

# **IRIS OIWFS Concept Study**

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#### **IRIS on TMT**





#### **Delivered Focal Plane**





# Key OIWFS Requirements

- Instrument rotator & services wrap alignment between NFIRAOS delivered beam and IRIS science instrument
  - Critical to IRIS image quality
- 2 tip/tilt and 1 tip/tilt/focus wavefront sensors
  - Fast guiding, focus and tilt anisoplanatism
- 2 mas positioning accuracy > 4.4 µm at focal plane
  - OIWFS are the positional reference for astrometric performance
- High sky coverage < 2 mas tip/tilt jitter at galactic pole</p>
  - Near infrared, 1.0 to 1.7  $\mu m$ , sensing on 'sharpened' guide stars
- High acquisition probability = low acquisition time



# Number & Type of Probes

(Lianqi Wang, Brent Ellerbroek, 2008)



5



- Object Select Mechanisms considered
  - Robot placed pickoff & pathlength mirrors (EAGLE concept)
  - Tip/tilt mirror tiling of focal plane (Caltech TIPI concept)
  - Theta-Phi, 2 rotation stages (Flamingos-2 OIWFS, etc)
  - Theta-R, 1 rotation + 1 linear stages (IRMOS, KMOS, etc)
- Requirements for positioning accuracy, dithering, and non-sidereal tracking
- Theta-R concept chosen for initial sky coverage analysis
- TT and TTF function change versus 2 probe planes



#### Patrol Geometry (Lianqi Wang, Brent Ellerbroek, 2008)

Cumulative Histogram of T/T Jitter, WindShake 20 mas, HgCdTe with 8e RoN



3 identical 'theta-R' probes at 120 degree spacing

Each probe capable of TT or TTF

Only need to reach 50% across 2' field











#### **Theta-R Probe Optics**







Results shown for 0, 0.3, 0.7, and 1.0 arcmin off axis

264000.00



#### **RMS Focus Error**

**RMS Focus Error: No Collimator Focus** 



**RMS Focus Error: Collimator Focussed** 



#### **Color Dependent Strehl**

Strehl: No Collimator Focus







# OIWFS ADC Current Version

•The doublet prism pair.

•Glass selection inspired by Drew Philips' glass-pair investigation.

•Diameter = 30 mm.

- •Thickness of a doublet = 14 mm.
- •Glasses are S-NPH2 and Spinel.

•It is quite well-behaved being in a collimated beam space.





ADC Residual Color Error vs. Wavelength S-NPH2 + SPINEL





#### **Macro Mechanics**





Exploded View





#### **OIWFS Probe Platform**

THIRTY METER TELESCOPE









## **Custom Linear Stage**



20



## Collimator focus stage





## **Back end optics**





## **OIWFS Cameras**

- TT pixel scale =  $\lambda_{\rm H}/2D$  = 5.67 mas/pixel
- TTF pixel scale =  $\lambda_H/D$  = 11.34 mas/pixel
  - Same detectors can be used for TT and TTF functions
- Analysis of tip/tilt jitter & image wander during acquisition shows worst case of ~250 mas
  - at  $4\sigma$  and 5.67 mas/pixel > minimum of ~180x180 pixels
  - Acquisition probability > larger probe FoV > more pixels
- Hawaii HxRG 1024x1024 HgCdTe detector w/new MBE material testing underway – read noise, dynamics
- HgCdTe e-APDs and InGaAs emerging technologies



## **Acquisition Probability**

(Corinne Boyer, Luc Simard, 2008)





## Noise vs. Frame Rate

measured for various frame sizes





25











## **Opto-mechanical follow**







# **On-chip follow**





![](_page_29_Picture_0.jpeg)

# Sky Coverage Update

(Lianqi Wang, Brent Ellerbroek, Jean-Pierre Veran, 2009)

- NGS mode errors combined with LGS mode errors of 178 nm RMS
- Overall on-axis budget of 187 nm RMS met at <u>45%</u> sky coverage
- Shortfall (in quadrature) of ~28 nm RMS at 50%
- System optimization still underway
  - Detector performance
  - "Fitting field" for LGS modes
  - Optimal choice of NGS modes and reconstructor

![](_page_29_Figure_10.jpeg)