

Palomar Adaptive Optics Test Plan

Title	Low-order servo loop optimization
Version	1.0
Date released	4/3/2007
Lead	A. Bouchez
Time requested	2 hours
Required conditions	Photometric, >90deg. from moon.

Purpose

1. Determine optimal low-order servo loop parameters as a function of NGS magnitude.
2. Practice acquisition procedure with faint NGS.

Previous analysis

Starting estimates (please note optimized values):

NGS V magnitude	Acq integ. (s)	Framerate (Hz)	Integral gain
10	0.03	1000	0.5
12	0.5	500	0.5
14	1.0	200	0.5
15	5.0	150	0.5
16	10.0	100	0.5
17	20.0	50	0.5
18	30.0	30	0.5

Test procedure

Targets for this test: isolated V=14, 15, 16, 17 stars near zenith. Use red stars (B-V = 1.0-2.5) to maximize PHARO sensitivity.

1. Follow *LGS Target Acquisition* testplan for star.
 - 1.1. Only perform NGS tune-up on the first target.
 - 1.2. Use LOWFS framerate from the table above.
 - 1.3. DO perform "zero LOWFS centroids" step!
 - 1.4. Record acquisition camera image before offsetting LOWFS.
2. Determine PHARO integration time for decent SNR images. Goal is 10-15k peak counts, but keep integration time <30s.
3. Take a PHARO sky
 - 3.1. Open TT loop from TAO command line (TAO button? / PHARO button??)
 - 3.2. Offset telescope from PHARO buttons (or TAO)
 - 3.3. Take a PHARO sky.
 - 3.4. Offset back to target.
 - 3.5. Close TT loop (same way as opened).
4. Record average Strehl of 3 frames in table below.
5. Adjust framerate
 - 5.1. Open TT loop.
 - 5.2. Set framerate to one "level" higher
 - 5.3. Take new LOWFS background (offset 60" to sky).
 - 5.4. Close TT loop.

