



High-Resolution Spectroscopy of Type Ia Supernovae

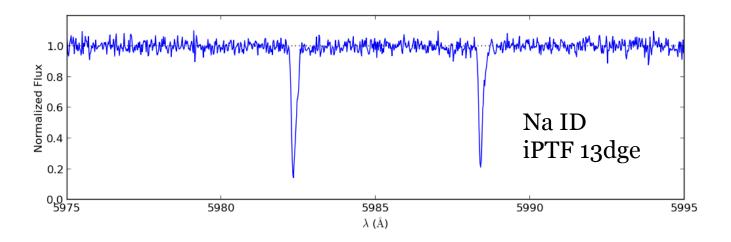
iPTF Workshop Stockholm, June 2014

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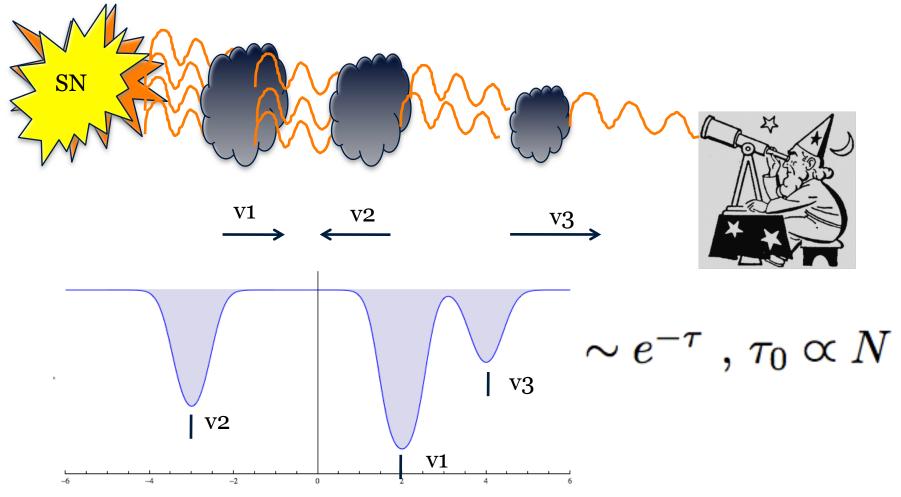
### What can we see?

- SN features have high velocity dispersions
  - High resolution spectroscopy unnecessary
- Intervening matter in line of sight
  - Narrow absorption features
  - Na I D, Ca II H&K, K I, DIBs





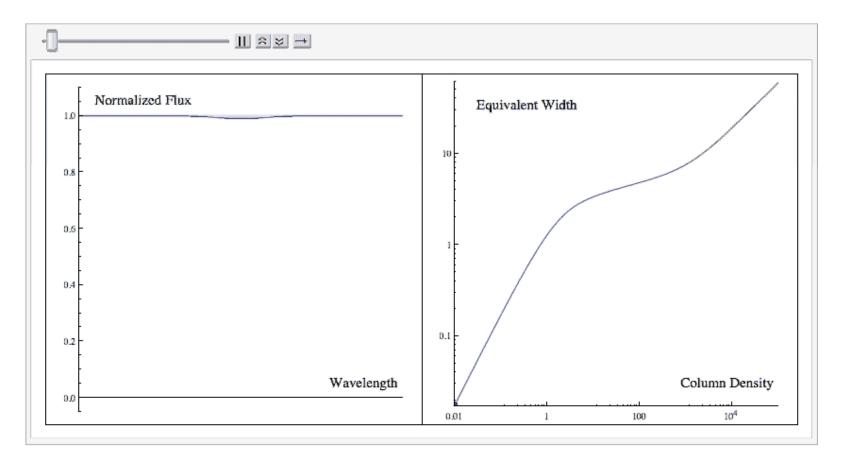
# Physical description of absorption spectra



The Oskar Klein Centre

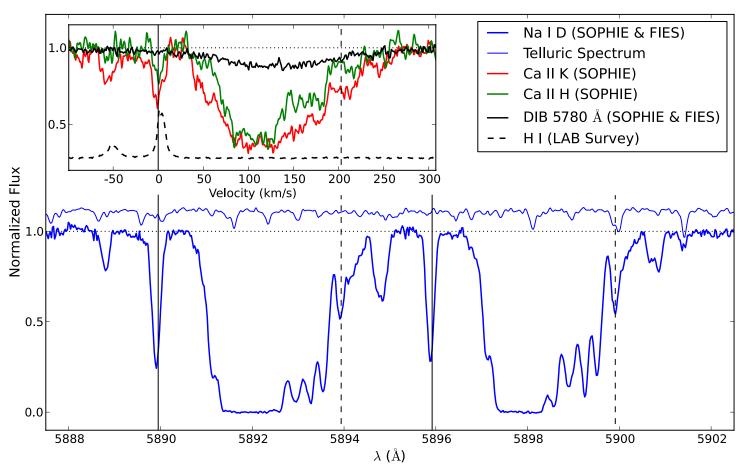


### Curve of Growth





## Example SN 2014J



Goobar et al. 2014



# What can we learn from the absorption lines?

#### Astrophysics

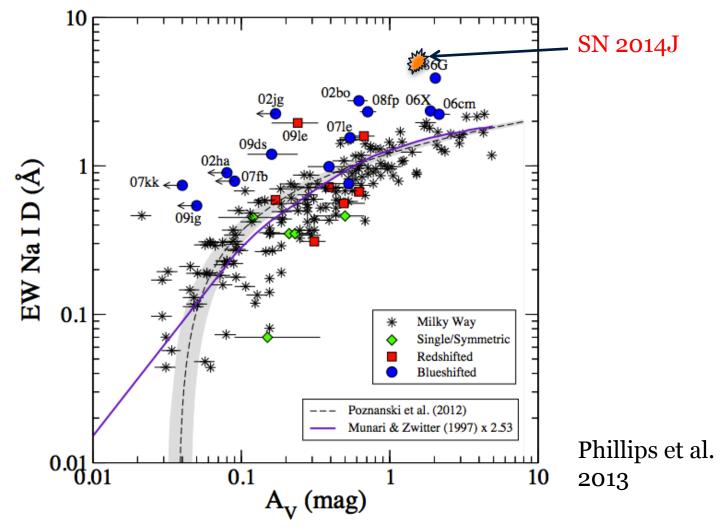
- Do the profiles and the equivalent width change?
  - Geometric effects?
  - Ionization → CSM!!
- Does the progenitor have a CSM?
  - SD CSM expected
  - DD possibly
- Statistically more blueshifted features (Sternberg et al. 2013)

#### Cosmology

- Absorption must correlate with reddening
- But ISM composition varies
- Equivalent width of Na I D doublet is a not an ideal proxy (Phillips et al. 2013)
- Must consider all features simultaneously

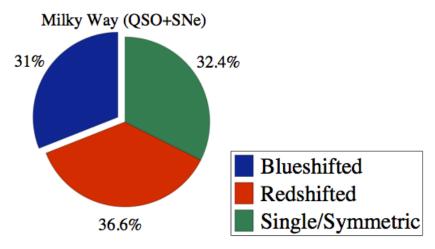


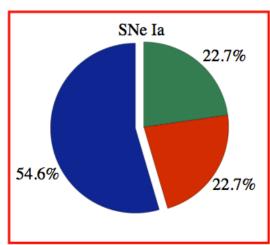
## Na ID as a reddening proxy?





## Is there Circum Stellar Matter?





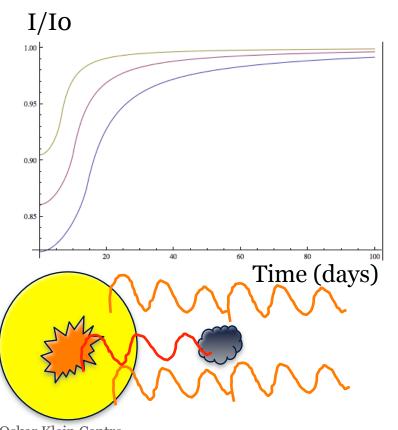
Sternberg et al. 2013

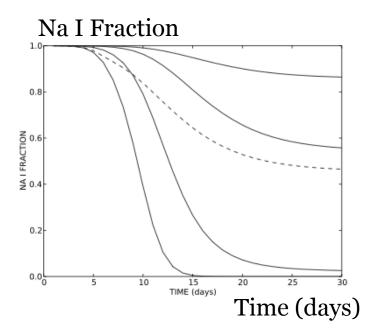


# Degeneracy of Geometric and Ionization effects

#### Geometric effects

#### Ionization of CSM

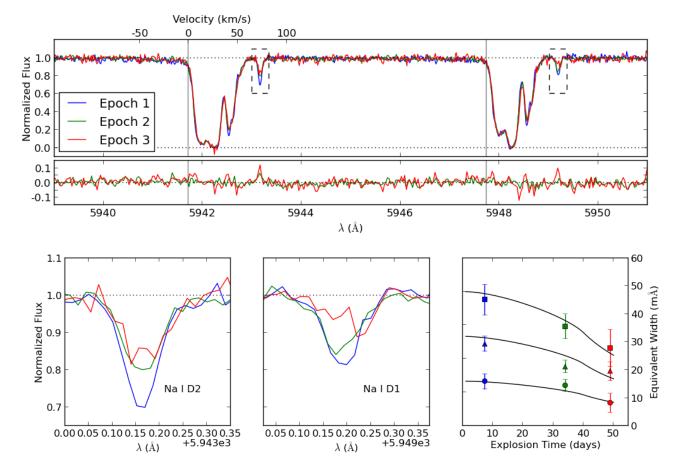




Borkowski et al. 2009



## SN Capricciosa (SN 2013gh)





### **Future Goals**

- Increase sample of multi-epoch (early!) high-resolution spectra
  - Are there changes apparent?
  - Can we verify/rule out the existence of a CSM?
  - What progenitor models does this favour?
- Identify useful reddening proxies
  - How do Na I D, Ca II H&K, K I, DIBs and others correlate with each other and to reddening
- VLT proposal with Assaf Sternberg



### References

- Goobar et al. (2014),
- Phillips et al. (2013), ApJ, 779, 38
- Sternberg et al. (2011), Science, 333, 856
- Patat et al.
- Borowski et al. (2009), ApJ, 699, L64-67
- Image of Astronomer: http://www.supercoloring.com/ pages/astronomer/