



# The Small Solar System Body Science by Using PTF Data

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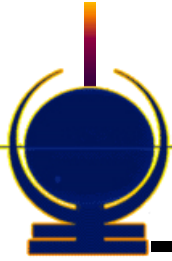
National Central University



Graduate Institute of Astronomy, NCU

國立中央大學天文研究所

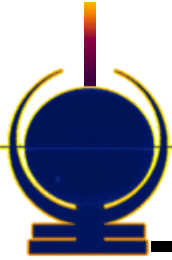




# The SSSB Projects

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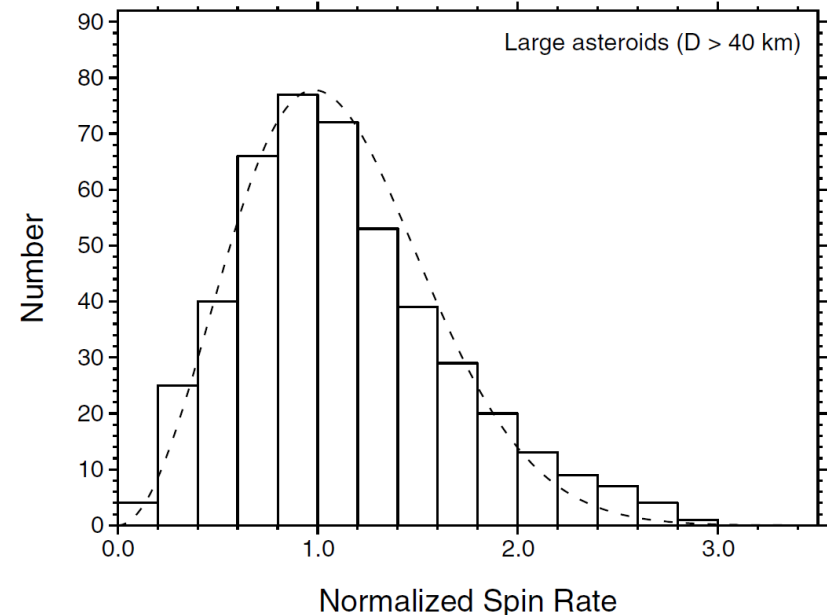
- **The Asteroid Spin Rate Study**
  - Chang, Chan-Kao (Rex)
- **The Phase-curve relation (H-G Analysis)**
  - Cheng, Yu-Chi
- **The Moving Object Detecting Pipeline**
  - Lin, Hsing-Wen (Edward) + CSIE dept.
- **The PTF Asteroid Lightcurve Database**
  - Chang, Chan-Kao (Rex)

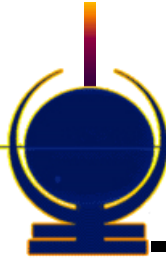


# Q for Asteroid Rotation (1)

## ■ The Spin Rate Distribution

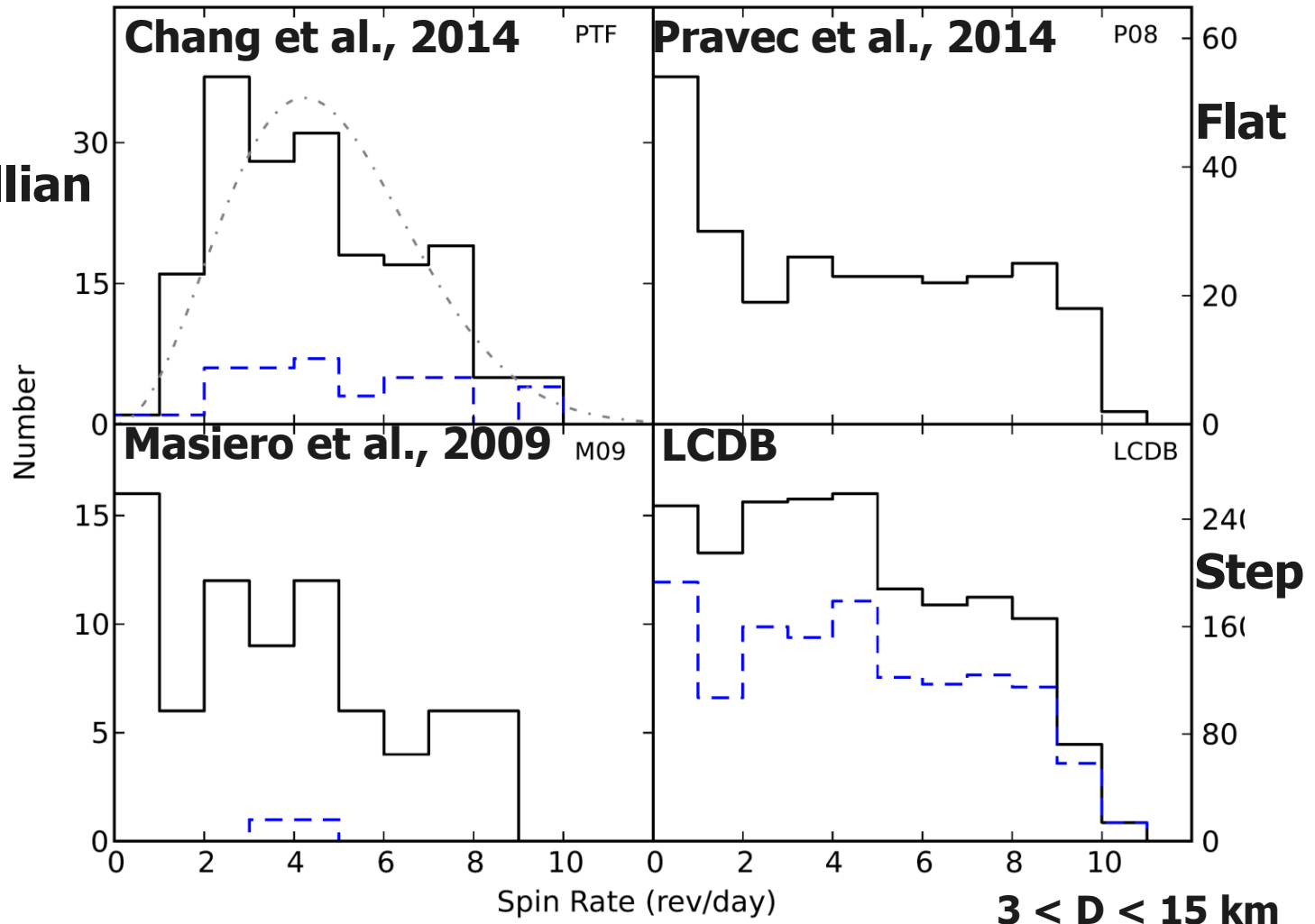
- $D > 40$  km
  - Maxwellian distribution (Pravec et al., 2002)
  - Collisionally evolved system (Salo, 1987)
  
- $D < 10$  km
  - Deviated from Maxwellian ?? (excesses of very slow/fast rotator; Vokrouhlicky, 2002)
  - YORP effect (angular momentum change due to sunlight absorption and re-emission)

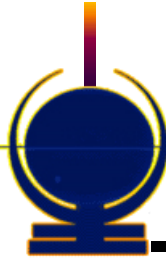




# Q for Asteroid Rotation (1)

de-Maxwellian





# Q for Asteroid Rotation (2)

## The Spin Rate Limit

### ■ Rubble pile & monolithic

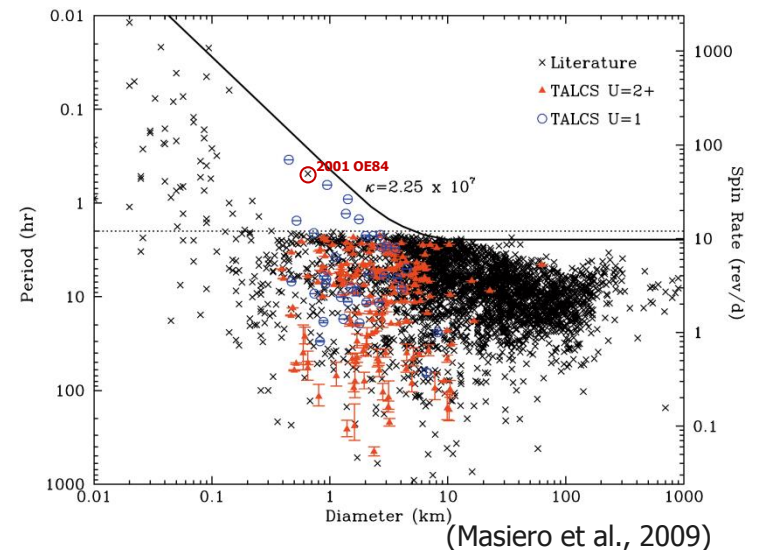
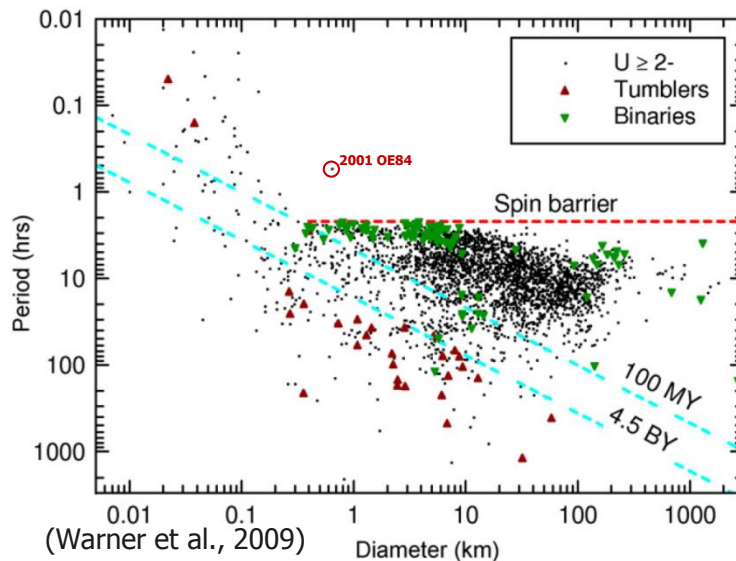
(Harris, 1996)

- spin-barrier at  $\sim 2$  hours for  $D > 150\text{m}$  (Gravitationally bounded)
- super-fast-rotators ( $P < 2\text{hr}$ ) for  $D < 150\text{m}$  (monolith)

### ■ Size-dependent strength

(Holsapple, 2007)

- spin-barrier at  $\sim 2$  hours for  $D > 10\text{km}$  (Gravitationally bounded)
- **super-fast-rotators for  $D < 5\text{km}$**  (tensile and cohesive; a transition from small to larger objects)





# Observations

- **12 fields**
- **4 adjacent nights**
- **On the ecliptic plane**

## Jan, 2014

	c.	Jan 6	Jan 7	Jan 8	Jan 9
	)	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt
3559	117.00	19.12	9.6, 28	9.8, 30	5.1, 16
3560	120.60	19.12	9.7, 29	9.8, 30	5.1, 16
3561	124.20	19.12	9.3, 28	9.4, 29	5.5, 17
3562	127.80	19.12	9.3, 28	9.1, 28	5.5, 17
3563	131.40	19.12	9.2, 28	8.8, 27	5.1, 16
3564	135.00	19.12	8.9, 27	8.7, 27	5.1, 16
3565	138.60	19.12	8.5, 26	8.4, 26	5.1, 16
3658	115.71	21.38	9.7, 29	9.8, 30	5.1, 16
3659	119.39	21.38	9.7, 28	9.8, 30	5.1, 16
3660	123.06	21.38	9.6, 29	9.4, 29	5.1, 16
3661	126.73	21.38	9.3, 28	9.5, 29	5.1, 16
3662	130.41	21.38	9.3, 28	9.1, 28	5.1, 16

## Feb, 2013

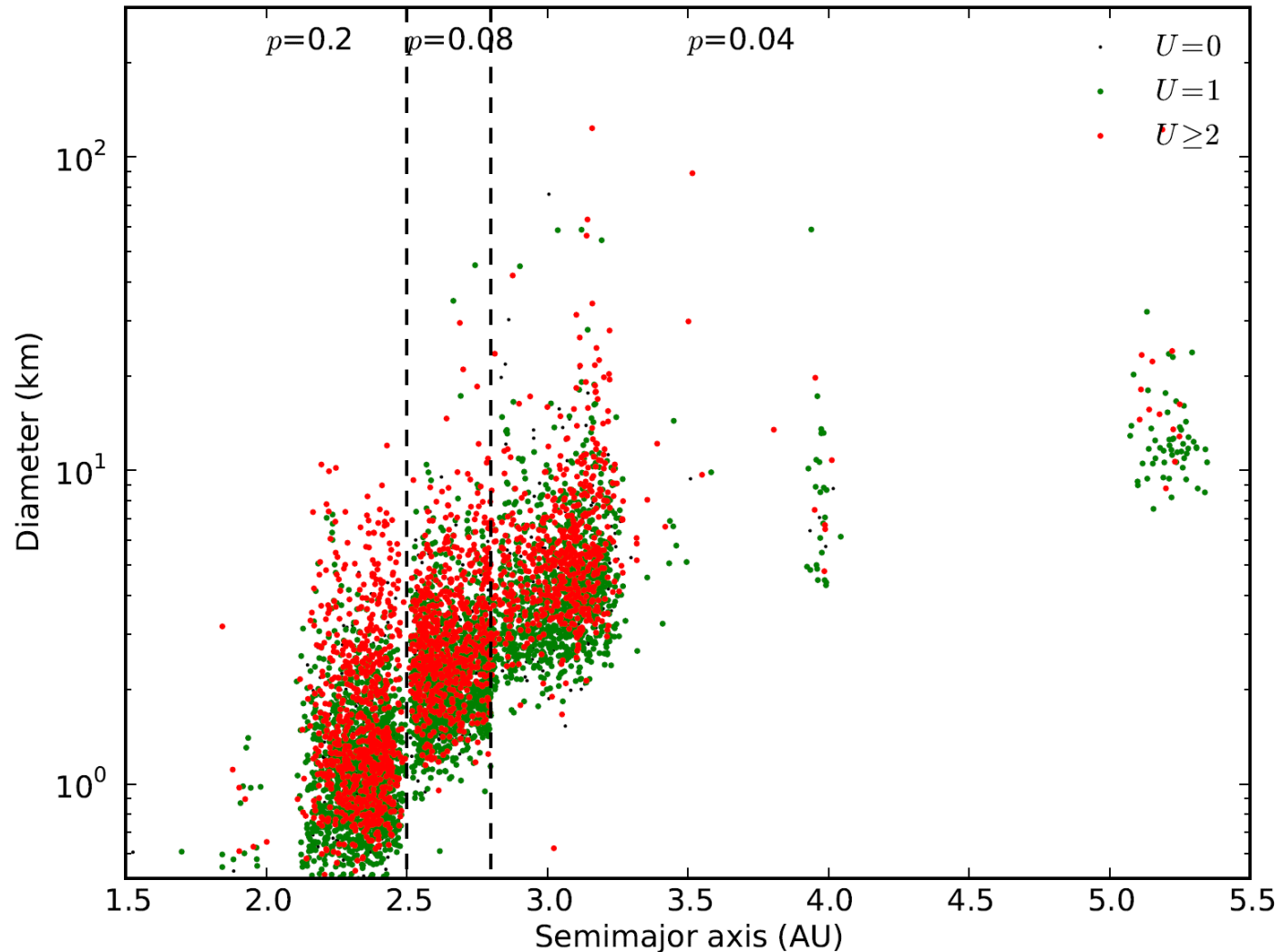
			Feb 15	Feb 16	Feb 17	Feb 18
	\ /	\ /	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt
3654	101.02	21.38	3.4, 10	5.1, 16	6.8, 19	6.4, 19
3655	104.69	21.38	3.4, 10	5.3, 15	6.8, 19	6.4, 19
3656	108.37	21.38	3.4, 10	5.3, 15	7.2, 20	6.7, 20
3657	112.04	21.38	3.4, 10	5.4, 15	7.2, 20	7.1, 21
3658	115.71	21.38	3.4, 10	5.5, 16	7.6, 21	7.4, 22
3749	90.93	23.62	3.4, 10	5.3, 14	6.4, 18	6.0, 18
3750	94.64	23.62	3.4, 10	5.1, 15	6.4, 18	6.0, 18
3751	98.35	23.62	3.4, 10	5.1, 15	6.8, 19	6.4, 19
3752	102.06	23.62	3.4, 10	5.0, 14	6.8, 19	6.7, 20
3753	105.77	23.62	3.4, 10	5.1, 15	7.2, 20	6.7, 20
3754	109.48	23.62	3.4, 10	5.1, 15	7.6, 21	7.1, 21
3755	113.20	23.62	3.4, 10	5.4, 15	7.6, 21	7.4, 22

## Feb, 2014

			Feb 20	Feb 21	Feb 22	Feb 23
			$\Delta t$ , Nexpt	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt	$\Delta t$ , Nexpt
3158	143.65	10.12	7.9, 19	7.6, 22	7.9, 22	7.2, 20
3159	147.12	10.12	8.0, 17	7.3, 20	7.9, 21	7.7, 22
3160	150.58	10.12	7.6, 18	7.7, 21	7.7, 22	7.7, 22
3161	154.04	10.12	8.0, 20	8.0, 22	7.9, 23	7.9, 23
3162	157.50	10.12	8.3, 20	7.9, 22	8.0, 24	7.9, 23
3163	160.96	10.12	8.0, 19	7.8, 22	8.0, 24	8.0, 24
3261	141.55	12.38	7.7, 20	7.6, 21	8.2, 23	7.1, 20
3262	145.05	12.38	8.1, 20	7.9, 23	8.2, 23	7.5, 20
3263	148.54	12.38	8.2, 20	7.6, 22	8.1, 23	7.7, 21
3264	152.04	12.38	7.9, 20	8.0, 23	8.2, 24	8.2, 24
3265	155.53	12.38	8.7, 19	8.0, 22	8.1, 24	8.2, 23
3266	159.03	12.38	8.0, 21	8.2, 23	8.1, 25	8.2, 24

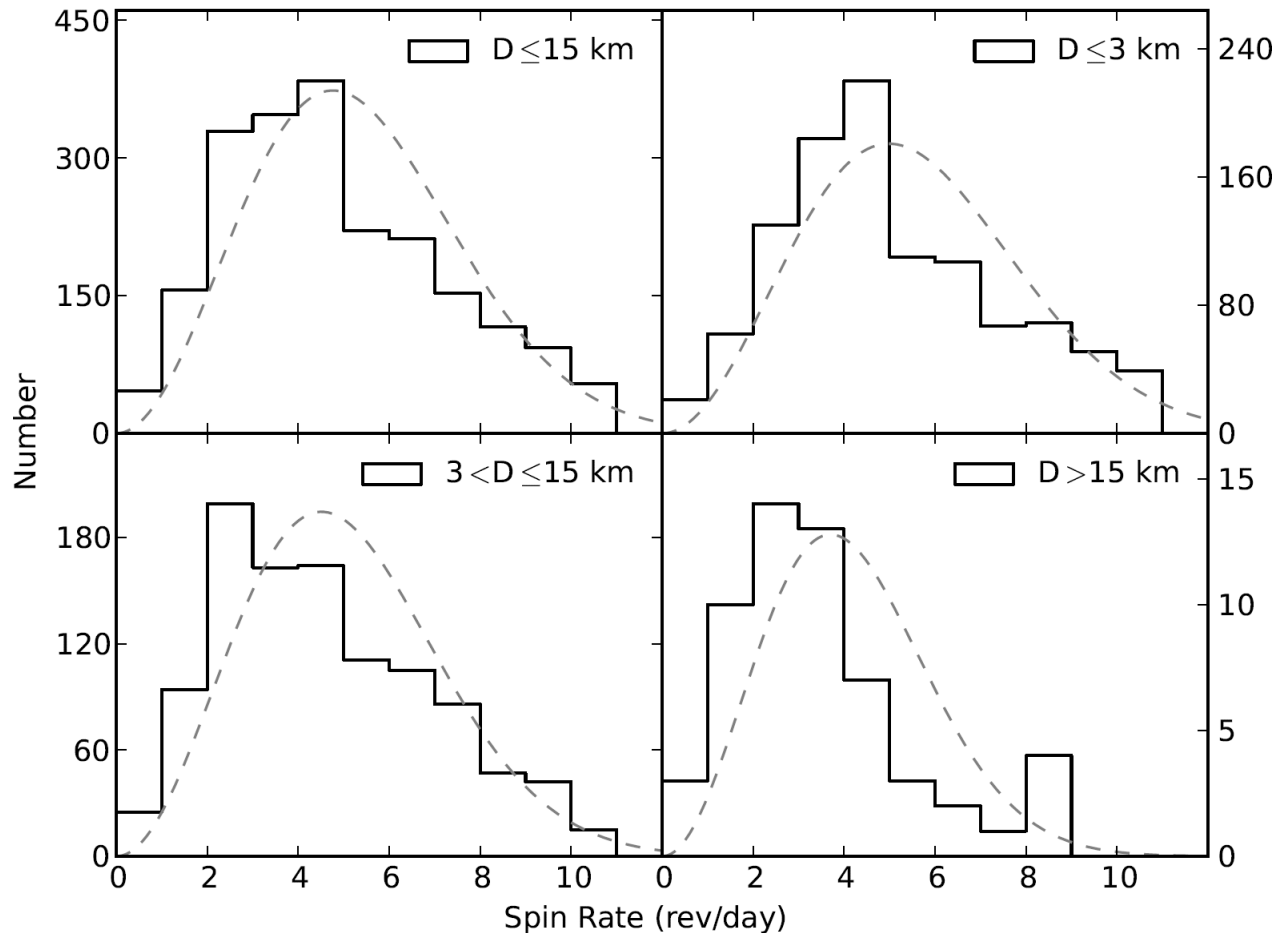


# $\sim 2,220$ Good Rotation Periods





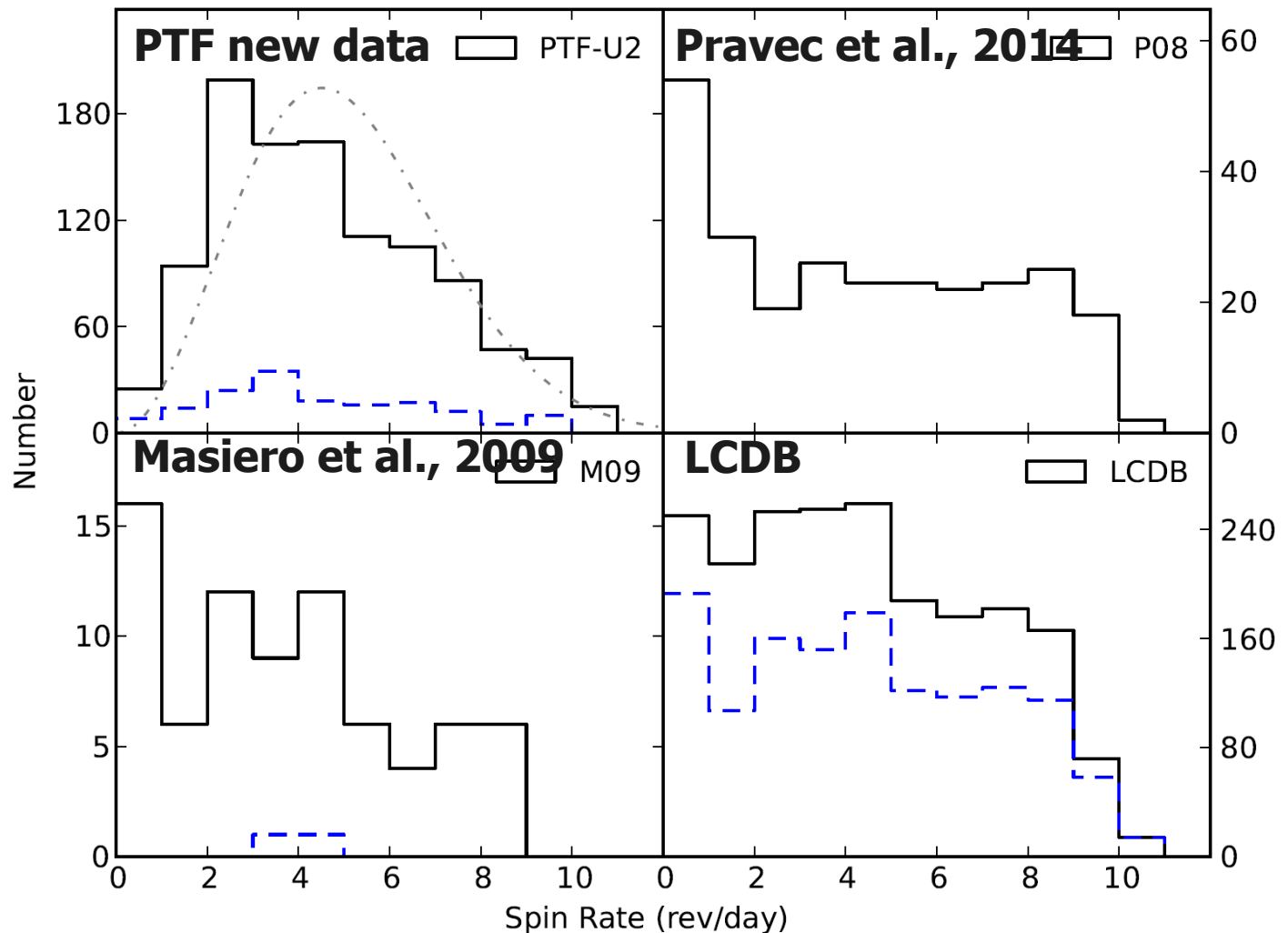
# The PTF Spin Rate Distribution vs Diameters

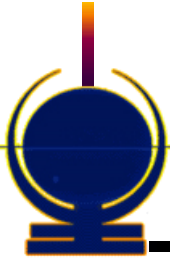


**Deviated-Maxwellian at small asteroids.**

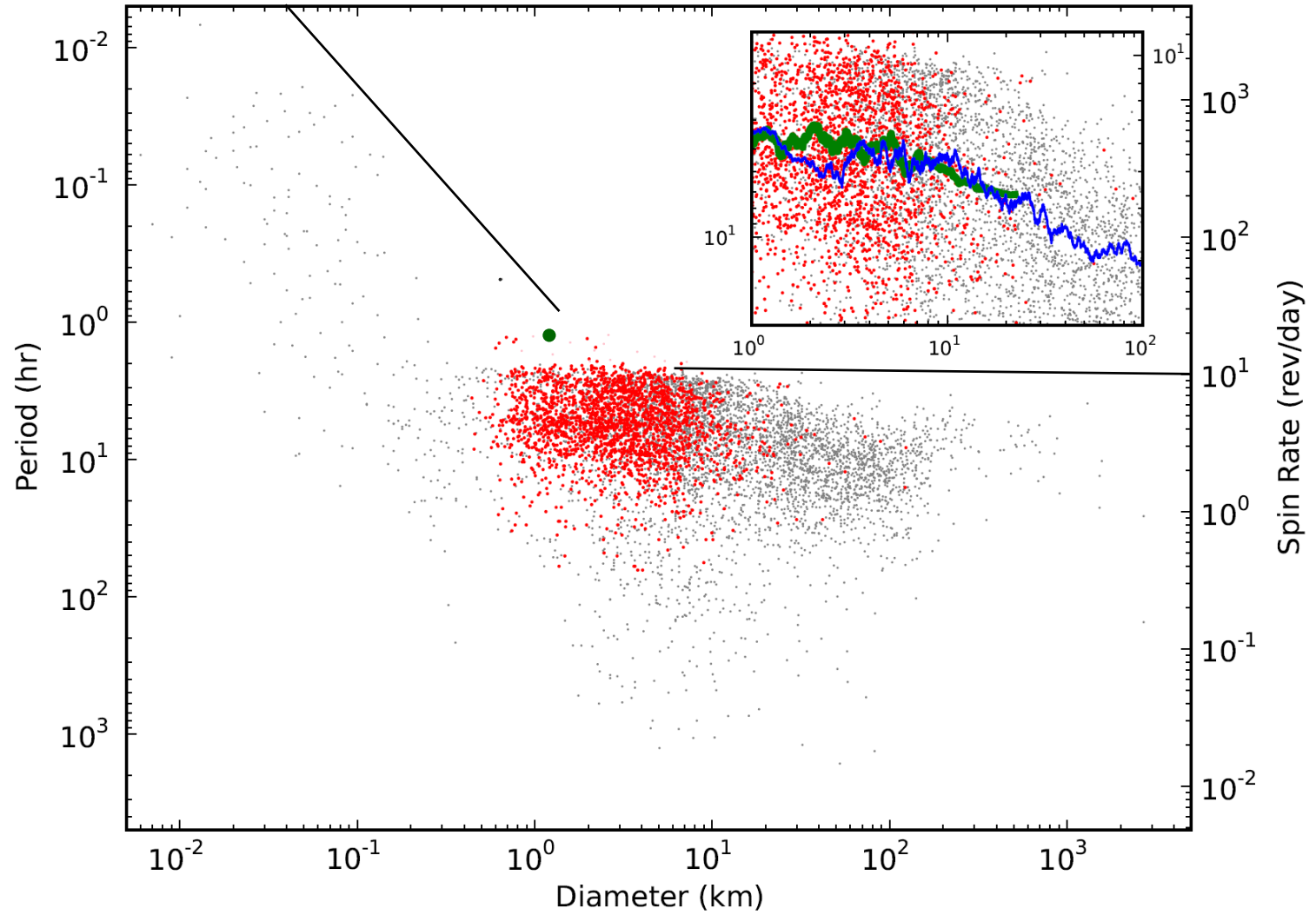


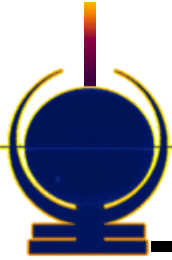
# The PTF Spin Rate Distribution comparing w/ others



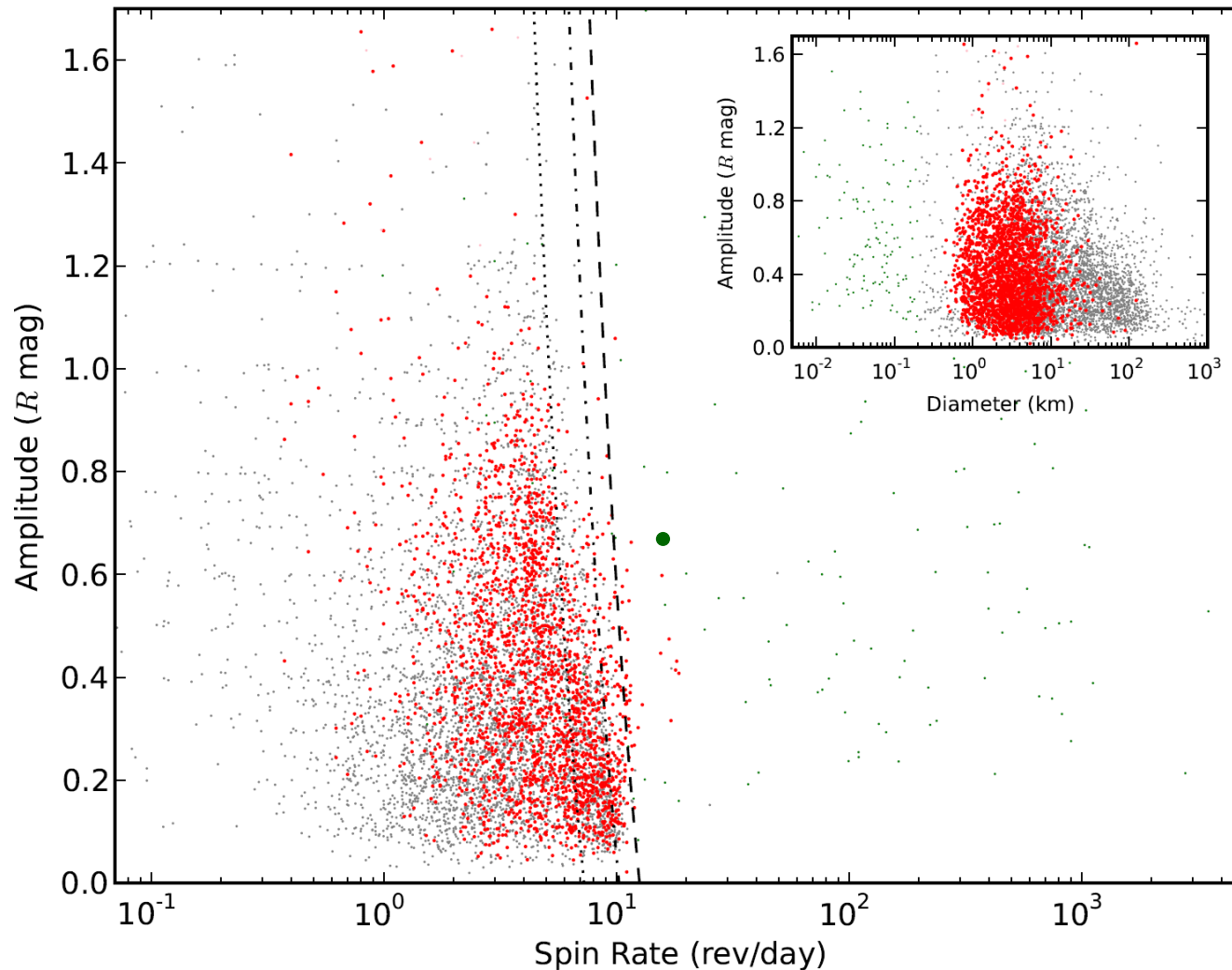


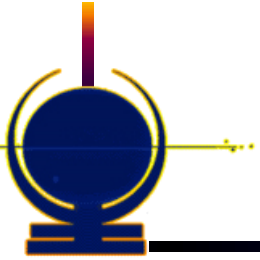
# The Spin Rate Limit – PTF (1)





# The Spin Rate Limit – PTF(2)

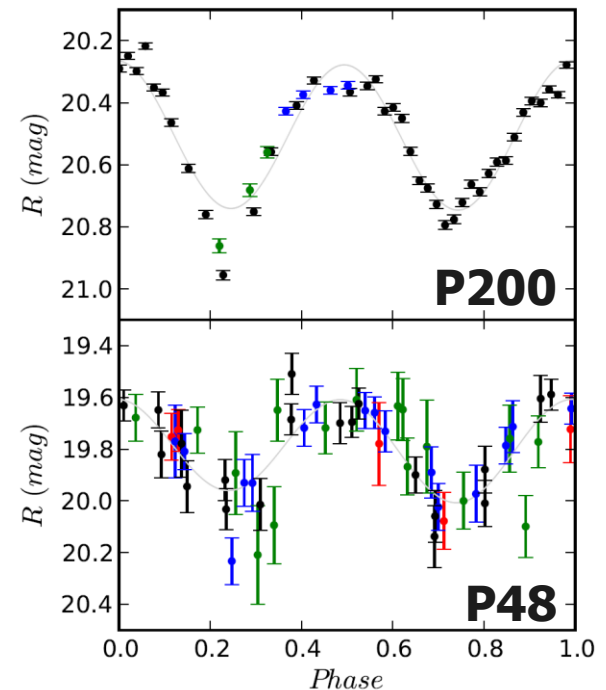
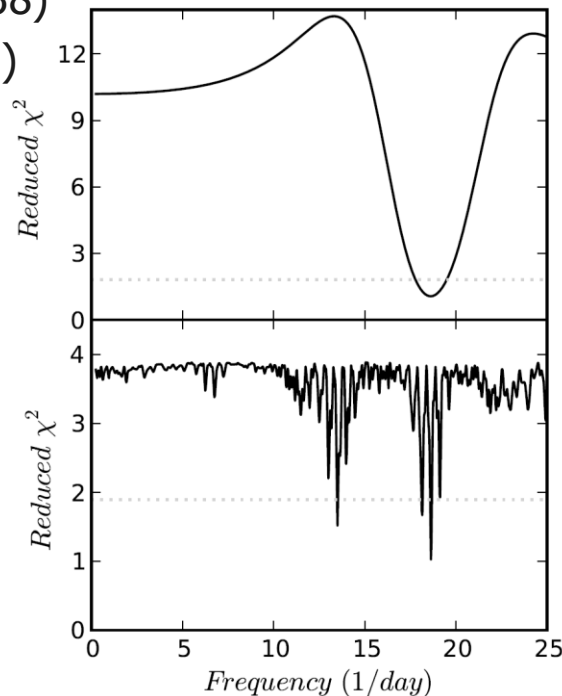




# The sub-km Super-Fast-Rotator

## The PTF super-fast-rotator

- A S-type main belt asteroid
- $D \sim 0.8$  km
- $P = 1.29$  hours ( $f = 18.6$  rev/day)
- $A \sim 0.7$  R mag ( $b/a \sim 1.38$ )
- Density  $\sim 6$  g/cm<sup>3</sup> (metal)
- A huge monolithic object  
(2x Taipei 101)  
(4x TF running track)



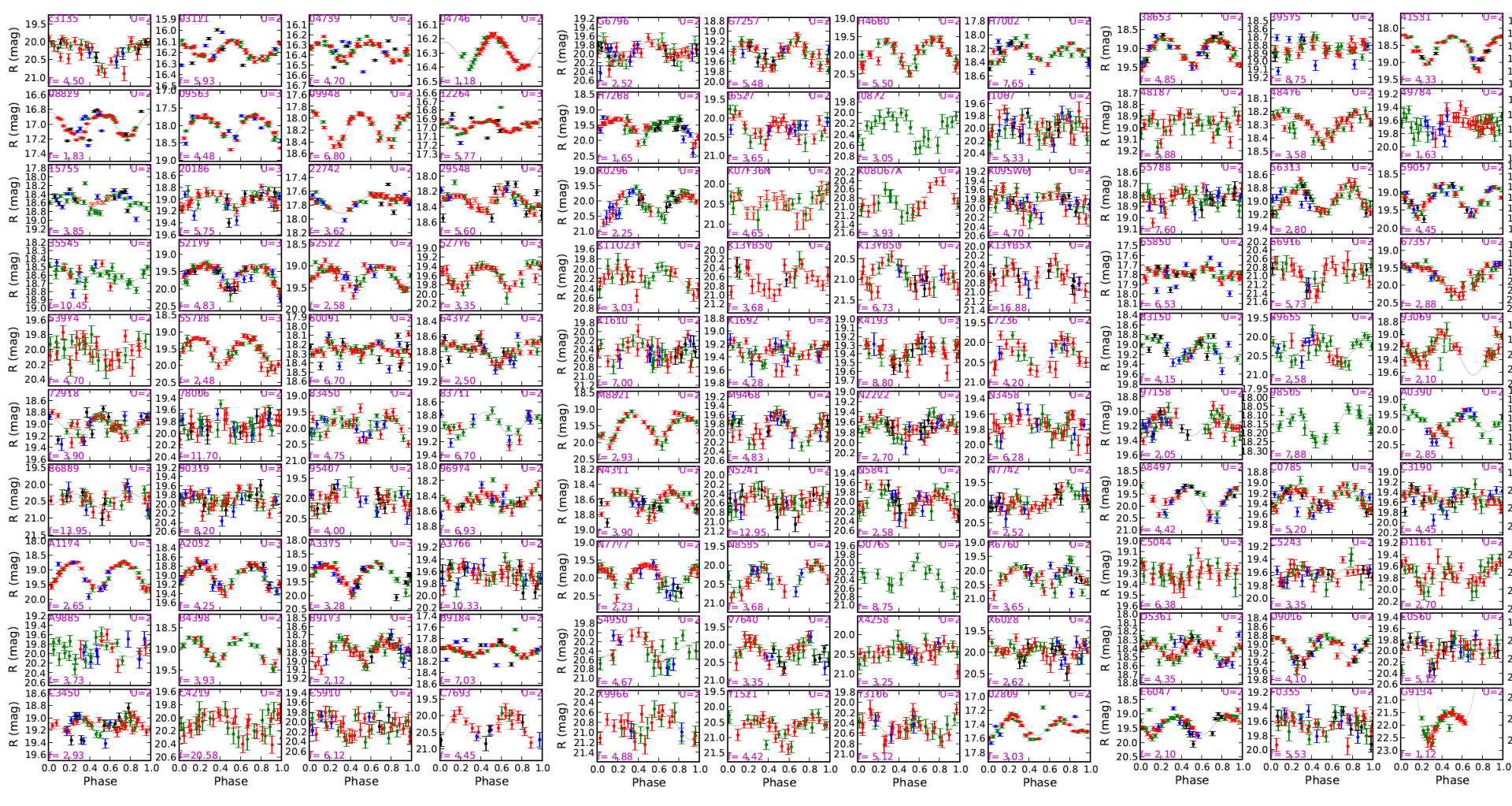


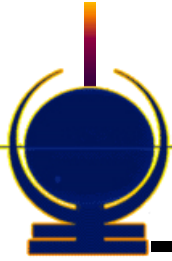
# The sub-km Super-Fast-Rotator

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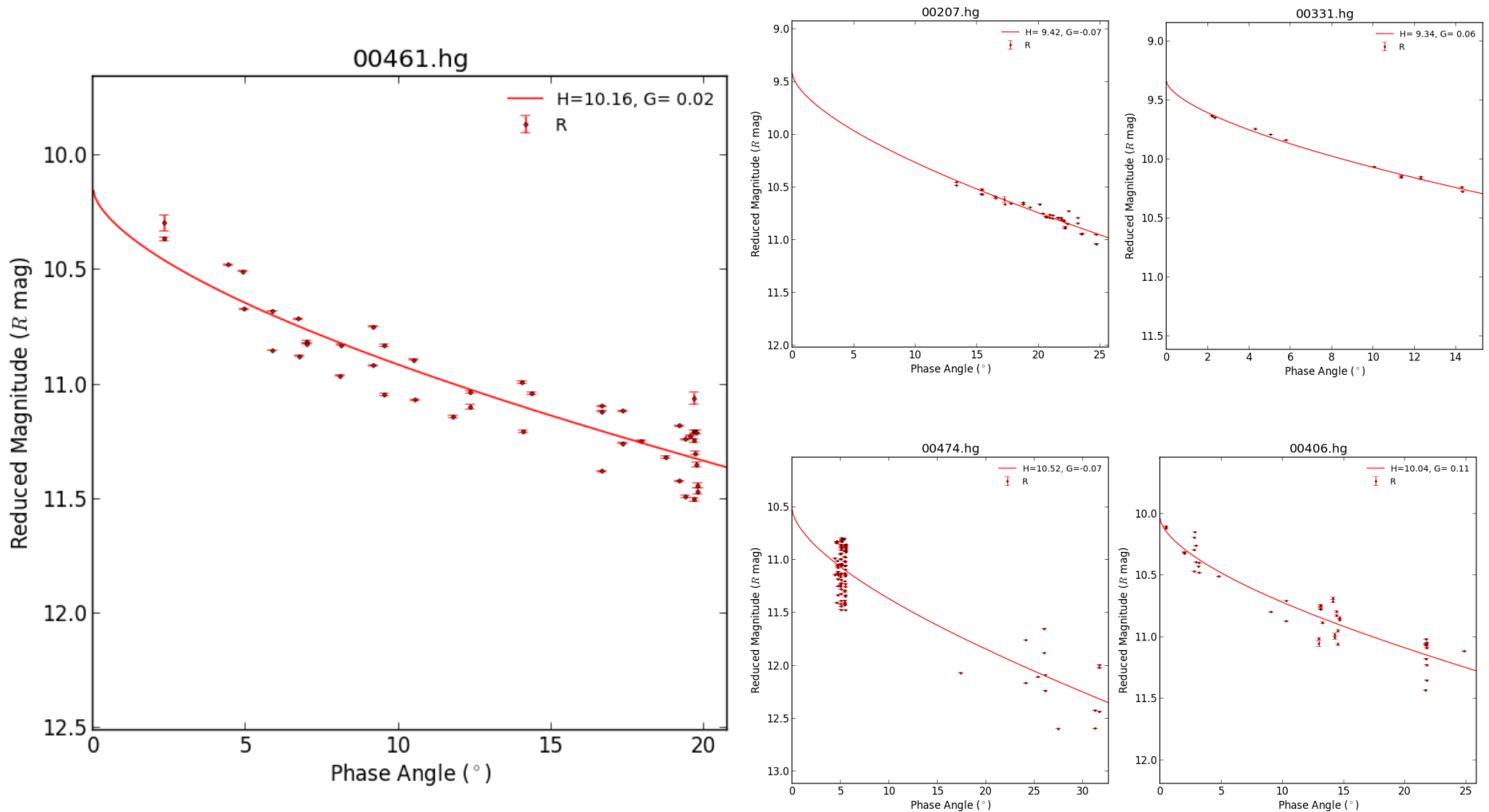


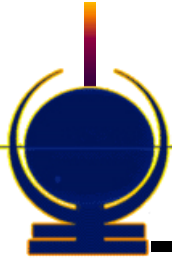






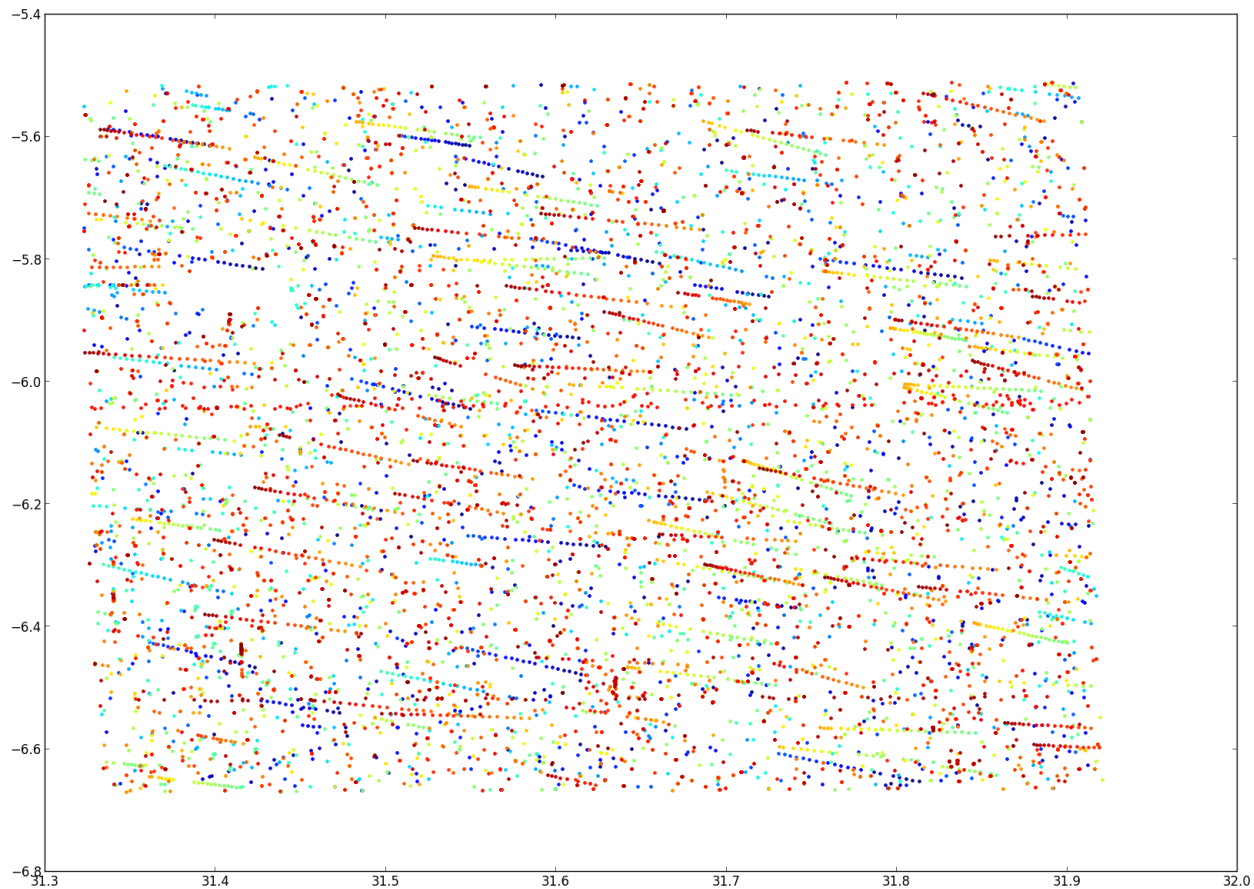
# The H-G Relation





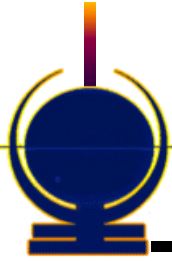
# The Moving Object Pipeline

Using the “Hough Transfer” to link detections belonging to the same “line”



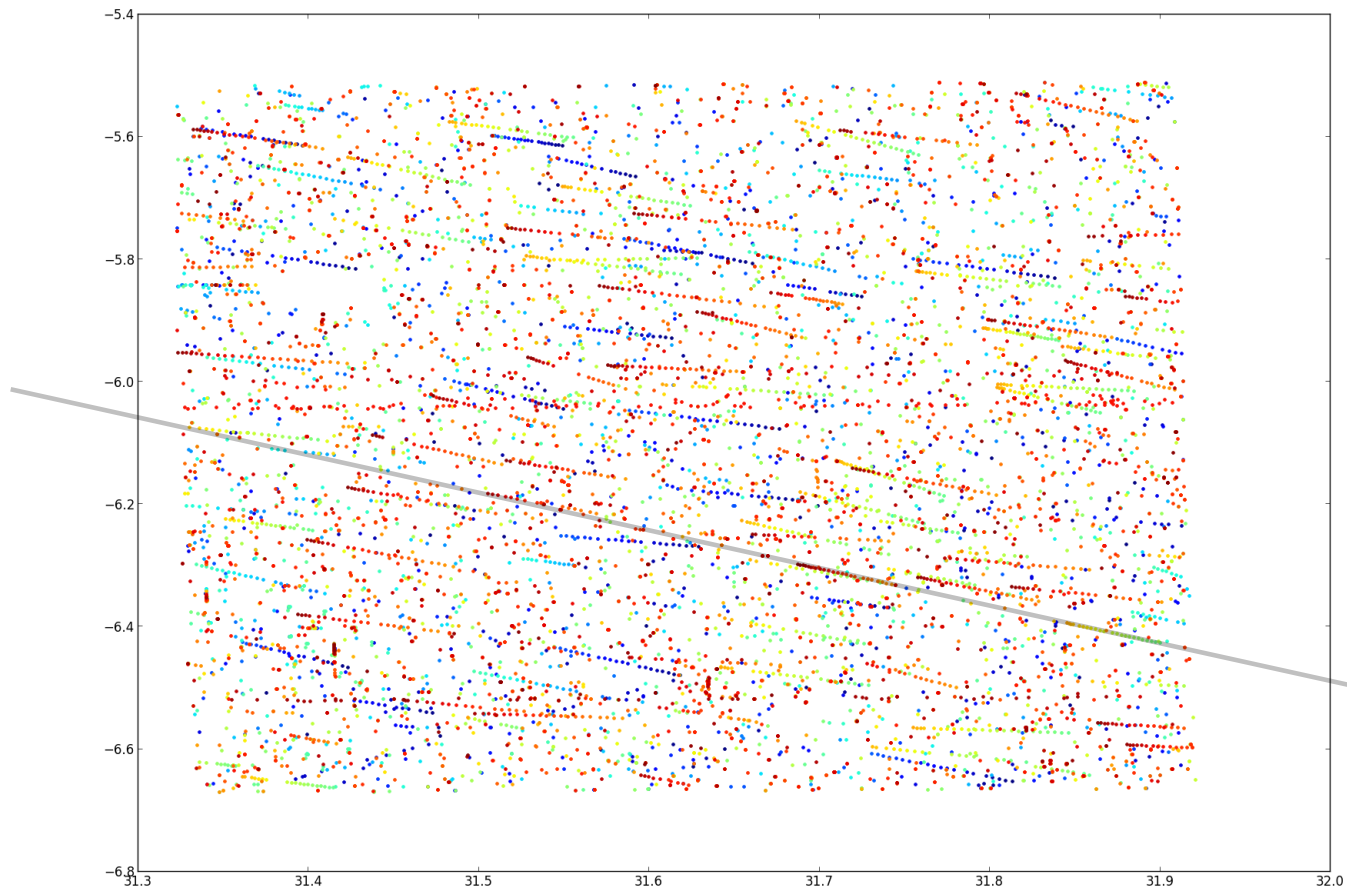
Field 3266, CCD 0, Feb 20-21, 2014



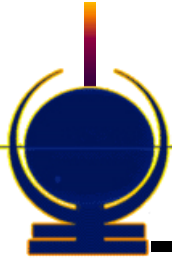


# The Moving Object Pipeline

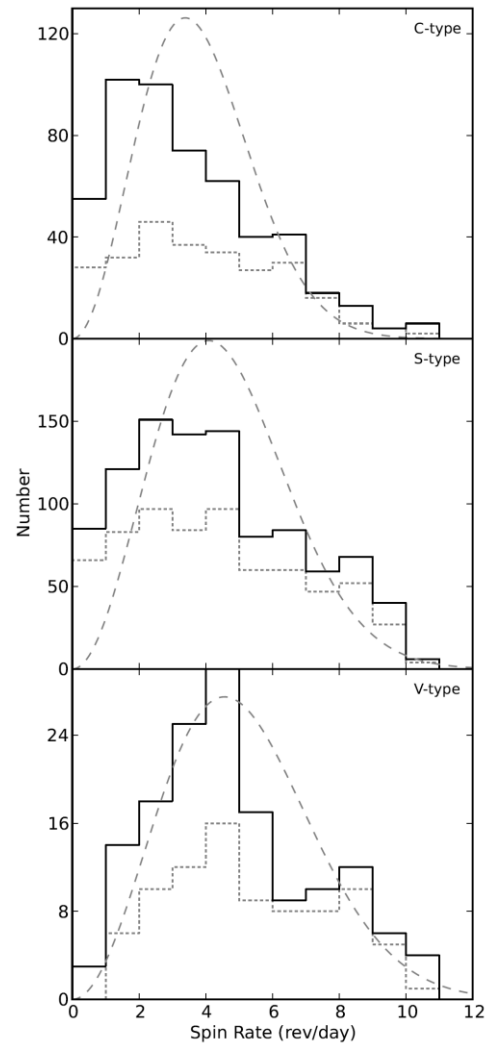
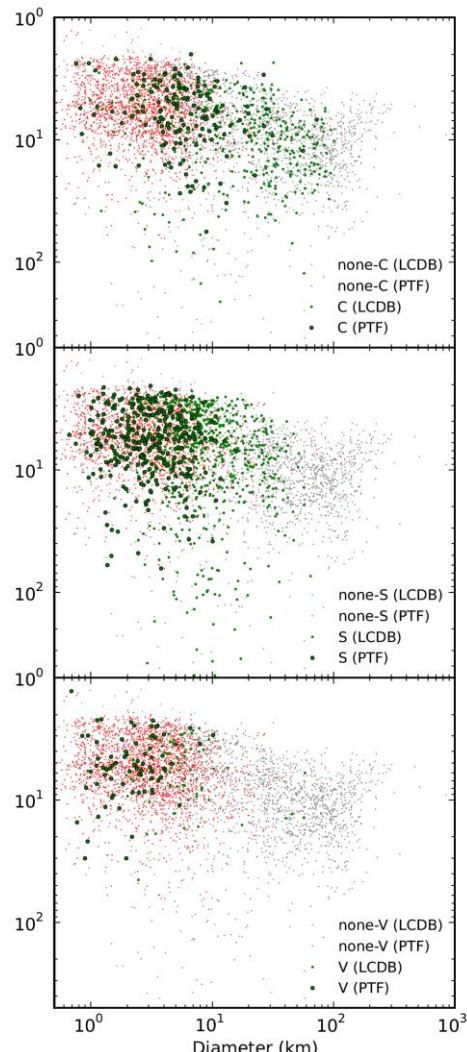
Using the “Hough Transfer” to link detections belonging to the same “line”



Field 3266, CCD 0, Feb 20-21, 2014



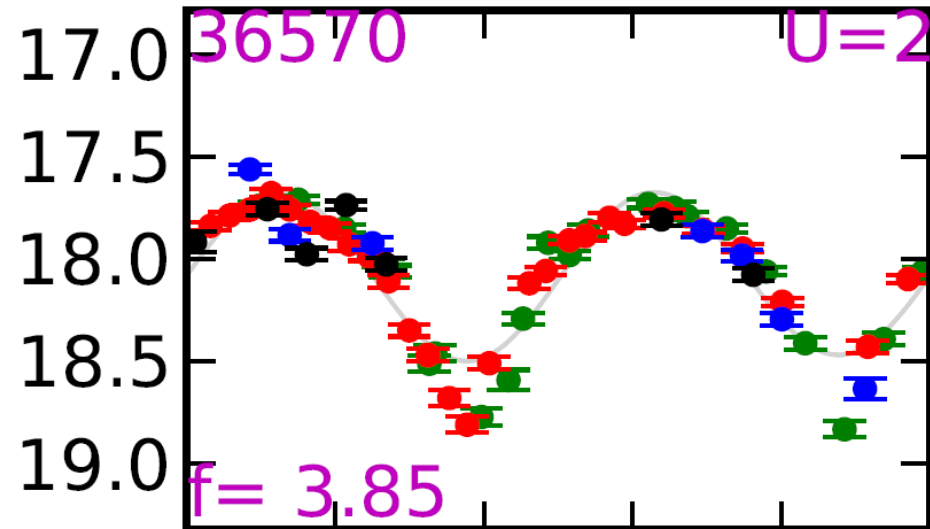
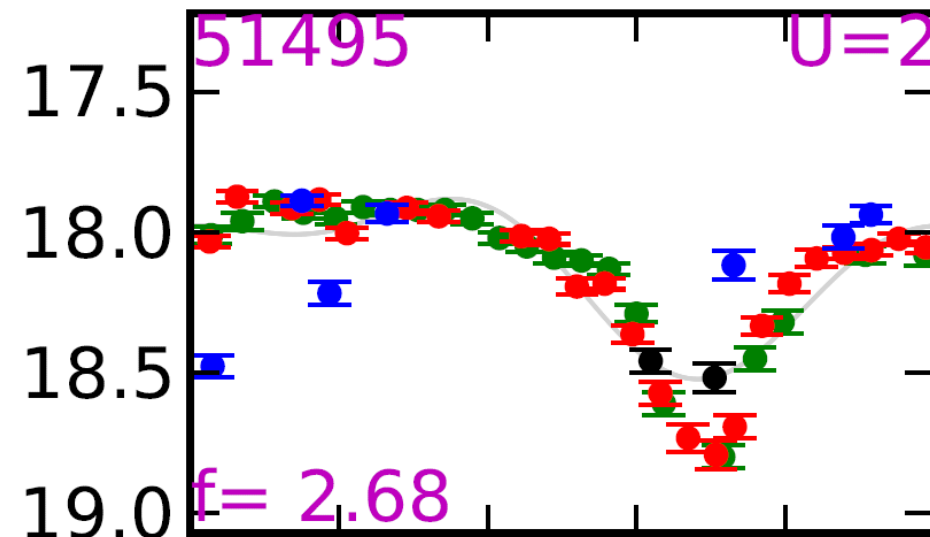
# Spin Rate vs Taxonomy

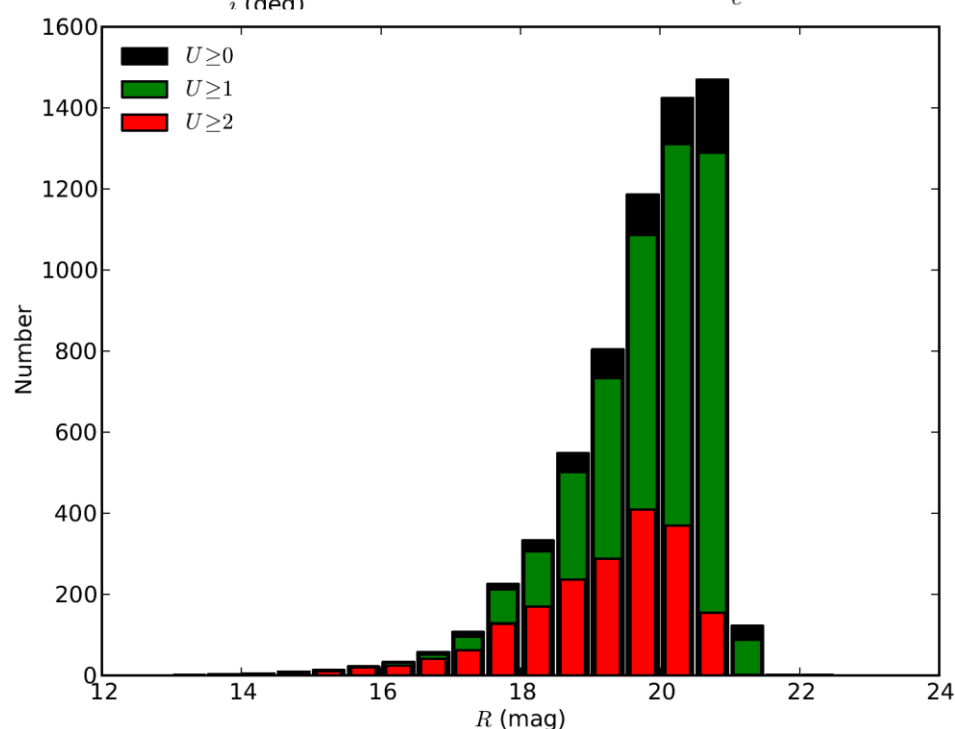
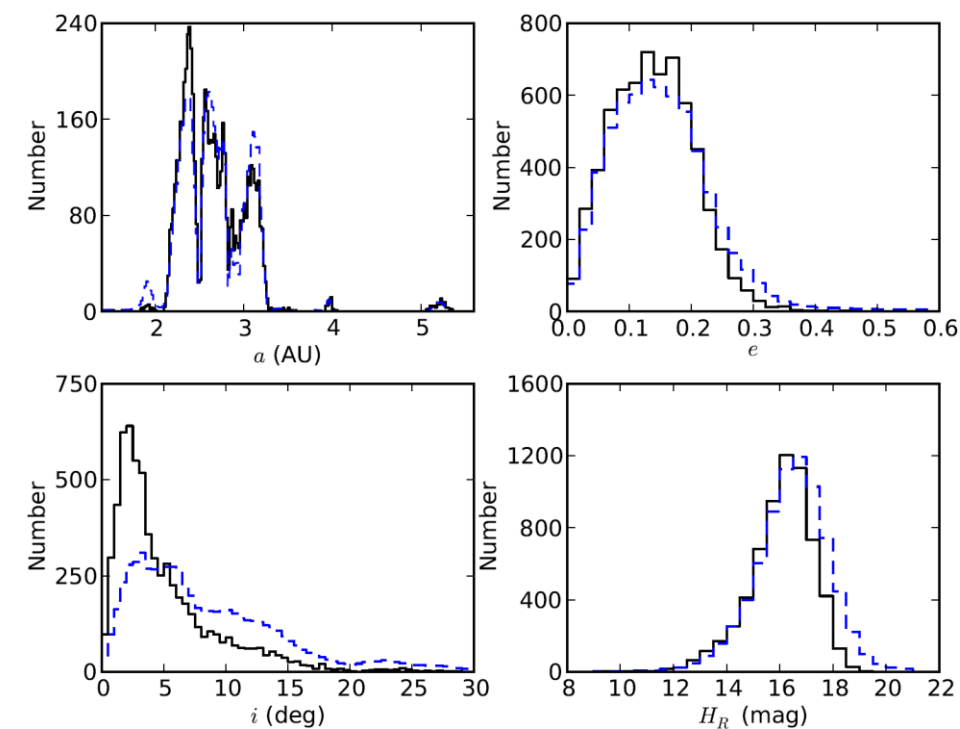
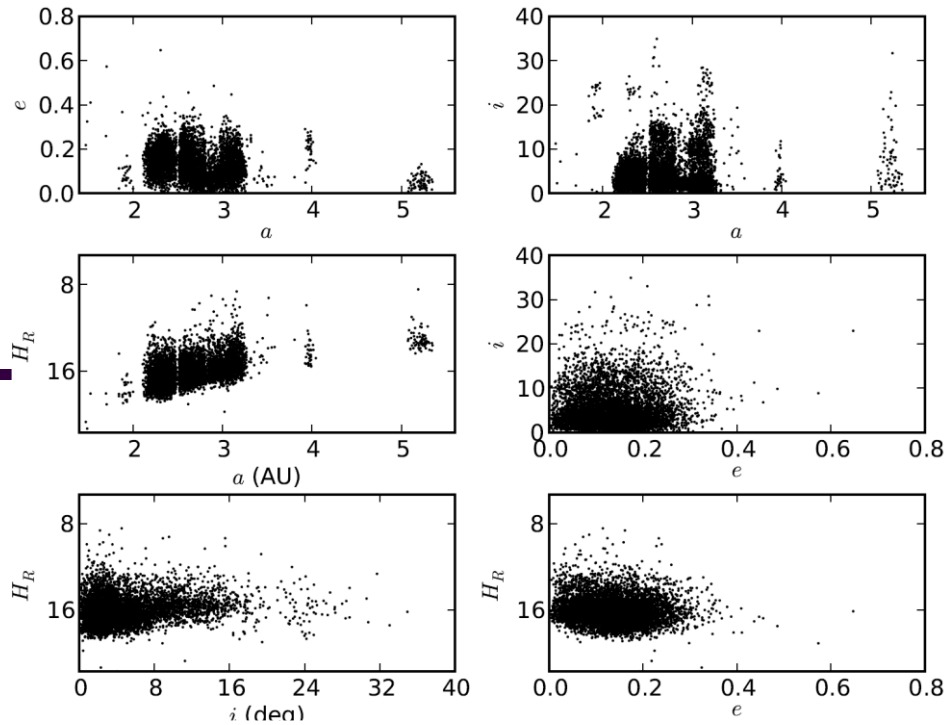


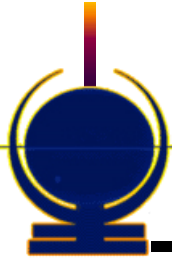


# Binary Asteroid

- **Binary asteroids** (e.g. Polishook, Brosch & Prialnik 2011)
  - Deep V-shaped minima and wide inverted-U-shaped maxima (Pravec et al., 2006)
  - Determination of the mass and/or density of the asteroids (Gnat & Sari, 2010)
  - Fractions in different environments to reveal formation model







# Q for Asteroid Rotation (1)

## ■ The Spin Rate Distribution

- $D > 40$  km
  - ➔ Maxwellian distribution (Pravec et al., 2002)
  - ➔ Collisionally evolved system (Salo, 1987)
- $D < 10$  km
  - ➔ Deviated from Maxwellian ?? (excesses of very slow/fast rotator; Vokrouhlicky, 2002)
  - ➔ YORP effect

