

KAON 548 - Science Requirements Summary in Support of System Architecture Evaluations

Requirement	Visible		Near-IR		Near-IR	Interferometer
	Imager	Spectrograph	Imager	Spectrograph	Deployable IFU	
$\lambda$ ( $\mu\text{m}$ )	0.7-1.0	0.7-1.0	1.0-2.4 (+Y&z)	1.0-2.4 (+Y&z)	1.0-2.4 (+Y&z)	J,H,K,L (N-band goal?)
Field of view diameter (")	$\geq 15$	$\geq 2$ (goal $\geq 3$ )	$\geq 15$ for X4b	$\geq 1 \times 3$ (goal 4)	$\geq 1 \times 3$	$\geq 1$
Field of regard diameter (")	na	na	na	na	$\geq 120$	$\geq 60$
Pixel size (mas)	$\leq 10$ (Nyquist at I band)	na	$\leq 13$ (Nyquist at J)	na	$\leq 35$ (2 pixels/spaxel)	na
Minimum # of IFUs	na	na	na	na	6	na
IFU separation	na	na	na	na	$> 1$ IFU in $10 \times 10''$ field	na
AO Background	na	na	$\leq 20\%$ of (sky + tel)	$\leq 20\%$ of (sky + tel)	$\leq 20\%$ of (sky + tel)	na
Sky coverage	$\geq 30\%$ for X3	$\geq 30\%$ for X3	$\geq 30\%$ for X1,X3,X4b	$\geq 30\%$ for X3,X4a	$\geq 30\%$ for X2	na
High order WFE (nm) for $\leq 5''$ fov	$\leq 170$	$\leq 170$	$\leq 170$	$\leq 170$	derived	$\leq 250$
Tip/tilt error (mas)	$\leq 15$	$\leq 15$	$\leq 15$ for sky cover; $\leq 3$ for G2	$\leq 15$	derived	$\leq 15$
50% Ensquared energy (mas)	na	follows from 170nm & 15mas	na	$\leq 25$ at J band (X3)	$\leq 70$	na
Companion sensitivity	$\Delta J \geq 7.5$ at 0.75" for S1b	na	$\Delta J \geq 5.5$ at 0.5" for S1b; $\Delta J \geq 8.5$ at 0.1", $\Delta J \geq 11$ at 0.2" for G1	na	na	na
Photometry (mag)	$\leq 0.05$ relative for S1b	na	$\leq 0.05$ relative for S1&G1	na	na	na
Astrometry (mas)	$\leq 1.5$ relative for S1b	na	$\leq 1.5-2$ for S1b&G1; $\leq 0.1$ for G2a	na	na	na
Polarimetry (%)	na	na	na	na	na	na
PSF estimation	required	goal	required	goal	PSF spectrum reqd	not required
Differential tracking	required up to 50"/hr	required up to 50"/hr	required up to 50"/hr	required up to 50"/hr	goal: 1 tip/tilt sensor	not required
Acquisition accuracy (mas or % of instrument field)	$\leq 10\%$ of field	$\leq 10\%$ for IFU; $\leq 0.25\lambda/D$ for slit	$\leq 10\%$	$\leq 10\%$ for IFU; $\leq 0.25\lambda/D$	$\leq 10\%$ ( $\leq 35$ relative)	$\leq 200$ mas
Dither dist ( " or % of inst field)	$\leq 50\%$	$\leq 50\%$ of longest dimension	$\leq 50\%$	$\leq 50\%$	$\leq 50\%$ of longest dimension	na
Dither accuracy (mas)	$\leq \lambda/D$	$\leq \lambda/D$	$\leq \lambda/D$	$\leq \lambda/D$	$\leq 70$	na
Dither time (sec)	$\leq 3$	$\leq 3$	$\leq 3$	$\leq 3$	$\leq 10$	na
Micro dither distance (mas)	$\leq 0.5\lambda/D$	$\leq 0.5\lambda/D$	$\leq 0.5\lambda/D$	$\leq 0.5\lambda/D$	$\leq 35$	na
Micro dither accuracy (mas)	$\leq 0.25\lambda/D$	$\leq 0.25\lambda/D$	$\leq 0.25\lambda/D$	$\leq 0.25\lambda/D$	$< 10$	na
Micro dither time (sec)	$\leq 3$	$\leq 3$	$\leq 3$	$\leq 3$	$\leq 3$	na
Nod reacquisition time (sec)	$\leq 10$	$\leq 10$	$\leq 10$	$\leq 10$	$\leq 60$	na
Positioning knowledge (mas)	$\leq 0.1\lambda/D$	$\leq 0.1\lambda/D$	$\leq 0.1\lambda/D$	$\leq 0.1\lambda/D$	$\leq 5$	na
Science image drift (mas/hr)	$\leq 5$	$\leq 5$	$\leq 5$	$\leq 5$	$\leq 5$	$\leq 5$
NGS mode	required	required	required	required	single IFU	required
AO instrument switching	to vis spectro	to vis imager	to NIR spectro (goal: vis)	to NIR imager (goal: vis)	not required	not required
AO backup switching	goal: to NIR instrument	goal: to NIR instrument	not required	not required	single IFU	goal: to NIR instrument
Science Cases	S1b,S2,S3,X3,X4b	X3	S1,S3,S4,G1,G2a,G3,G4,X1,X3,X4b,X4d,X5,X6	S3,S4,G1,G2b,G3,G4,X1,X3,X4a	G2b,X2,X4c	

Science Cases	
S1a	Asteroid companions survey
S1b	Asteroid companions orbit determination
S2	Asteroid size and shape
S3	Gas Giants and Moons of giant planets
S4	NGS observations of Neptune & Uranus
G1	Planets around low mass stars
G2a	General Relativity & the Galactic Center - astrometry
G2b	General Relativity & the Galactic Center - radial velocities
G3	Debris Disks
G4	Young Stellar Objects
X1	QSO host galaxies
X2	High-z galaxies
X3	Nearby AGNs
X4a	Distant galaxies lensed by galaxies - spectroscopy
X4b	Distant galaxies lensed by galaxies - imaging
X4c	Distant galaxies lensed by clusters - spectroscopy
X4d	Distant galaxies lensed by galaxies - imaging
X5	Astrometry in Sparse Fields
X6	Resolved Stellar Populations in Crowded Fields

Color code
Physical requirements
Performance requirements
Operational requirements
Science cases
<b>Other</b>
Goal: AO transmits H $\alpha$ to visible instruments
NGS WFS field of regard $\geq 30''$ radius
Interferometer req'ments must be met by a NGAO mode (matched field, polarization, etc. for K1&2)
non-AO backup required for all cases
Goal: Provide full field (20" vis, 40" NIR) to 2k Nyquist sampled detector
Fixed field & fixed pupil modes required
Tip/tilt & LGS acquisition capabilities required
Alignment, calibration & diagnostic tools req'd

Seeing Assumptions
All values at $\lambda = 0.5 \mu\text{m}$
Challenging NGAO
37.5th percentile
$r_0 = 14 \text{ cm}; \theta_0 = 2.15''$
Median NGAO
62.5th percentile
$r_0 = 18 \text{ cm}; \theta_0 = 2.9''$
Good NGAO
87.5th percentile
$r_0 = 22 \text{ cm}; \theta_0 = 4.0''$