



Mickael Rigault | ZTF meeting @ Stockholm | August 2018

Pysedm

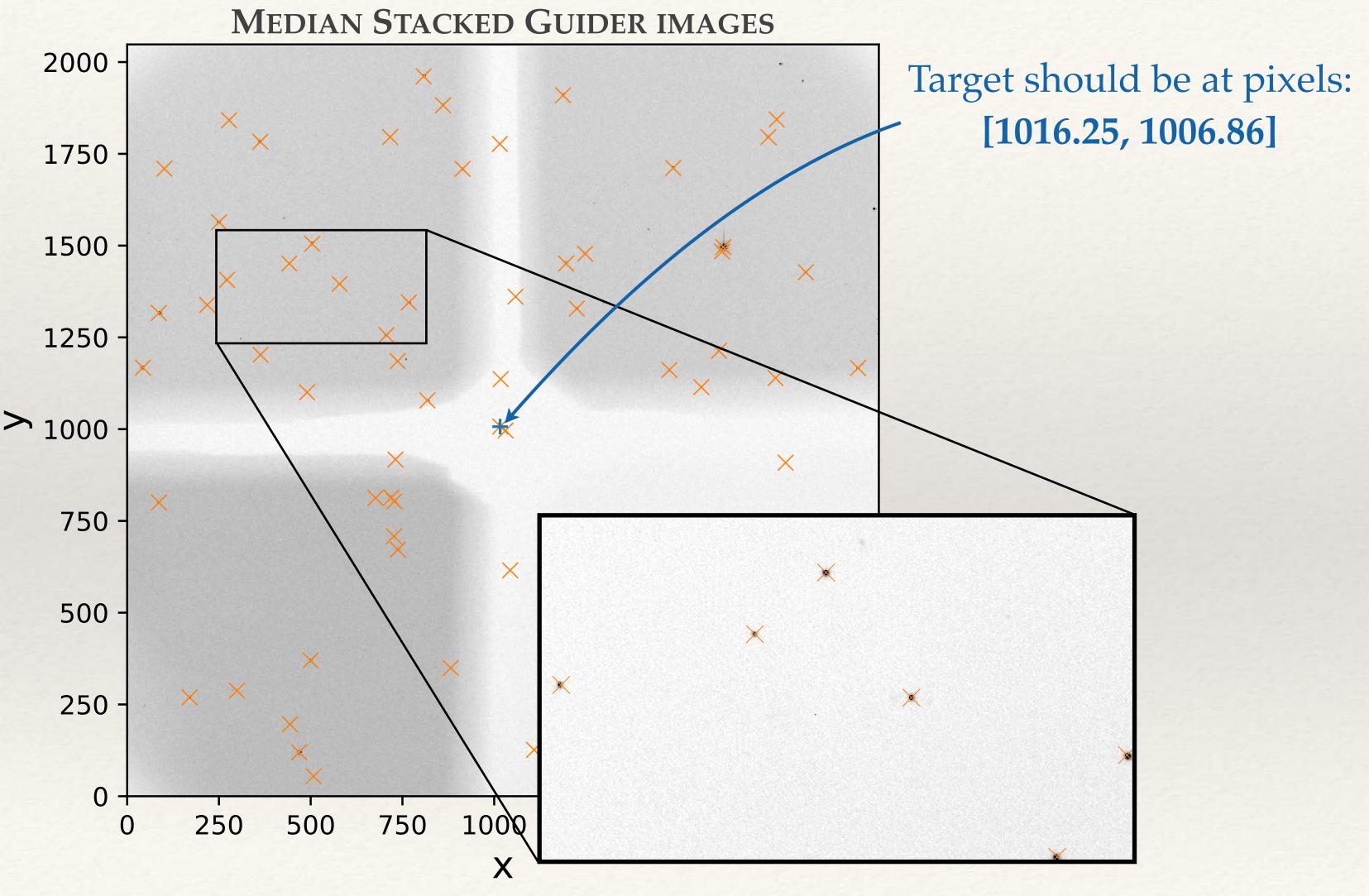
Status report

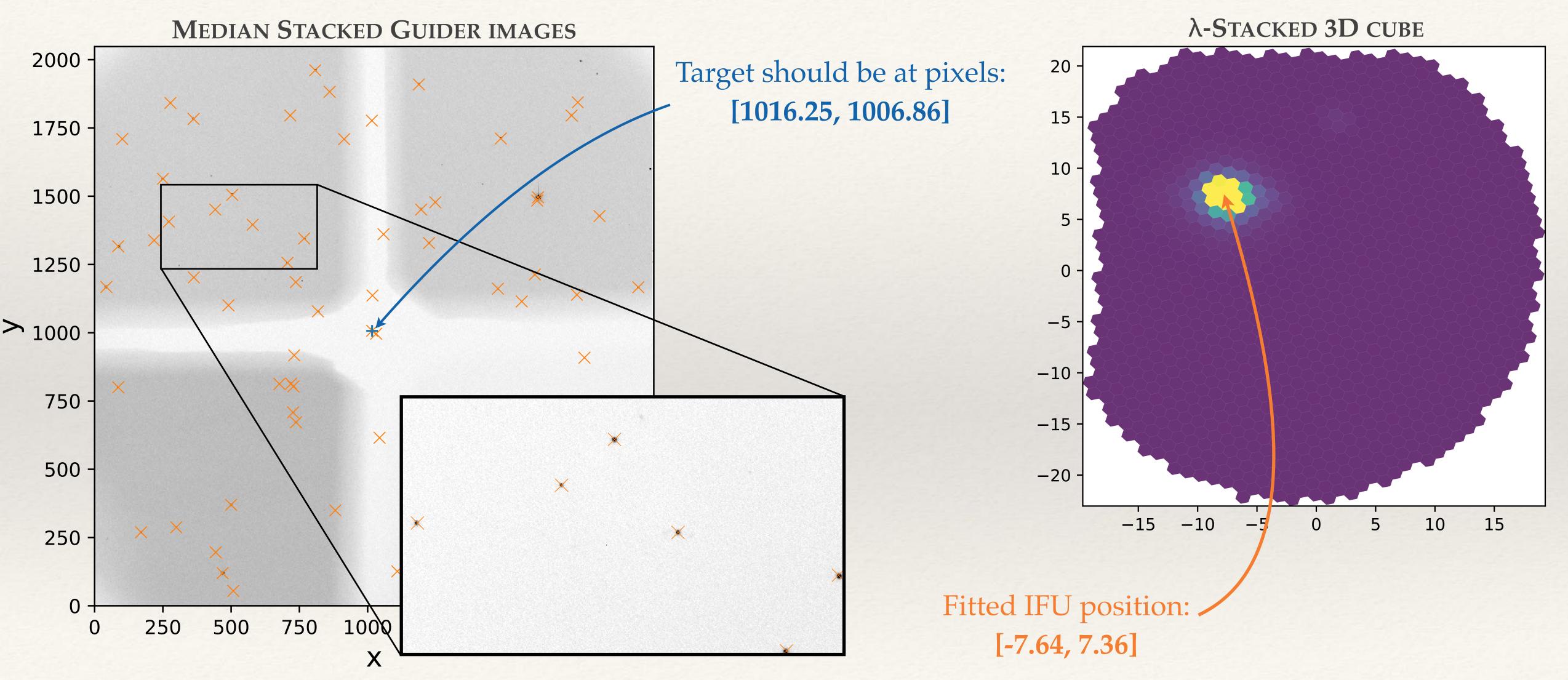
What's new?

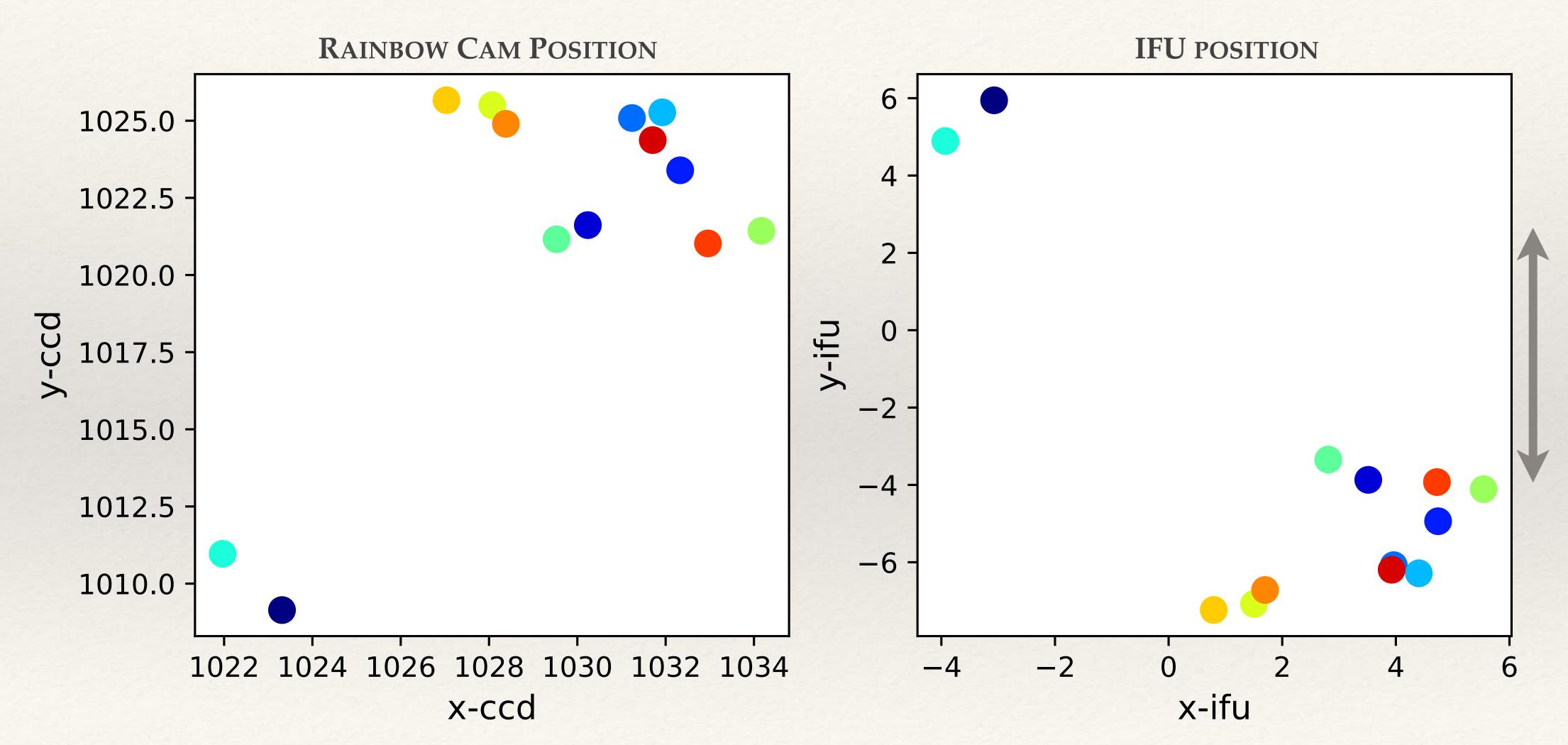
Fully compatible with the new lens-array

Use of guider image (from the rainbow camera) to guess the target position inside the IFU

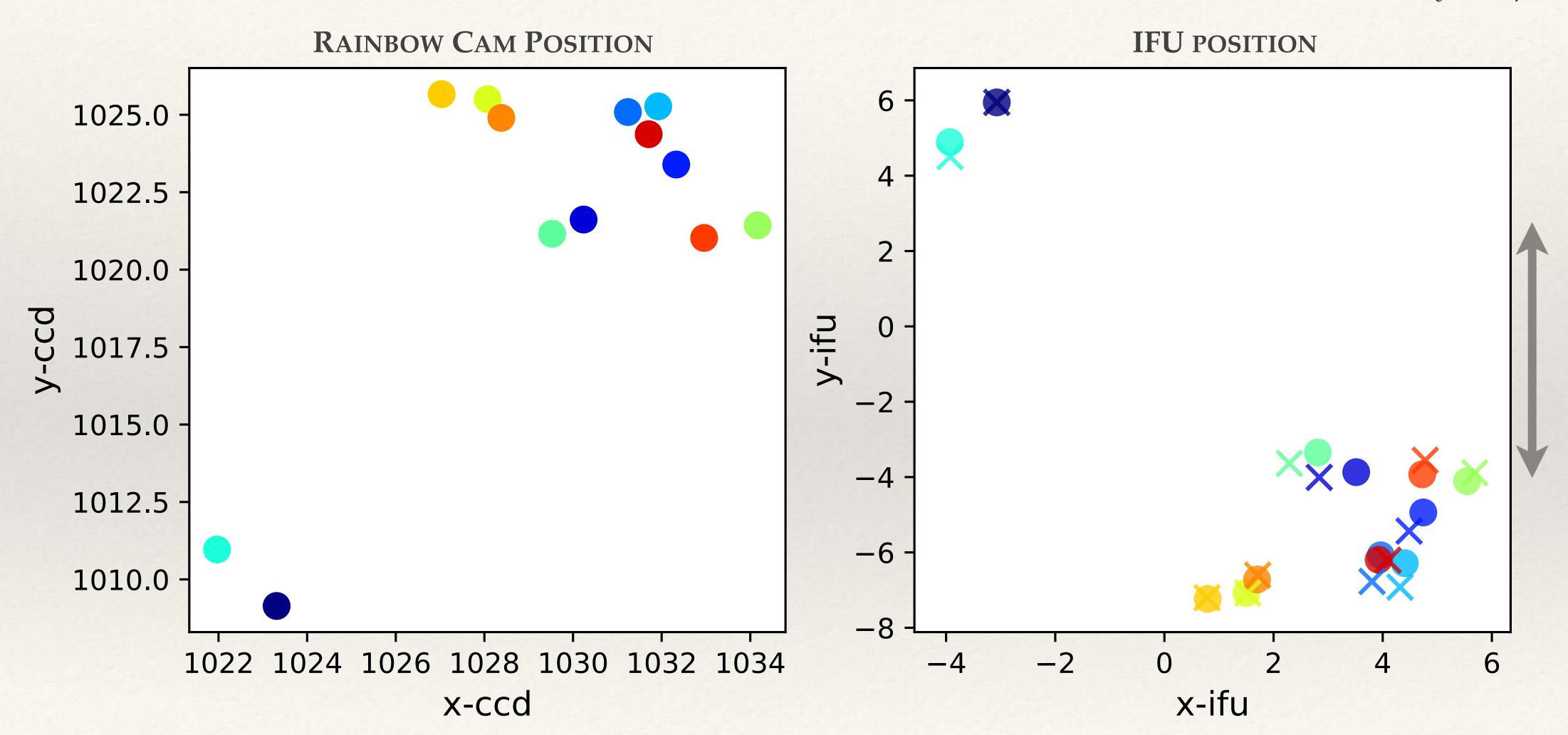
pipeline fully automated
ccd → spectra → typing & Marshall



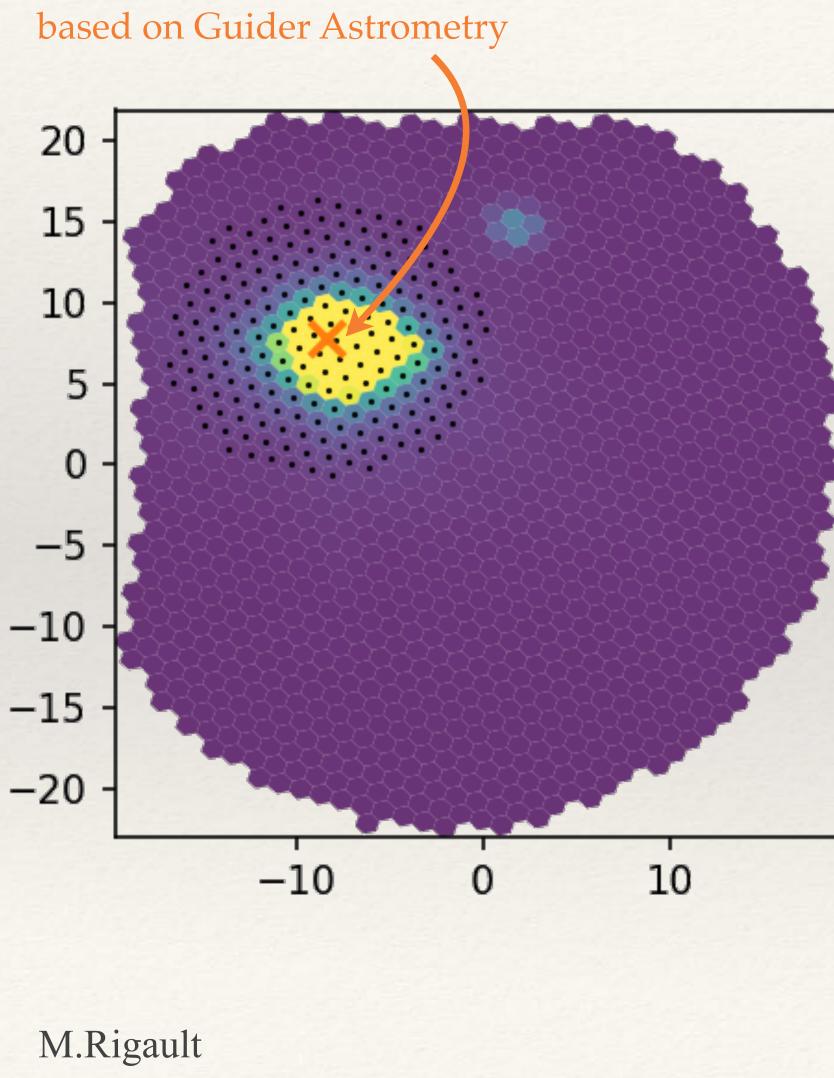


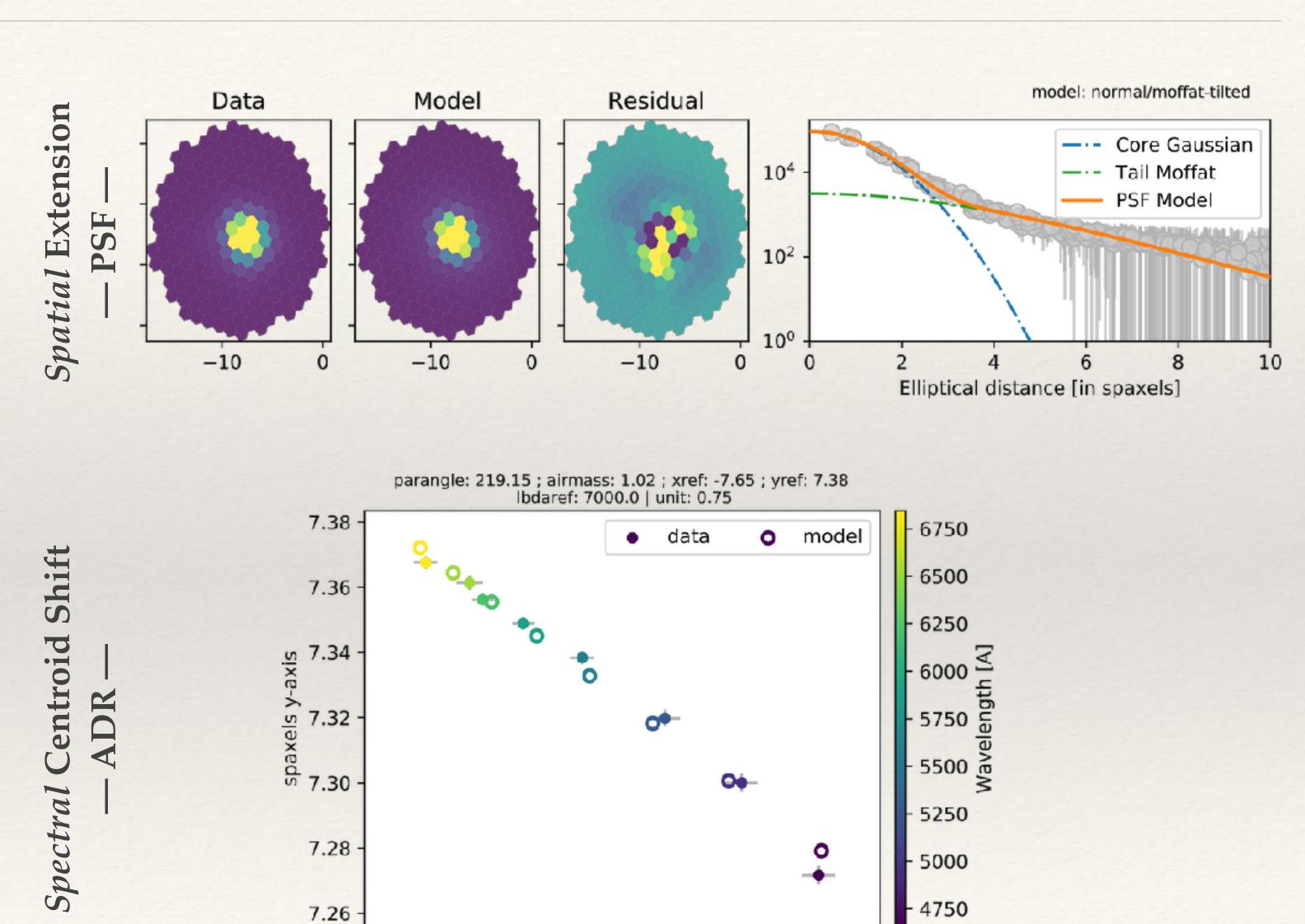


Currently: Simple 2x2 matrix









-7.60

spaxels x-axis

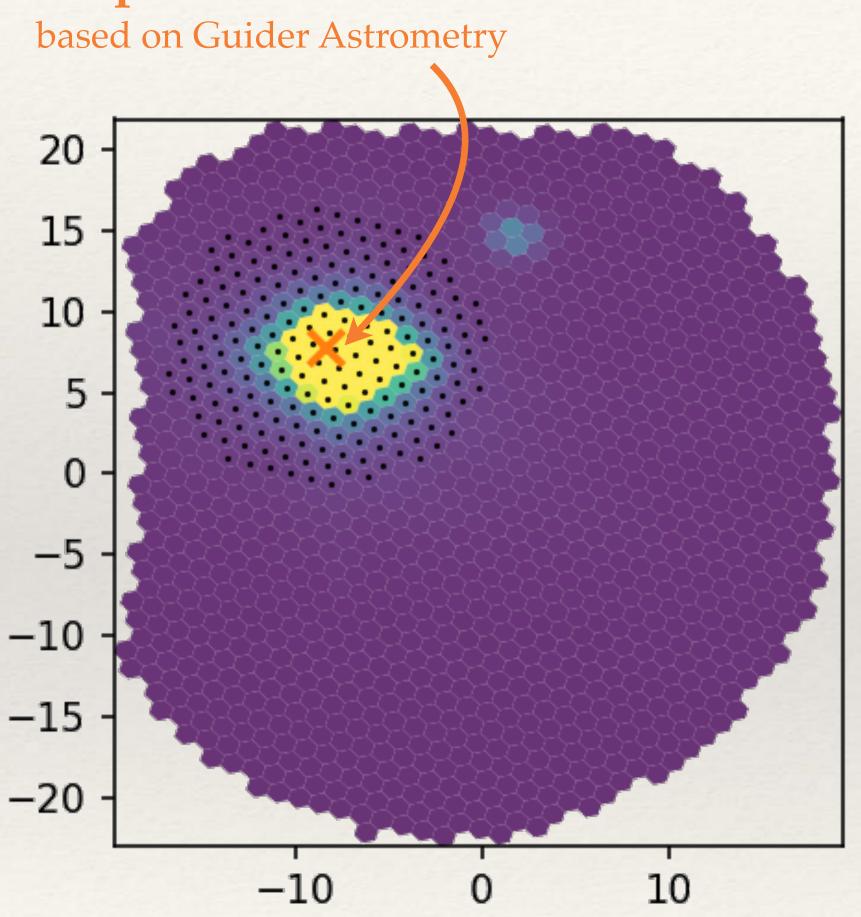
-7.58

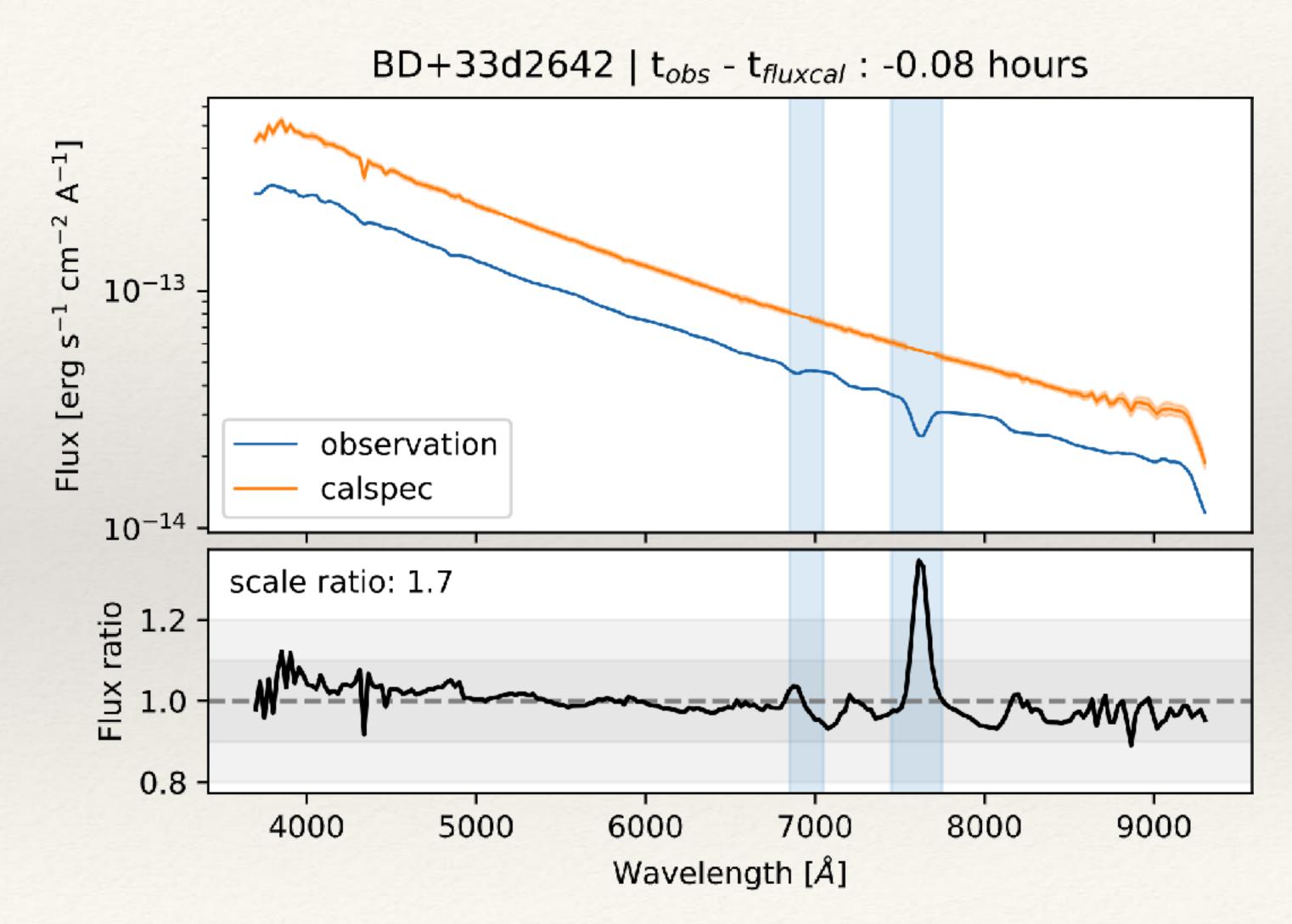
-7.56

-7.62

-7.64

Expected Location





pysedm fully automated pipeline

CCD → 3D cube: ~5min

Flexure corrections
(trace-position & wavelength)

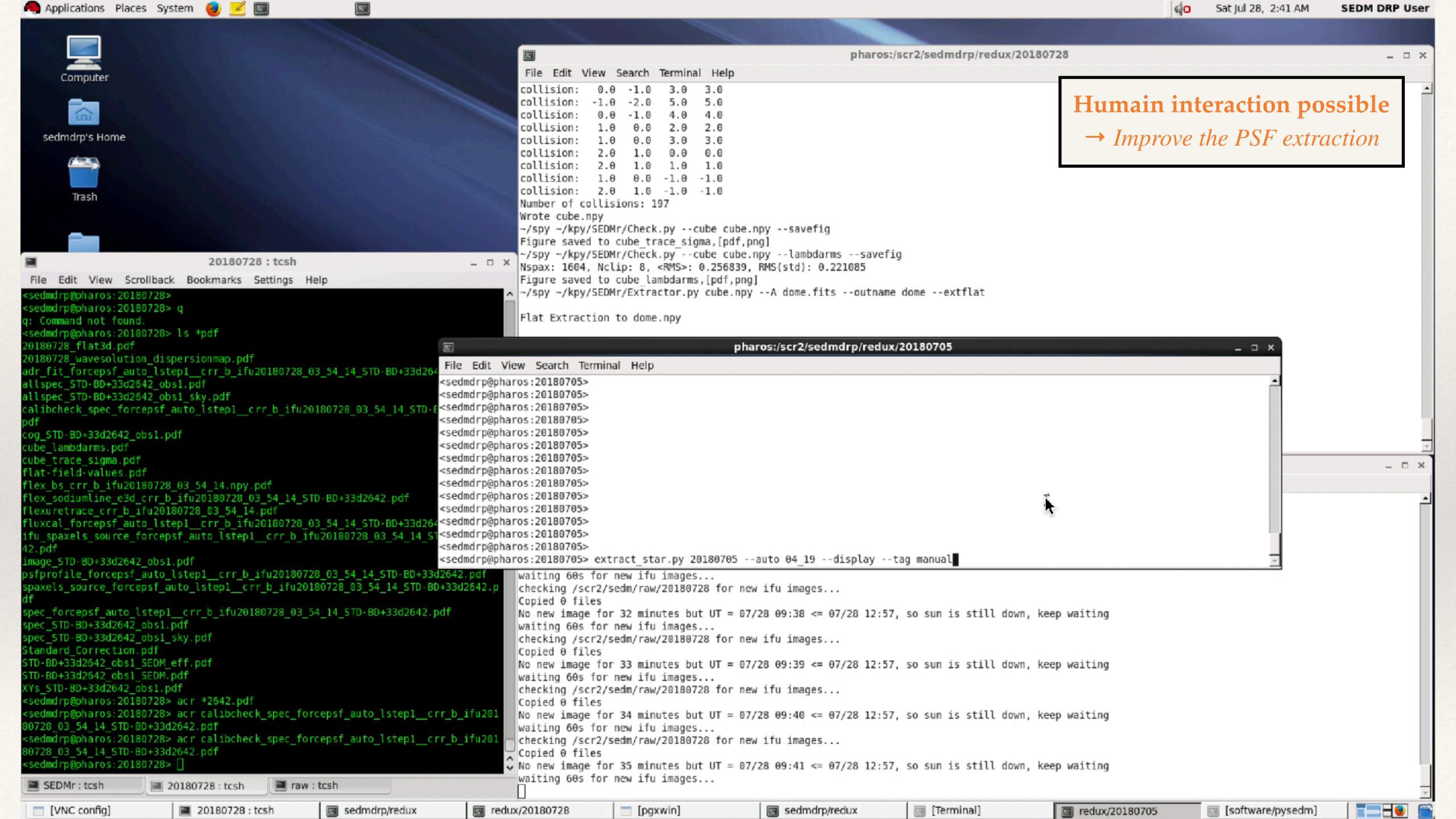
Meta-guider Astrometry (building + solving wcs)

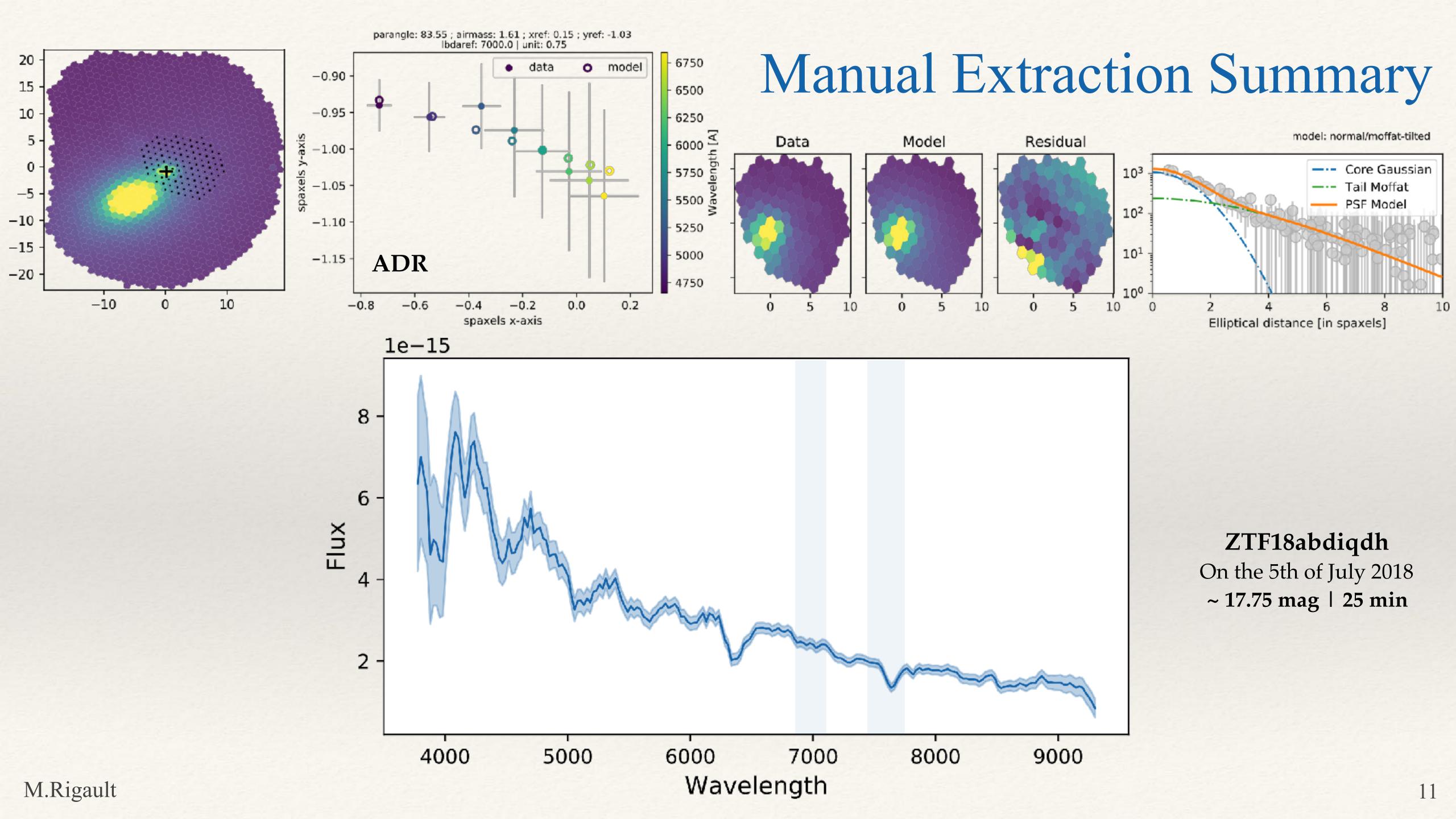
3D cube → Spectra: ~1min

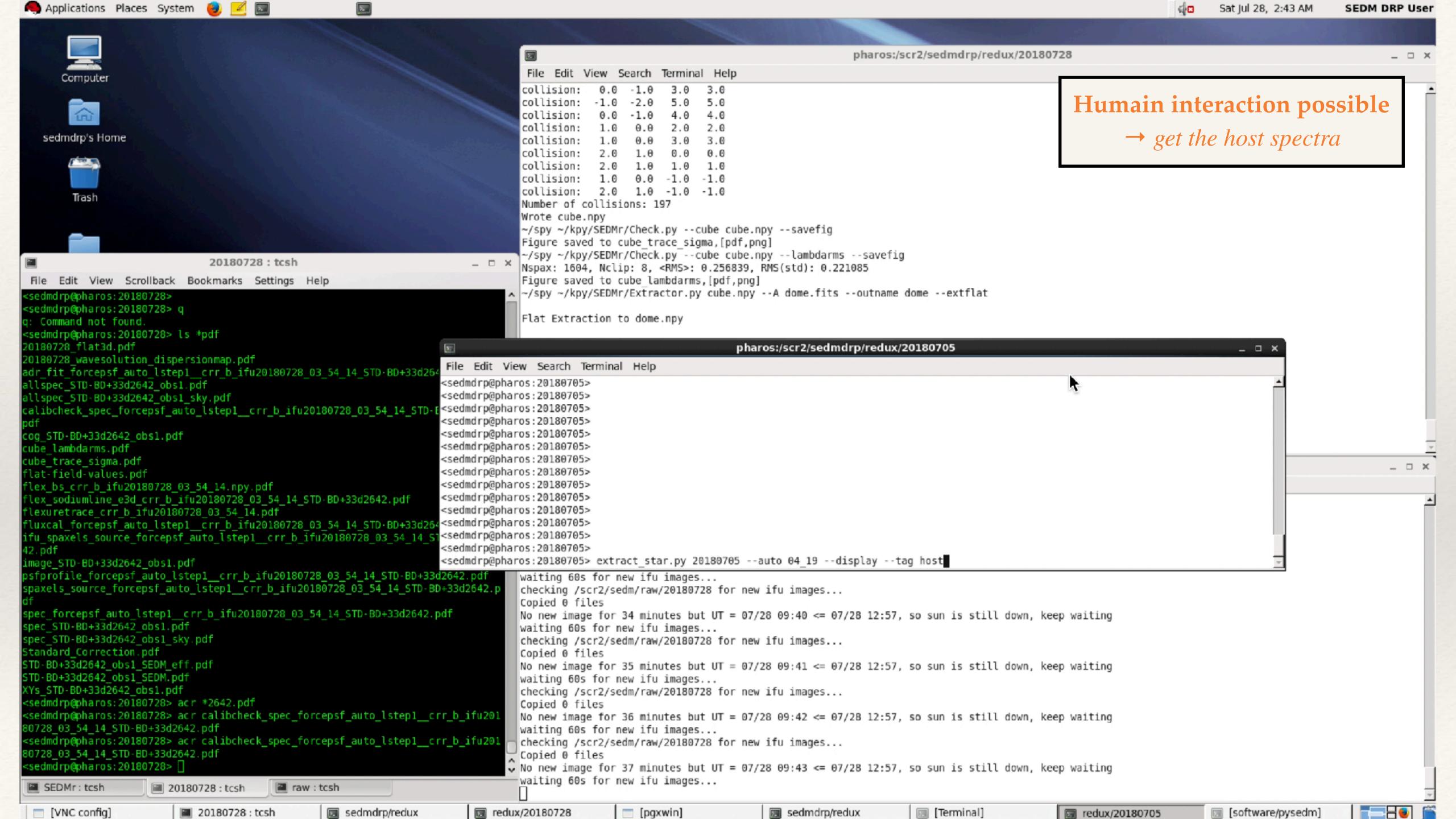
Centroid position (ADR model)

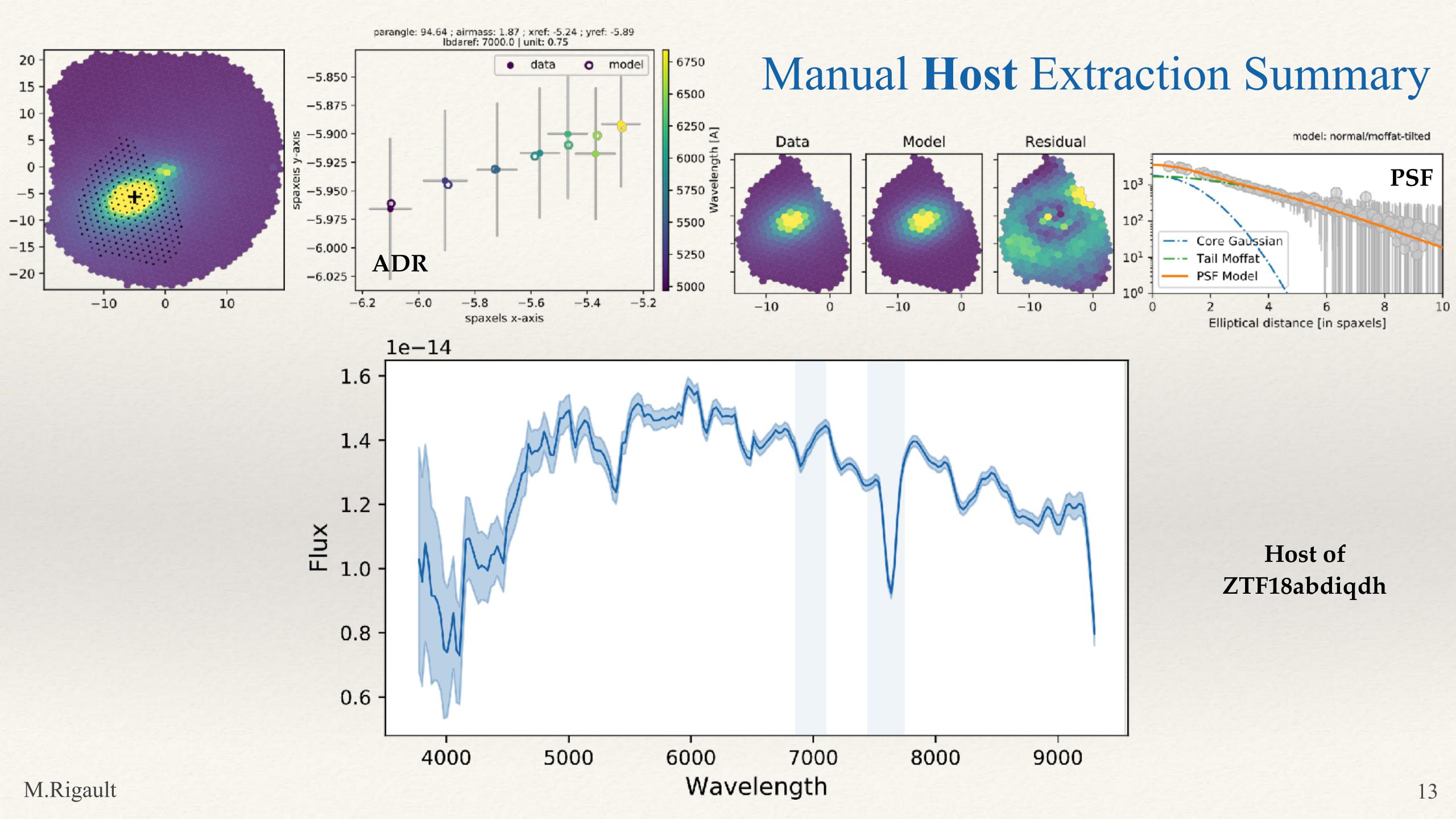
PSF modelling & fitting (based on meta-slices)

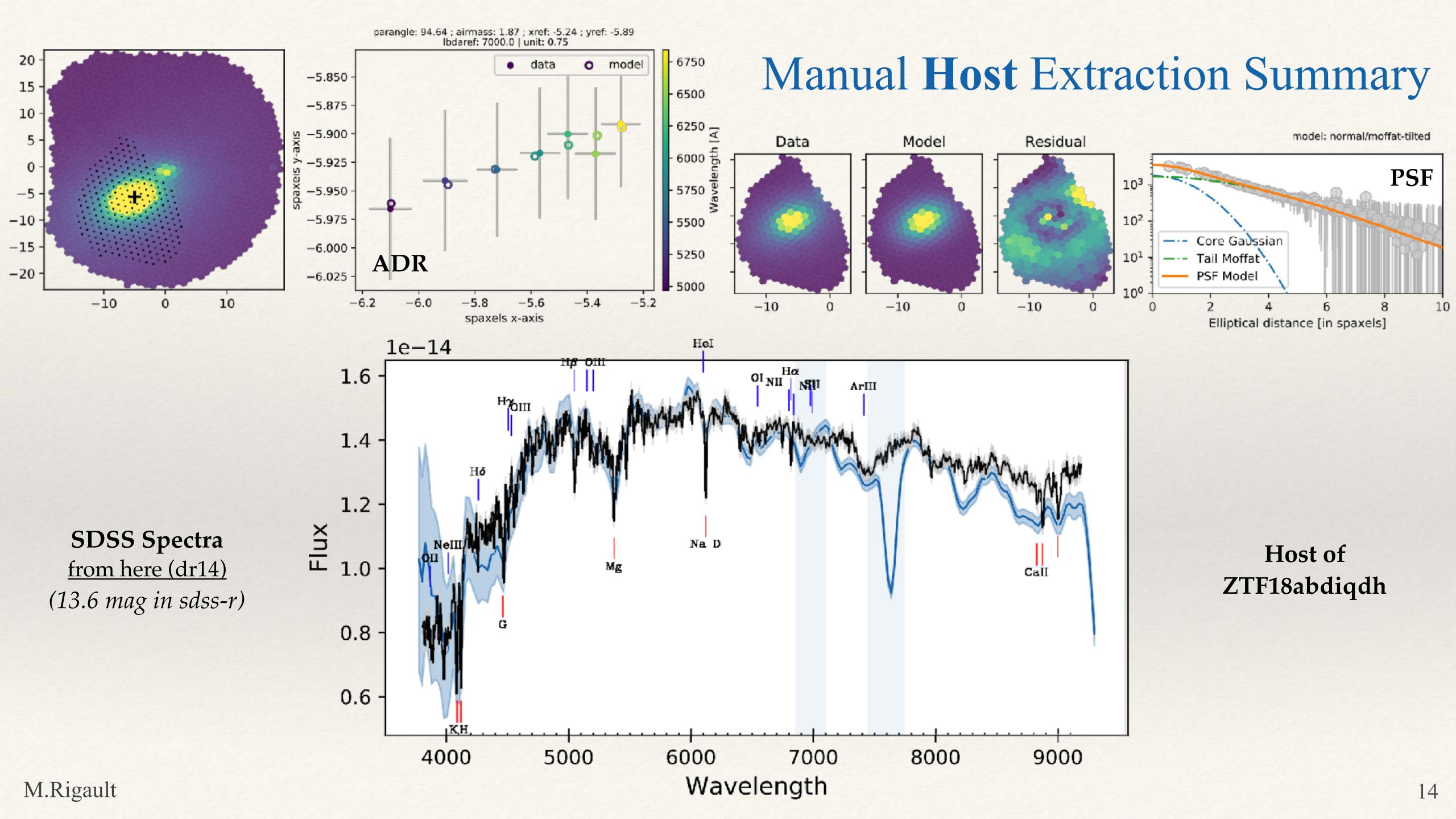
No Humain interaction needed



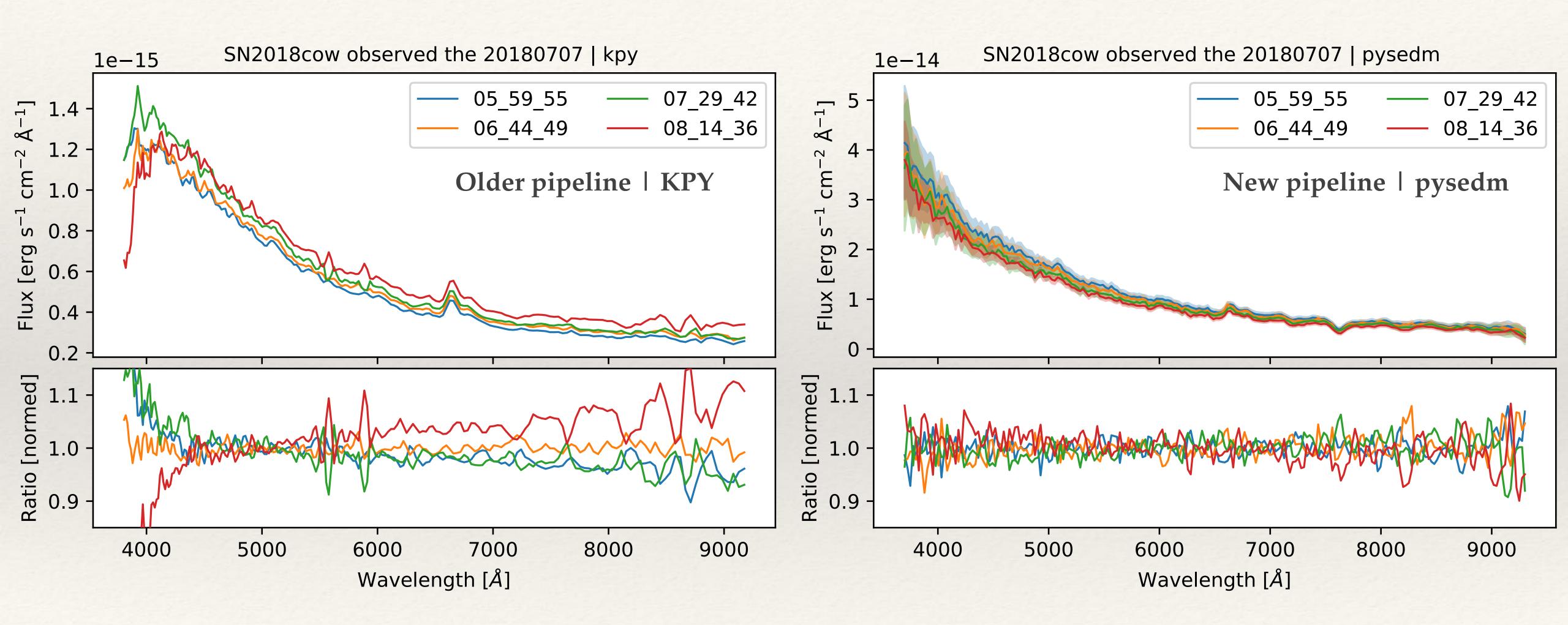




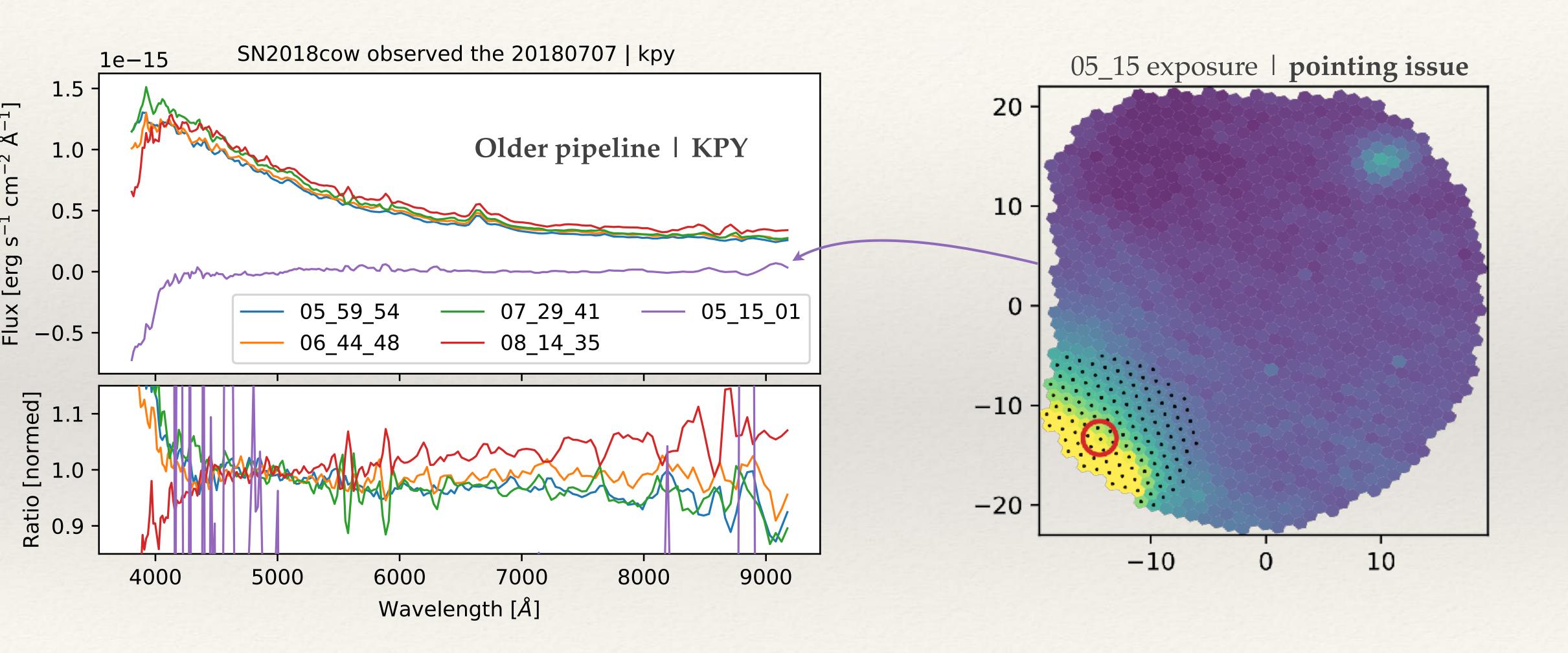




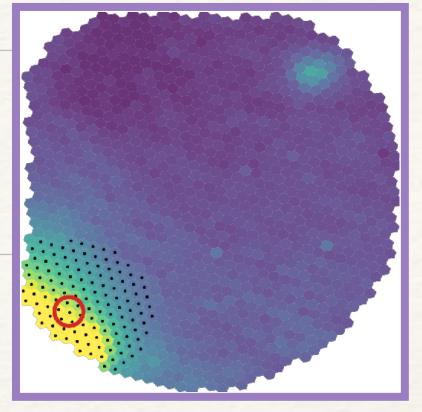
kpy vs. pysedm | Several observations during the night

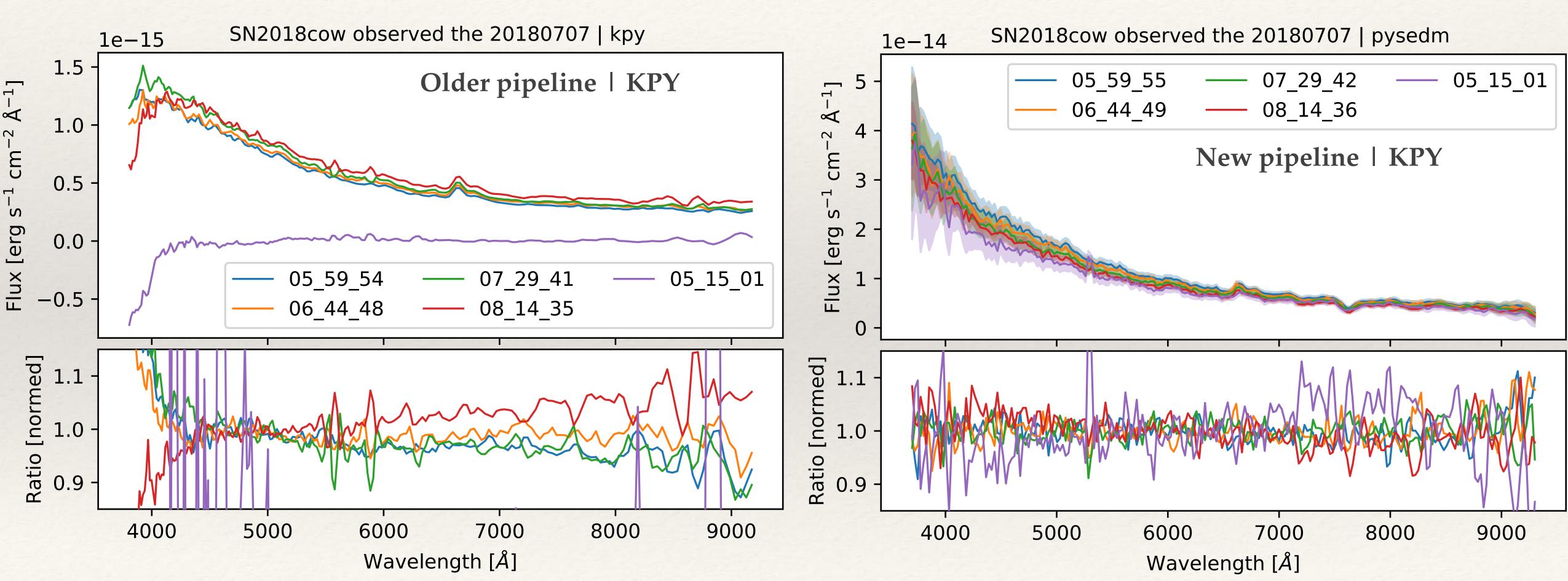


kpy vs. pysedm pointing issue

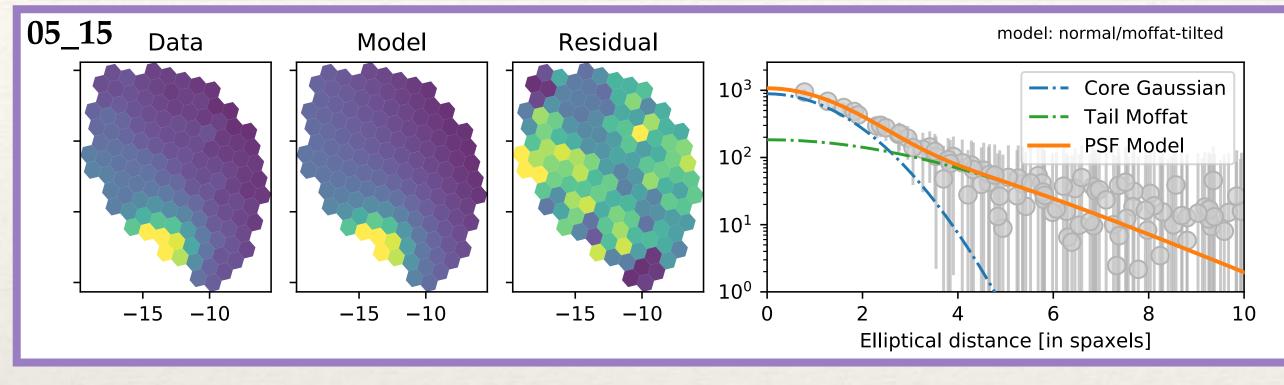


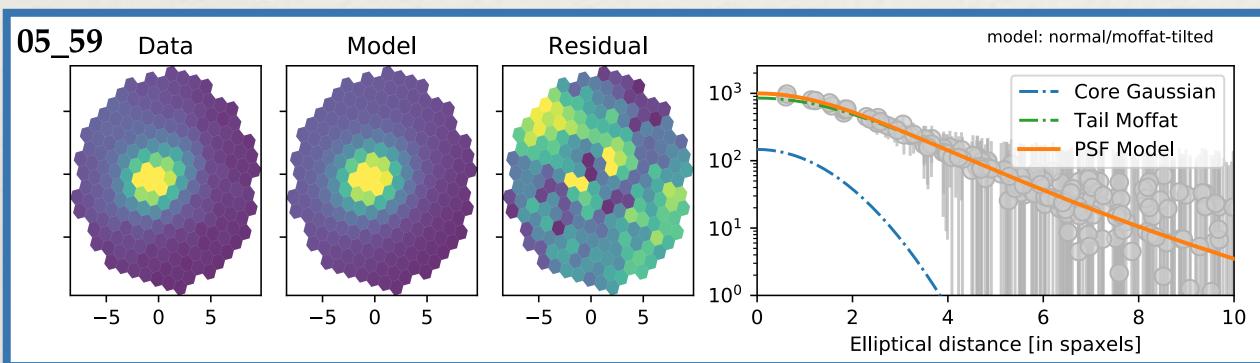
kpy vs. pysedm | pointing issue

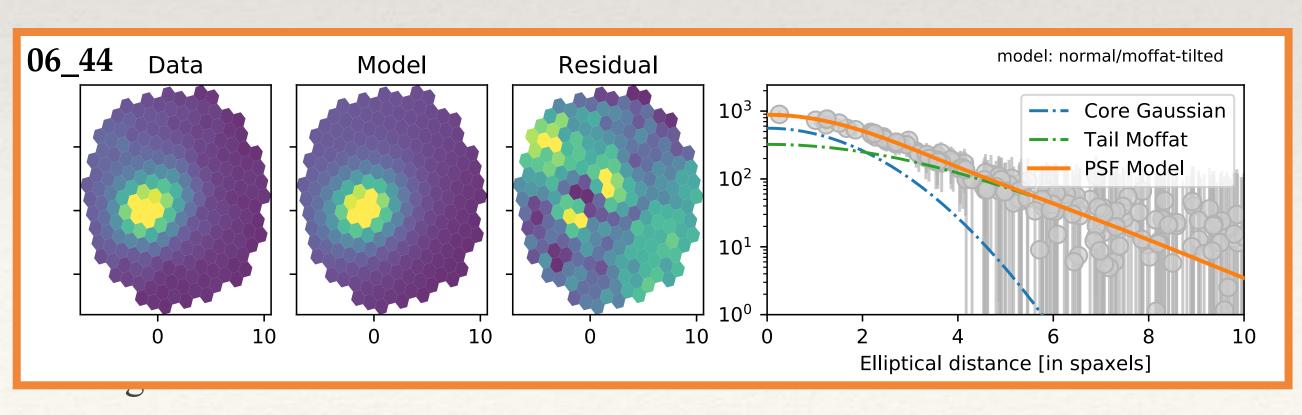


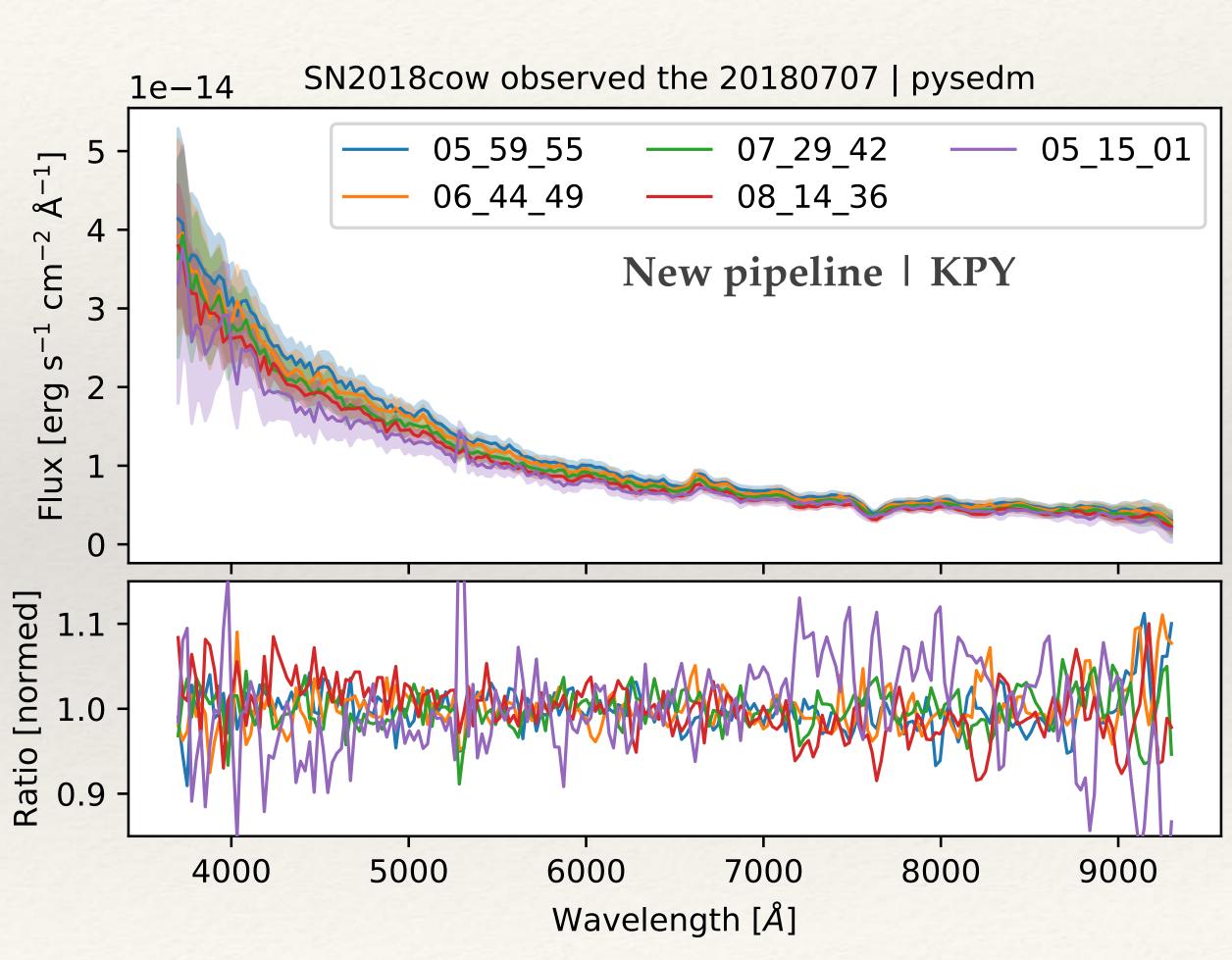


kpy vs. pysedm | pointing issue & PSF fitting



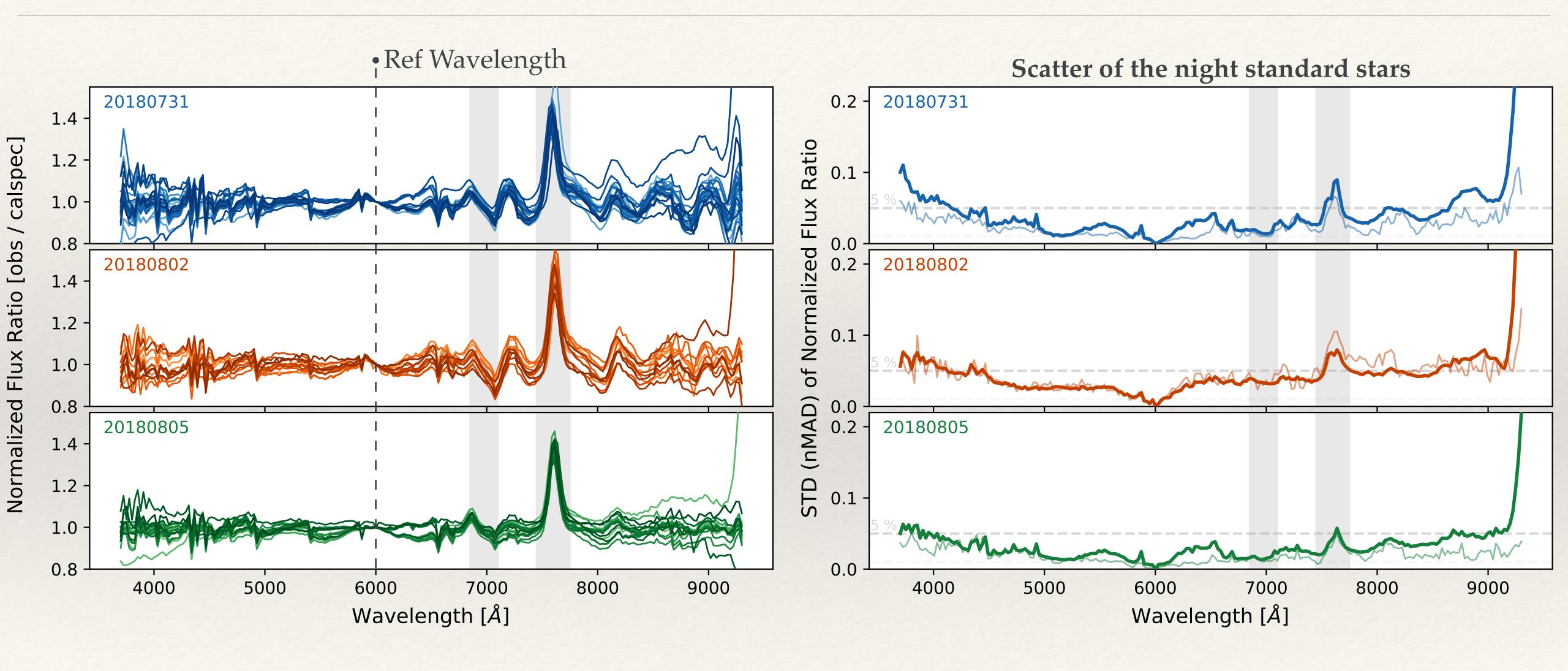




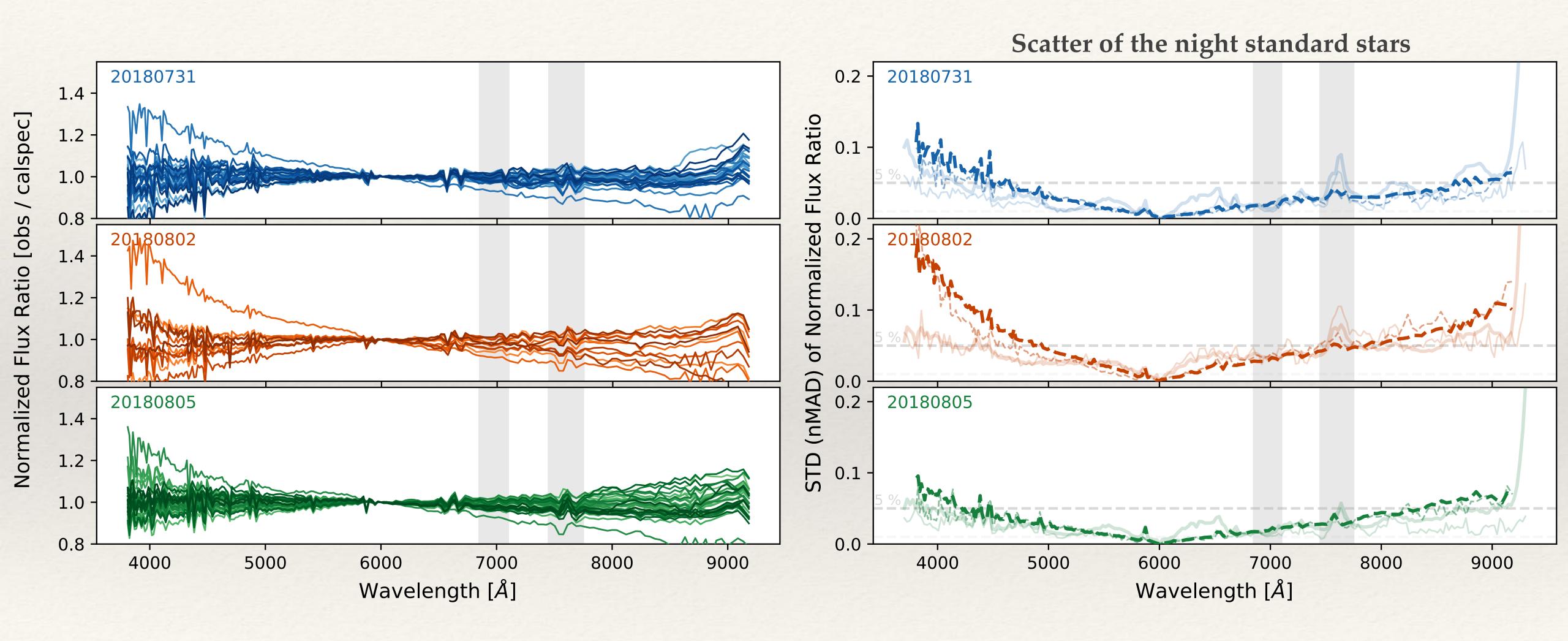


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Standard Star Scatter



Standard Star Scatter | kpy vs. pysedm



Conclusion

pysedm pipeline is now fully automated

Spectra are available ~6 min after end of exposure

Typing possible for >19.5mag in less than 45 min

Software fully compatible with python 2.7 and 3.x

To be finished:

accurate wcs | telluric correction

More development possible (like host subtraction)