KAON 548 - Science Requirements Summary in Support of System Architecture Evaluations

Minimum # of IFÚs IFU separation  AO Background Sky coverage High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)  Companion sensitivity  ΔI ≥	Imager  0.7-1.0 ≥ 15 na ≤ 10 (Nyquist at I band) na na  ≥ 30% for X3 ≤ 170 ≤ 15 na	Spectrograph  0.7-1.0  ≥ 2 (goal ≥ 3)  na  na  na  na  ≥ 30% for X3  ≤ 170  ≤ 15  follows from 170nm & 15mas	Imager  1.0-2.4 (+Y&z) ≥ 30" for S3 na ≤ 13 (Nyquist at J) na na Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170 ≤ 15 for sky cover; ≤ 3 for G2	Spectrograph  1.0-2.4 (+Y&z) ≥ 1 x 3 (goal 4) na na na na s > 30% of unattenuated(sky+tel) ≥ 30% for X3,X4a ≤ 170	Deployable IFU  1.0-2.4 (+Y&z)  ≥ 1 x 3  ≥ 120  ≤ 35 (2 pixels/spaxel)  6  > 1 IFU in 10x10" field  ≤ 30% of unattenuated(sky+tel)  ≥ 30% for X2  derived	Interferometer  J,H,K,L (N-band goal?  ≥ 1  ≥ 60  na  na  na  na
Field of view diameter (") Field of regard diameter (") Pixel size (mas) Minimum # of IFUs IFU separation  AO Background Sky coverage High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)	≥ 15 na s 10 (Nyquist at I band) na na   na  ≥ 30% for X3 ≤ 170 ≤ 15 na	≥ 2 (goal ≥ 3) na na na na a  na ≥ 30% for X3 ≤ 170 ≤ 15	≥ 30" for S3 na ≤ 13 (Nyquist at J) na na Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	≥ 1 x 3 (goal 4)  na  na  na  na  na  > 30% of unattenuated(sky+tel)  ≥ 30% for X3,X4a	≥ 1 x 3 ≥ 120 ≤ 35 (2 pixels/spaxel) 6 > 1 IFU in 10x10" field ≤ 30% of unattenuated(sky+tel) ≥ 30% for X2	≥ 1 ≥ 60 na na na
Field of regard diameter (") Pixel size (mas) Minimum # of IFUs IFU separation  AO Background Sky coverage High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)	na 10 (Nyquist at I band) na na  na  ≥ 30% for X3 ≤ 170 ≤ 15 na	na na na na ≥ 30% for X3 ≤ 170 ≤ 15	na ≤ 13 (Nyquist at J) na na Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	na na na na ≤ 30% of unattenuated(sky+tel) ≥ 30% for X3,X4a	≥ 120 ≤ 35 (2 pixels/spaxel) 6 > 1 IFU in 10x10" field ≤ 30% of unattenuated(sky+tel) ≥ 30% for X2	≥ 60 na na na
Field of regard diameter (") Pixel size (mas) Minimum # of IFUs IFU separation  AO Background Sky coverage High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)	na na ≥ 30% for X3 ≤ 170 ≤ 15 na	na na na na ≥ 30% for X3 ≤ 170 ≤ 15	≤ 13 (Nyquist at J) na na Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	na na na na ≤ 30% of unattenuated(sky+tel) ≥ 30% for X3,X4a	≤ 35 (2 pixels/spaxel) 6 > 1 IFU in 10x10" field  ≤ 30% of unattenuated(sky+tel) ≥ 30% for X2	na na na na
Pixel size (mas) Minimum # of IFUs IFU separation  AO Background Sky coverage High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)	na na ≥ 30% for X3 ≤ 170 ≤ 15 na	na na na ≥ 30% for X3 ≤ 170 ≤ 15	na na Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	na na na s ≤ 30% of unattenuated(sky+tel) ≥ 30% for X3,X4a	6 > 1 IFU in 10x10" field  ≤ 30% of unattenuated(sky+tel) ≥ 30% for X2	na na na
Minimum # of IFUs IFU separation  AO Background Sky coverage ligh order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)  Companion sensitivity  ΔI ≥	na na ≥ 30% for X3 ≤ 170 ≤ 15 na	na na ≥ 30% for X3 ≤ 170 ≤ 15	na na Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	na na  \$\frac{1}{2}\$ \$\leq 30\% of unattenuated(sky+tel) \$\geq 30\% for X3,X4a	6 > 1 IFU in 10x10" field  ≤ 30% of unattenuated(sky+tel) ≥ 30% for X2	na na na
AO Background Sky coverage High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)  Companion sensitivity  ΔI ≥	na ≥ 30% for X3 ≤ 170 ≤ 15 na	na ≥ 30% for X3 ≤ 170 ≤ 15	Background (mag/arcsec^2) J: 15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	na ≤ 30% of unattenuated(sky+tel) ≥ 30% for X3,X4a	≤ 30% of unattenuated(sky+tel) ≥ 30% for X2	na na
Sky coverage ligh order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas)  Companion sensitivity  ΔI ≥	≥ 30% for X3 ≤ 170 ≤ 15 na	≥ 30% for X3 ≤ 170 ≤ 15	15.9, H: 13.7, K wide field 13.6, K narrow field 13.2 ≥ 30% for X1,X3,X4b ≤ 170	≥ 30% for X3,X4a	≥ 30% for X2	
Sky coverage  High order WFE (nm) for ≤ 5" fov  Tip/tilt error (mas)  50% Ensquared energy (mas)  Companion sensitivity  ΔI ≥	≥ 30% for X3 ≤ 170 ≤ 15 na	≥ 30% for X3 ≤ 170 ≤ 15	≥ 30% for X1,X3,X4b ≤ 170	≥ 30% for X3,X4a	≥ 30% for X2	
High order WFE (nm) for ≤ 5" fov Tip/tilt error (mas) 50% Ensquared energy (mas) Companion sensitivity ΔI ≥	≤ 170 ≤ 15 na	≤ 170 ≤ 15	≤ 170			na
Tip/tilt error (mas) 50% Ensquared energy (mas)  Companion sensitivity ΔI ≥	≤ 15 na	≤ 15	_ ·· ·	≤ 170	dorived	
50% Ensquared energy (mas)  Companion sensitivity ΔI ≥	na		≤ 15 for sky cover; ≤ 3 for G2		delived	≤ 250
Companion sensitivity ΔI ≥		follows from 170nm & 15mas		≤ 15	derived	≤ 15
	1> 7 5 -4 0 75 1 for 04		na	≤ 25 at J band (X3)	≤ 70	na
			$\Delta$ J ≥ 5.5 at 0.5" for S1b; $\Delta$ J ≥ 8.5			
Photometry (mag) ≤ 0	I ≥ 7.5 at 0.75" for S1b	na	at 0.1", ∆J ≥ 11 at 0.2" for G1	na	na	na
, ( 0,	≤ 0.05 relative for S1b	na	≤ 0.05 relative for S1&G1 ≤ 1.5-2 for S1b&G1 ≤ 0.1 for	na	na	na
	≤ 1.5 relative for S1b	na	G2a	na	na	na
Polarimetry (%)		na		na	na	na
PSF estimation	required	goal	required	goal	PSF spectrum reqd	not required
	required up to 50"/hr	required up to 50"/hr	required up to 50"/hr	required up to 50"/hr	goal: 1 tip/tilt sensor	not required
acquisition accuracy (mas or % of		≤ 10% for IFU; ≤ 0.25λ/D				
instrument field)	≤ 10% of field	for slit	≤ 10%	$\leq$ 10% for IFU; $\leq$ 0.25 $\lambda$ /D for slit	≤ 10% (≤ 35 relative)	≤ 200 mas
Dither dist (" or % of inst field)	≤50%	≤ 50% of longest dimension	≤ 50%	≤ 50%	≤ 50% of longest dimension	na
Dither accuracy (mas)	≤ λ/D	≤λ/D	≤ λ/D	≤ \(\lambda/\)D	≤ 70	na
Dither time (sec)	≤ 3	≤ 3	≤ 3	≤ 3	≤ 10	na
Micro dither distance (mas)	≤ 0.5λ/D	≤ 0.5λ/D	≤ 0.5λ/D	≤ 0.5λ/D	≤ 35	
Micro dither accuracy (mas)	≤ 0.25λ/D	≤ 0.25λ/D	≤ 0.25\/D	≤ 0.25λ/D	<10	
Micro dither time (sec)	≤ 3	≤ 3	≤ 3	≤ 3	≤ 3	
Nod reacquisition time (sec)	≤ 10	≤ 10	≤ 10	≤ 10	≤ 60	
Positioning knowledge (mas)	≤ 0.1λ/D	≤ 0.1λ/D	≤ 0.1λ/D	≤ 0.1 λ/D	≤ 5	na . <del>-</del>
Science image drift (mas/hr)	≤ 5	≤ 5	≤ 5	≤ 5	≤ 5	≤ 5
NGS mode	required	required	required	required	single IFU	required
AO instrument switching	to vis spectro	to vis imager	to NIR spectro (goal: vis)	to NIR imager (goal: vis)	not required	not required
AO backup switching goa	goal: to NIR instrument	goal: to NIR instrument	not required	not required	single IFU	goal: to NIR instrume
Science Cases S	S1b,S2,S3,X3,X4b	X3	S1,S3,S4,G1,G2a,G3,G4,X1,X3 ,X4b,X4d,X5,X6	S3,S4,G1,G2b,G3,G4,X1,X3,X4a	G2b,X2,X4c	

Ocience Cases	310,02,03,X3,X40 X3	,X4b,X4d,X5,X6	55,54,61,62b,65,64,71,75,74a	020,X2,X40	
	Science Cases	Color code		Seeing Assumptions	
S1a	Asteroid companions survey		Physical	requirements	All values at $\lambda = 0.5  \mu \text{m}$
S1b	Asteroid companions orbit determination		Performano	ce requirements	Challenging NGAO
S2	Asteroid size and shape		Operational requirements		37.5th percentile
S3	Gas Giants and Moons of giant planets		Scien	nce cases	$r_0 = 14 \text{ cm}; \theta_0 = 2.15$ "
S4	NGS observations of Neptune & Uranus			Other	Median NGAO
G1	Planets around low mass stars		Goal: AO transmits	$H\alpha$ to visible instruments	50th percentile
G2a	General Relativity & the Galactic Center - astrometry		NGS WFS field of	of regard ≥ 30" radius	$r_0 = 16 \text{ cm}; \theta_0 = 2.7"$
G2b	General Relativity & the Galactic Center - radial velocit	ties	Interferometer req'ments must be	met by a NGAO mode (matched field,	Good NGAO
G3	Debris Disks		polarization	n, etc. for K1&2)	62.5th percentile
G4	Young Stellar Objects		non-AO backup	required for all cases	$r_0 = 18 \text{ cm}; \theta_0 = 2.9$ "
X1	QSO host galaxies				Excellent NGAO
X2	High-z galaxies			" NIR) to 2k Nyquist sampled detector	87.5th percentile
X3	Nearby AGNs		Fixed field & fixed	I pupil modes required	$r_0 = 22 \text{ cm}; \theta_0 = 4.0"$
X4a	Distant galaxies lensed by galaxies - spectroscopy		·	ition capabilities required	
X4b	Distant galaxies lensed by galaxies - imaging		Alignment, calibratio	n & diagnostic tools req'd	j
X4c	Distant galaxies lensed by clusters - spectroscopy				
X4d	Distant galaxies lensed by galaxies - imaging				
X5	Astrometry in Sparse Fields				
X6	Resolved Stellar Populations in Crowded Fields				