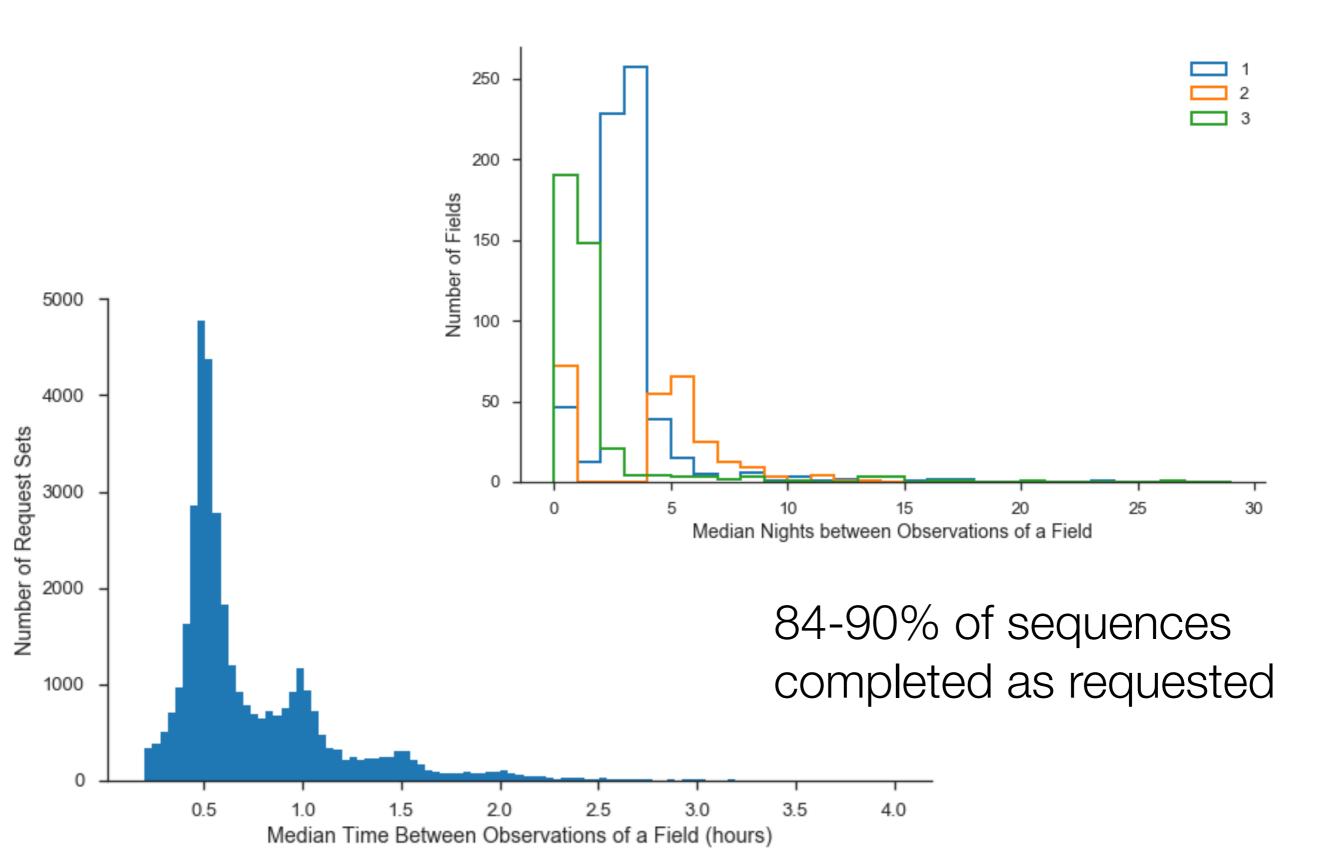
#### Program balance is enforced correctly.



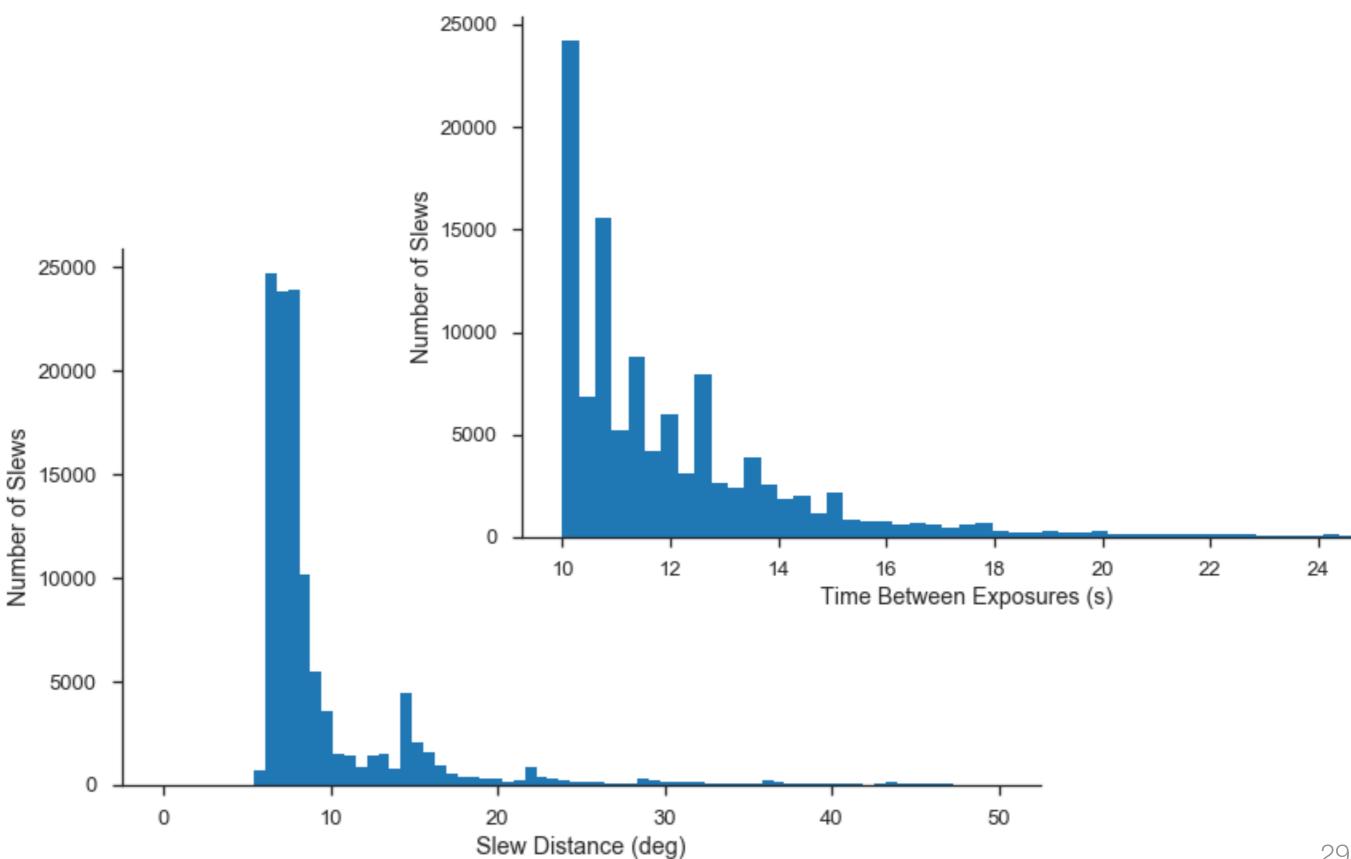
#### i-band filter scheduling needs tweaking.



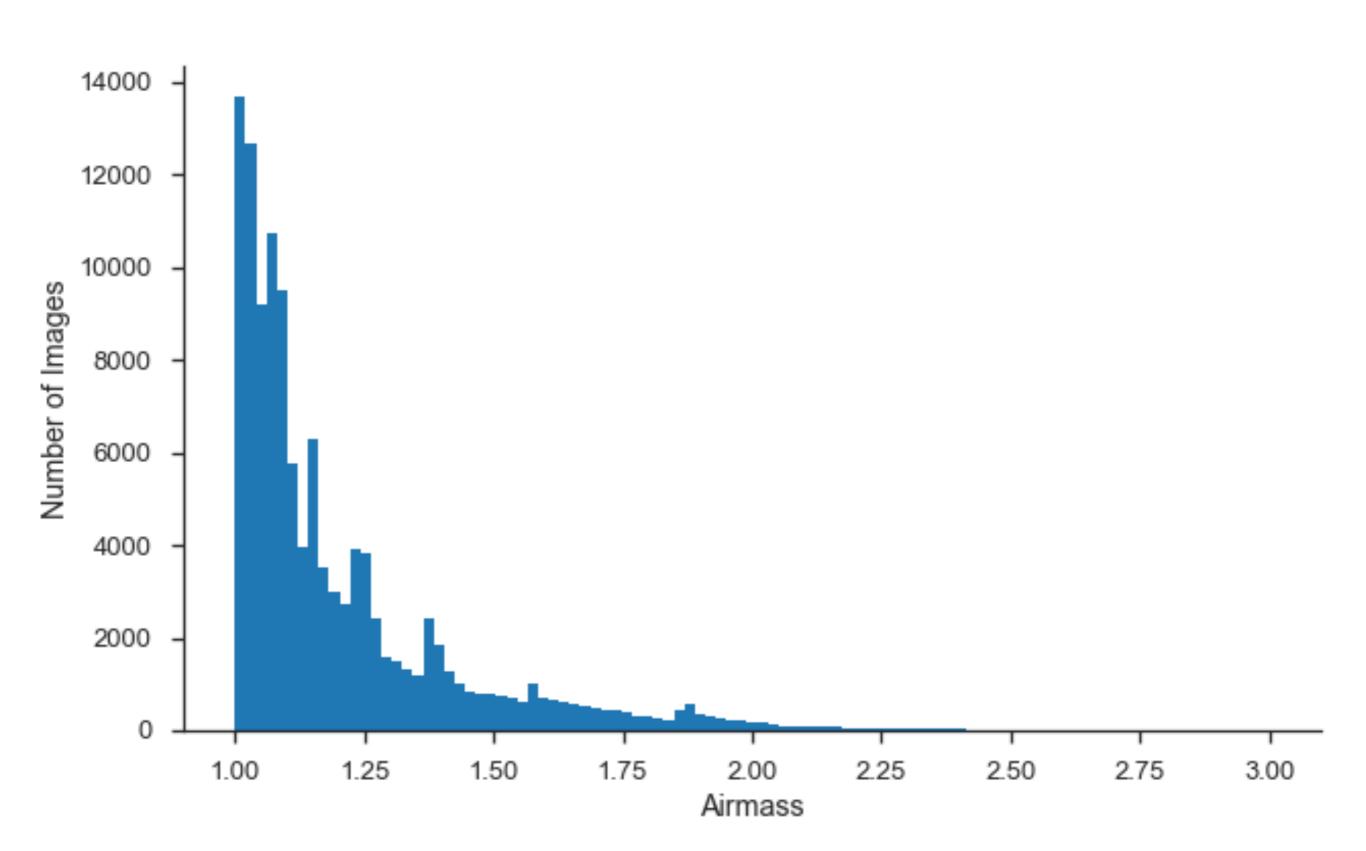
#### Cadences requested are being scheduled correctly.



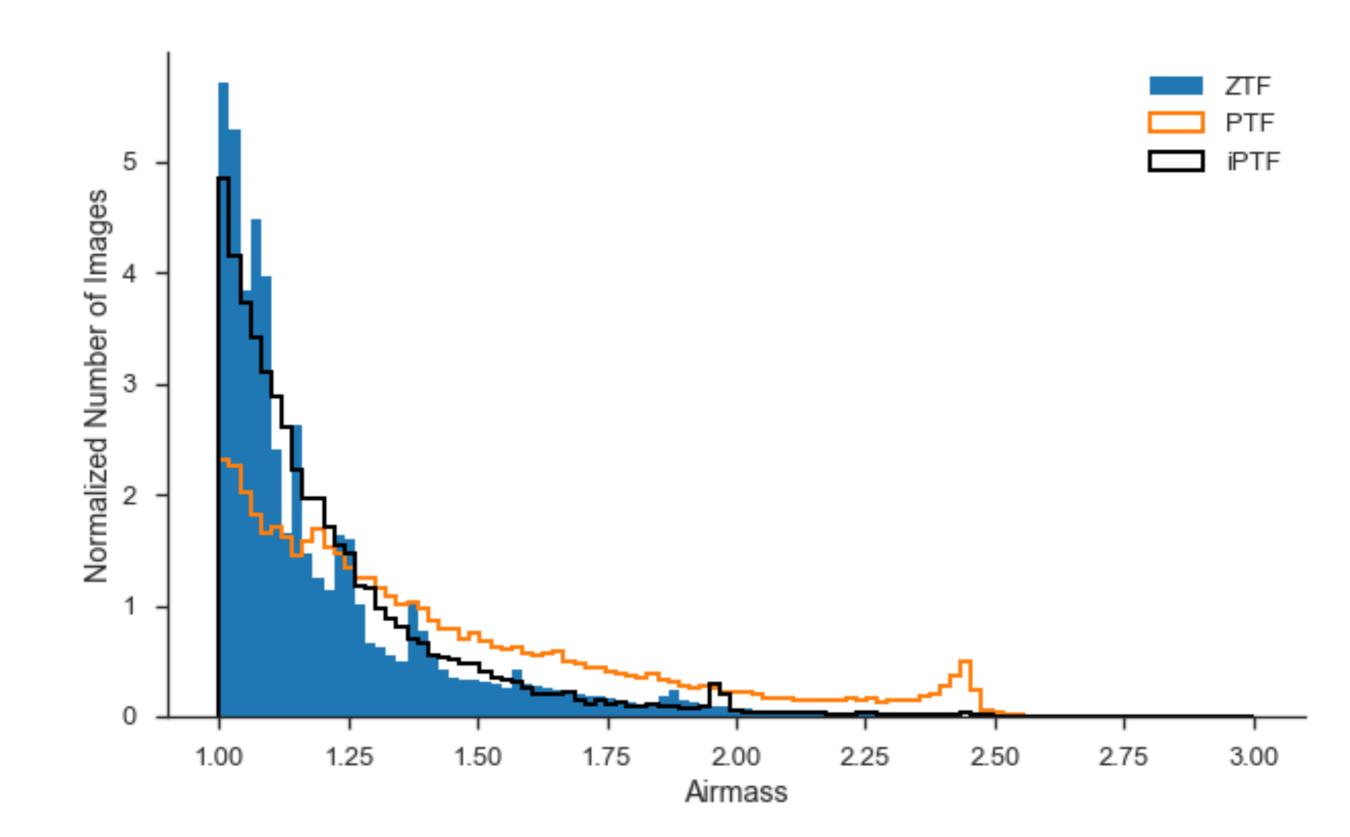
## Slew behavior is reasonable; there is room for additional optimization.



#### The scheduler preferentially observes near zenith.



#### The scheduler preferentially observes near zenith.



#### TOO observations are not yet enabled.

Can combine the volume-probed slot based scheduling with localization probability for each tile

Envision having separate TOO queues that replace the default queue and run until completion within a validity window

Working with TOO marshal team to define the API to talk to the scheduler (which has a REST interface)

Integrate TOOs & special programs as swappable block queues

Model as-delivered ZTF performance

Within-night recomputes and other second-order optimizations

Bugfixes and maintenance!

Help with analysis of simulated or actual schedules is welcome.