Available Keck NGAO Point Spread Functions

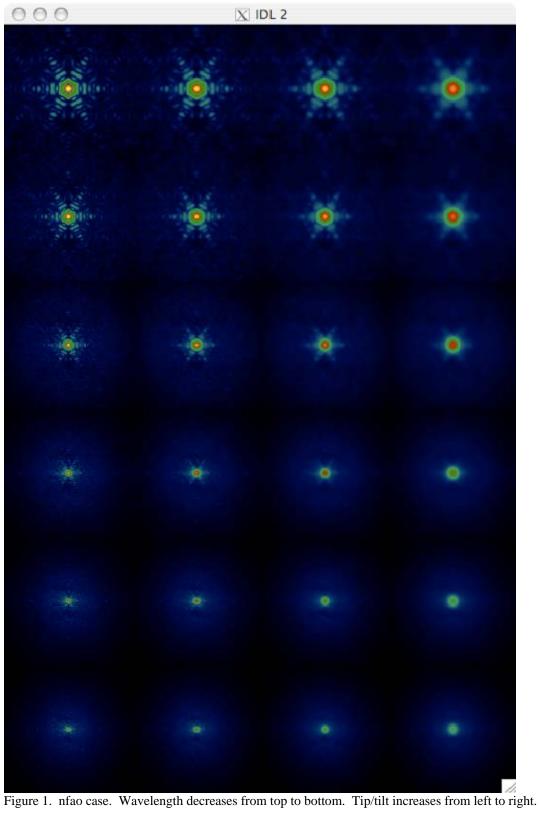
P. Wizinowich & C. Neyman April 15, 2006

This note is intended to provide a brief guide to the PSFs that can be found at <u>ftp://ftp.keck.hawaii.edu/outgoing/kpao/</u>. All of these are simulations, with the exception of the RYScuti folder. More details can be found in the README files in some of the folders.

Folder Name	Conjugate Type	RMS wavefront error (nm) on-axis	Wavelength Bands	Off-axis blurring (mas)
KPAO120	SCAO	120	R,I,J,H,K	0
nfao	SCAO	105	V,R,I,J,H,K	0
(see Fig. 1)				
nfaoSeeingVar	SCAO	90-130	V,R,I.J,H,K	0
nfao170	SCAO	~170	V,R,I,J,H,K	0
nfao140	SCAO	~140	V,R,I,J,H,K	0
singleLGS	SCAO (super K1	226	V,R,I,J,H,K	0
(see Fig. 2)	LGS)			
mcaoKPAO	MCAO	100	V,R,I,J,H,K	0
(see Fig. 3)				
RYScuti	SCAO (actual K2		Lp,Ms	
	NGS AO images)			

- The first seven folders contain simulations produced by Chris Neyman.
- A side-by-side comparison of some of the images can be found in the figures below (thanks to Ralf Flicker).
- The RYScuti folder contains images from the existing K2 AO system for comparison (thanks to David Le Mignant).
- Some K2 LGS AO images can be found at http://www2.keck.hawaii.edu/optics/lgsao/performance.html.
- In answer to requests about how the blurring relates to on-axis star magnitude and sky coverage Chris provided the following table with the following comments: "In the on-axis case, we can take advantage of the sharpening of the tip/tilt tracking star by the AO system. If we design a R band tracking with the correct plate scale, then the R magnitudes for the errors I generated in the PSF's would be (*those listed in the following table*). Remember these stars are on axis (almost never the case) and I'm taking full advantage of the sharpening of the guide star by AO."

Blurring (mas)	Rmag	% sky coverage at galactic NP
8	18.2	50
15	20.6	90
25	22.3	Complete



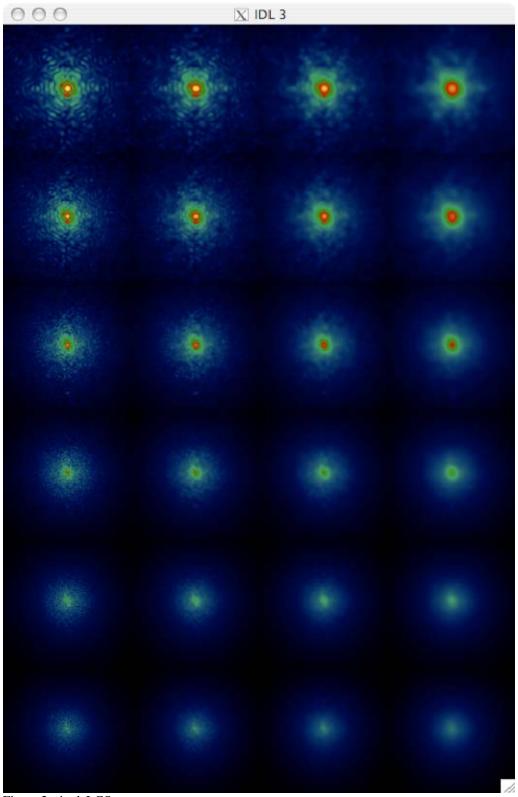


Figure 2. singleLGS case.

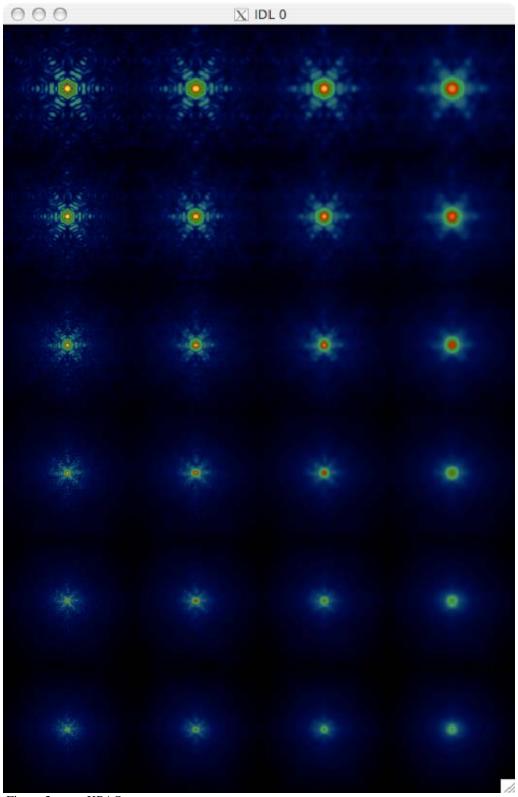


Figure 3. mcaoKPAO case.