KAON 548 - Science Requirements Summary in Support of System Architecture Evaluations

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	Visible		Near-IR		Near-IR	
Requirement	Imager	Spectrograph	Imager	Spectrograph	Deployable IFU	Interferometer
λ (μm)	0.7-1.0	0.7-1.0	1.0-2.4 (+Y&z)	1.0-2.4 (+Y&z)	1.0-2.4 (+Y&z)	J,H,K,L (N-band goal?)
Field of view diameter (")	≥ 15	≥ 2 (goal ≥ 3)	≥ 15 for X4b	≥ 1 x 3 (goal 4)	≥1 x 3	≥ 1
Field of regard diameter (")	na	na	na	na	≥ 120	≥ 60
Pixel size (mas)	≤ 10 (Nyquist at I band)	na	≤ 13 (Nyquist at J)	na	≤ 35 (2 pixels/spaxel)	na
Minimum # of IFUs	na	na	na	na	6	na
IFU separation	na	na	na	na	> 1 IFU in 10x10" field	na
AO Background	na	na	≤ 20% of (sky + tel)	≤ 20% of (sky + tel)	≤ 20% of (sky + tel)	na
Sky coverage	≥ 30% for X3	≥ 30% for X3	≥ 30% for X1,X3,X4b	≥ 30% for X3,X4a	≥ 30% for X2	na
High order WFE (nm) for ≤ 5" fov	≤ 170	≤ 170	≤ 170	≤ 170	derived	≤ 250
Tip/tilt error (mas)	≤ 15	≤ 15	≤ 15 for sky cover; ≤ 3 for G2	≤ 15	derived	≤ 15
50% Ensquared energy (mas)	na	follows from 170nm & 15mas	na	≤ 25 at J band (X3)	≤ 70	na
			$\Delta J \ge 5.5$ at 0.5" for S1b; $\Delta J \ge 8.5$			
Companion sensitivity	ΔI ≥ 7.5 at 0.75" for S1b	na	at 0.1", ∆J ≥ 11 at 0.2" for G1	na	na	na
Photometry (mag)	≤ 0.05 relative for S1b	na	≤ 0.05 relative for S1&G1	na	na	na
i notemouy (mag)	_ 0.00 .0.00.0.0 .0. 0.0		≤ 1.5-2 for S1b&G1 ≤ 0.1 for			
Astrometry (mas)	≤ 1.5 relative for S1b	na	G2a	na	na	na
Polarimetry (%)		na		na	na	na
PSF estimation	required	goal	required	goal	PSF spectrum regd	not required
Differential tracking	required up to 50"/hr	required up to 50"/hr	required up to 50"/hr	required up to 50"/hr	goal: 1 tip/tilt sensor	not required
Acquisition accuracy (mas or % of		≤ 10% for IFU; ≤ 0.25λ/D		≤ 10% for IFU; ≤ 0.25λ/D		·
instrument field)	≤ 10% of field	for slit	≤ 10%	for slit	≤ 10% (≤ 35 relative)	≤ 200 mas
Dither dist (" or % of inst field)	≤50%	≤ 50% of longest dimension	≤ 50%	≤ 50%	≤ 50% of longest dimension	na
Dither accuracy (mas)	≤ λ/D	≤λ/D	≤ \(\lambda/\)D	≤ λ/D	≤ 70	na
Dither time (sec)	≤ 3	≤ 3	≤ 3	≤ 3	≤ 10	na
Micro dither distance (mas)	≤ 0.5λ/D	≤ 0.5λ/D	≤ 0.5λ/D	≤ 0.5λ/D	≤ 35	
Micro dither accuracy (mas)	≤ 0.25λ/D	≤ 0.25λ/D	≤ 0.25λ/D	≤ 0.25λ/D	<10	
Micro dither time (sec)	≤ 3	≤ 3	≤ 3	≤ 3	≤ 3	
Nod reacquisition time (sec)	≤ 10	≤ 10	≤ 10	≤ 10	≤ 60	
Positioning knowledge (mas)	≤ 0.1λ/D	≤ 0.1λ/D	≤ 0.1λ/D	≤ 0.1 λ/D	≤ 5	na
Science image drift (mas/hr)	≤ 5	≤ 5	≤ 5	≤ 5	≤ 5	≤ 5
NGS mode	required	required	required	required	single IFU	required
AO instrument switching	to vis spectro	to vis imager	to NIR spectro (goal: vis)	to NIR imager (goal: vis)	not required	not required
AO backup switching	goal: to NIR instrument	goal: to NIR instrument	not required	not required	single IFU	goal: to NIR instrument
Saionea Casas	C1h C2 C2 V2 V4h		S1,S3,S4,G1,G2a,G3,G4,X1,X3,	S3,S4,G1,G2b,G3,G4,X1,	- J	
Science Cases	S1b,S2,S3,X3,X4b	X3	X4b,X4d,X5,X6	X3,X4a	G2b,X2,X4c	
Science Cases Color code See						Socing Assumptions

Science Cases				
S1a	Asteroid companions survey			
S1b	Asteroid companions orbit determination			
S2	Asteroid size and shape			
S3	Gas Giants and Moons of giant planets			
S4	NGS observations of Neptune & Uranus			
G1	Planets around low mass stars			
G2a	General Relativity & the Galactic Center - astrometry			
G2b	General Relativity & the Galactic Center - radial velocities			
G3	Debris Disks			
G4	Young Stellar Objects			
X1	QSO host galaxies			
X2	High-z galaxies			
X3	Nearby AGNs			
X4a	Distant galaxies lensed by galaxies - spectroscopy			
X4b	Distant galaxies lensed by galaxies - imaging			
X4c	Distant galaxies lensed by clusters - spectroscopy			
X4d	Distant galaxies lensed by galaxies - imaging			
X5	Astrometry in Sparse Fields			
X6	Resolved Stellar Populations in Crowded Fields			
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Color code Physical requirements Performance requirements Operational requirements Science cases Other A O transmits Ho to visible instruments

Goal: AO transmits Hα to visible instruments
NGS WFS field of regard ≥ 30" radius
Interferometer req'ments must be met by a NGAO
mode (matched field, polarization, etc. for K1&2)
non-AO backup required for all cases
Goal: Provide full field (20" vis, 40" NIR) to 2k Nyquist
sampled detector

Fixed field & fixed pupil modes required Tip/tilt & LGS acquisition capabilities required Alignment, calibration & diagnostic tools req'd

Seeing Assumptions All values at $\lambda = 0.5 \ \mu m$ Challenging NGAO 37.5th percentile $r_0 = 14 \ cm; \ \theta_0 = 2.15''$ Median NGAO 50th percentile $r_0 = 16 \ cm; \ \theta_0 = 2.7''$ Good NGAO 62.5th percentile $r_0 = 18 \ cm; \ \theta_0 = 2.9''$ Excellent NGAO 87.5th percentile $r_0 = 22 \ cm; \ \theta_0 = 4.0''$